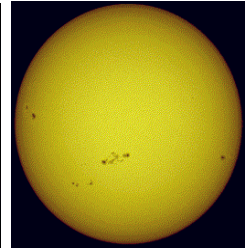
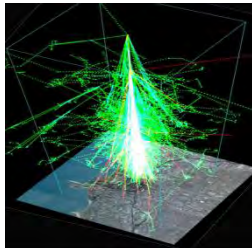
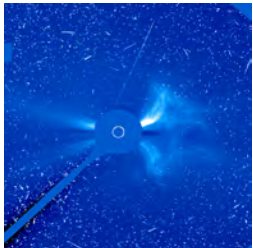


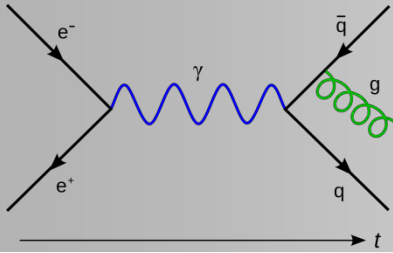
Long-term solar change and solar influences on global and regional climates

Mike Lockwood

*(University of Reading,
& Space Science and Technology Department,
STFC/Rutherford Appleton Laboratory)*

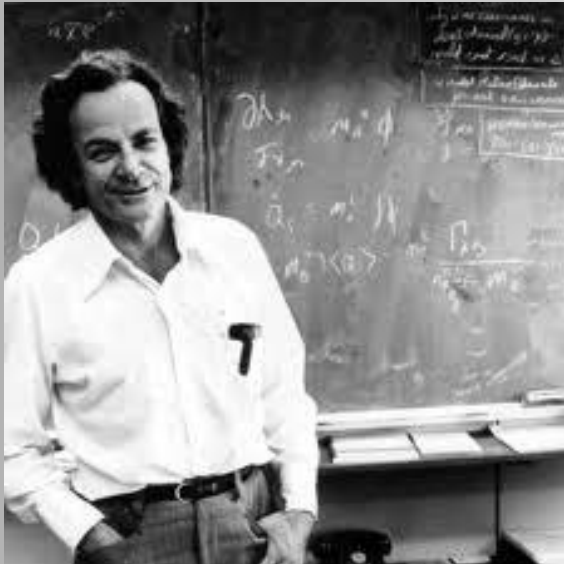
STFC Introductory Solar System Plasma Physics Summer School
Newcastle, 13th September 2017





“The first principle is that you must not fool yourself and you are the easiest person to fool”

“reality must take precedence over public relations, for Nature cannot be fooled”



Richard P. Feynman
(1918-1988)



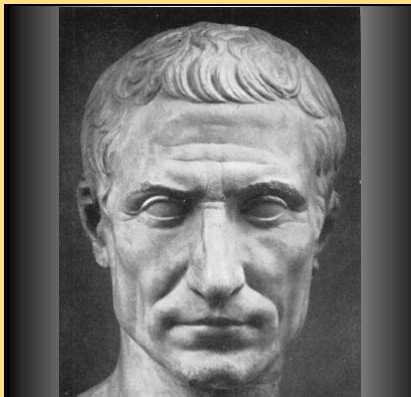
“Still, a man hears what he wants to hear and disregards the rest”

(Paul Simon, The Boxer, 1970)



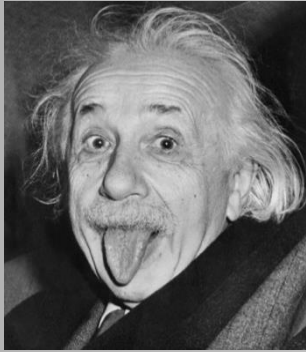
“men may construe things after their fashion, clean from the purpose of the things themselves”

(William Shakespeare, Julius Ceasar, 1599)



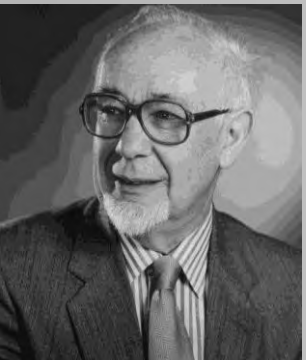
“men, in general are quick to believe that which they wish to be true.”

(Julius Ceasar, 50BC)



Science

🔊 *SAIƏNS*
(noun)



- Cambridge Dictionary: “(knowledge from) the careful study of the structure & behaviour of the physical world, especially by watching, measuring, and doing experiments, and the development of theories to describe the results of these activities”
- Wikipedia: “(from Latin *scientia*, meaning knowledge) is a systematic enterprise that builds and organizes knowledge in the form of testable explanations and predictions about the universe.”
- OED: “A systematically organized body of knowledge on a particular subject.”
- John Michael Ziman (1925-2005): “....‘consensibility’, leading to **consensus**, is the touchstone of reliable knowledge”



Science Consensus

🔊 *saɪəns kən 'sɛnsəs*
(compound noun)



- Wikipedia: “the collective judgment, position, and opinion of the community of scientists in a particular field of study. Consensus implies general agreement, though not necessarily unanimity”

Climate change: there IS an overwhelming scientific consensus



OPEN ACCESS

IOP PUBLISHING

Environ. Res. Lett. 8 (2013) 024024 (7pp)

ENVIRONMENTAL RESEARCH LETTERS

doi:10.1088/1748-9326/8/2/024024

Quantifying the consensus on anthropogenic global warming in the scientific literature

John Cook^{1,2,3}, Dana Nuccitelli^{2,4}, Sarah A Green⁵, Mark Richardson⁶, Bärbel Winkler², Rob Painting², Robert Way⁷, Peter Jacobs⁸ and Andrew Skuce^{2,9}

¹ Global Change Institute, University of Queensland, Australia

² Skeptical Science, Brisbane, Queensland, Australia

³ School of Psychology, University of Western Australia, Australia

⁴ Tetra Tech, Incorporated, McClellan, CA, USA

⁵ Department of Chemistry, Michigan Technological University, USA

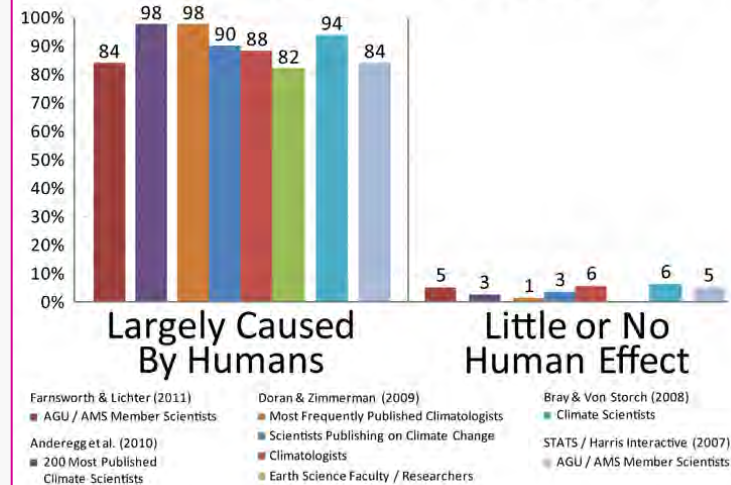
⁶ Department of Meteorology, University of Reading, UK

⁷ Department of Geography, Memorial University of Newfoundland, Canada

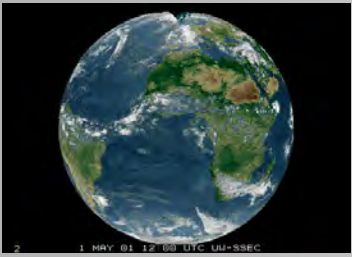
⁸ Department of Environmental Science and Policy, George Mason University, USA

⁹ Salt Spring Consulting Ltd, Salt Spring Island, BC, Canada

Opinions of Climate and Earth Scientists on Global Warming



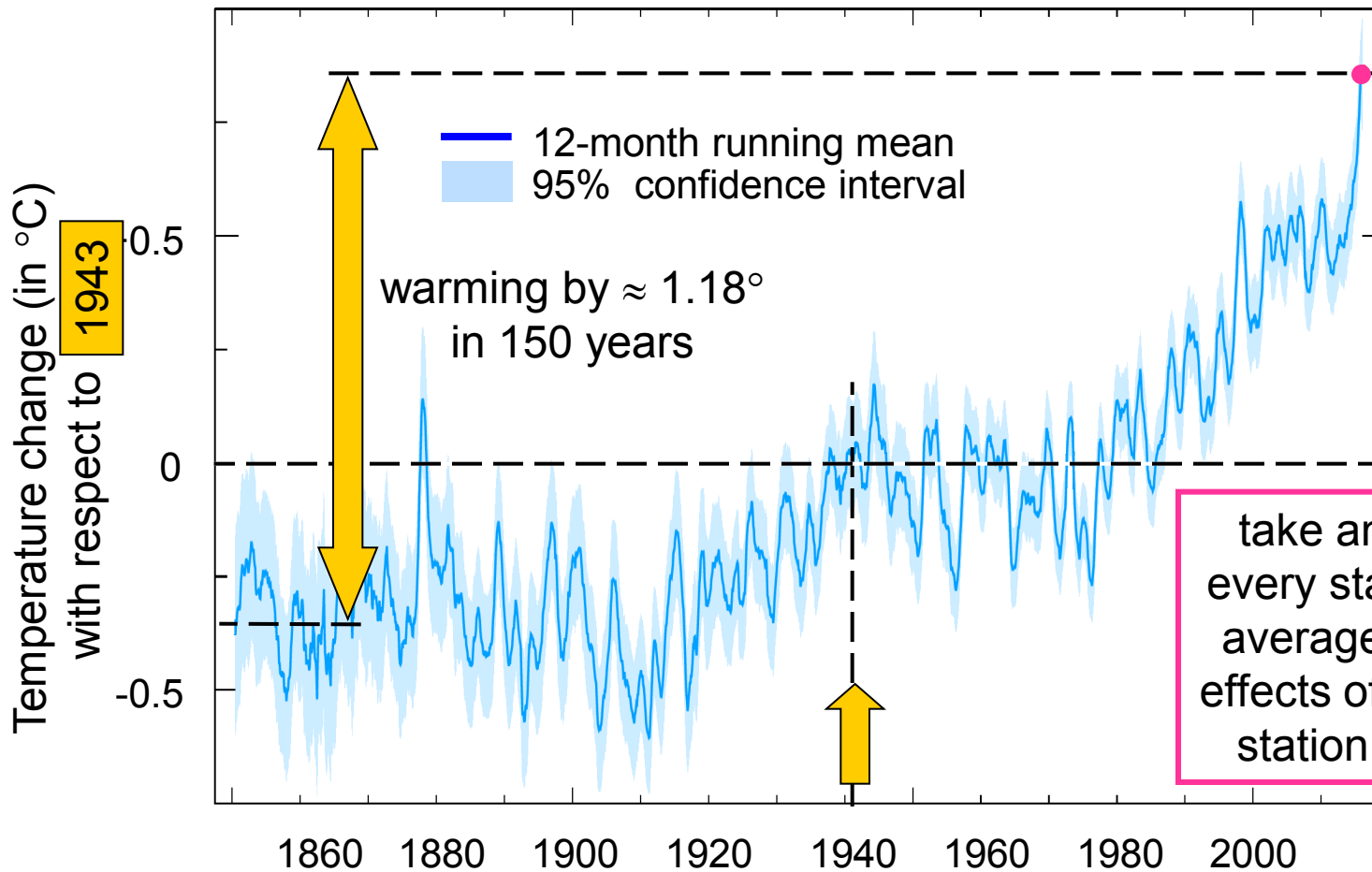
↑ Survey of all papers published 1991-2011 using keywords “climate change” and “global warming” (11944 of them)
97% of papers offering an opinion on climate change agreed that human activities are causing global warming



Is the Earth Warming?



Average surface temperature anomaly measured by the global network of weather stations (data from CRU, UEA)

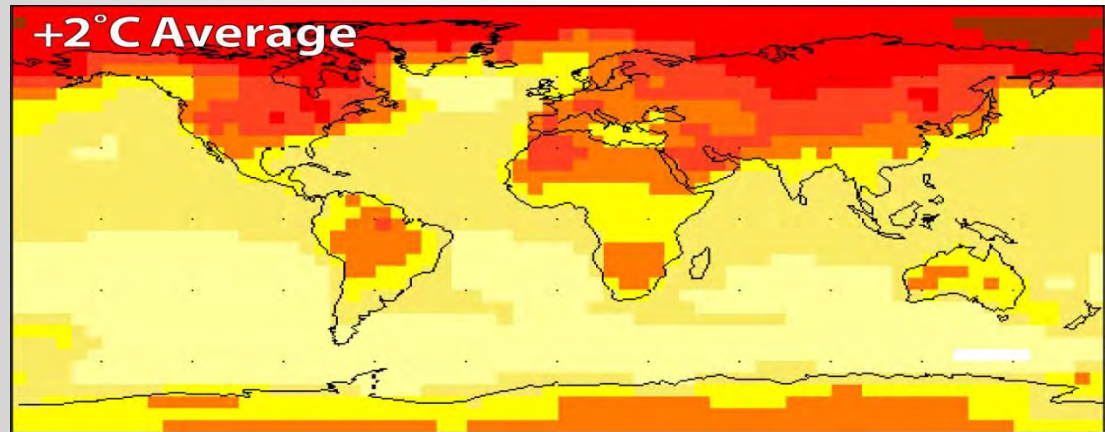


take anomaly for every station & then average (limits the effects of changes in station locations)

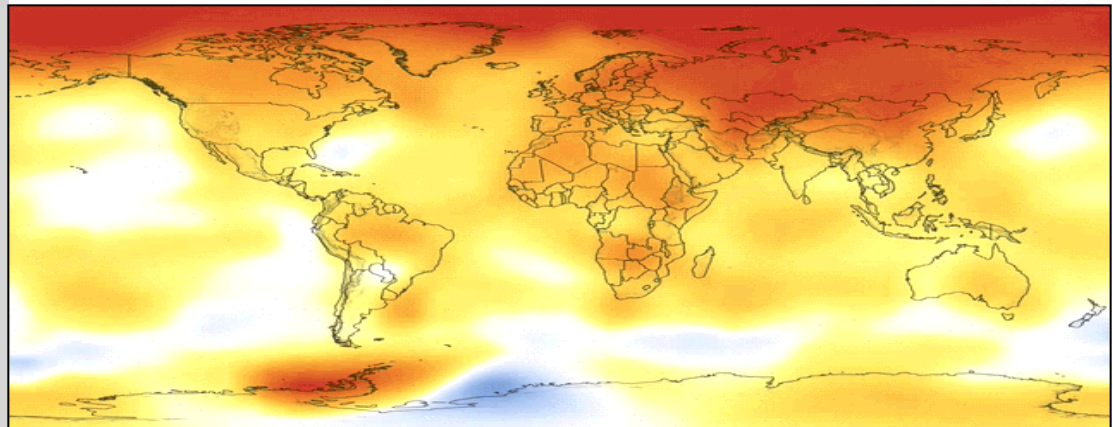


Map of Air Surface Temperature rise predicted in 1988

MODELLED AST MAP
– for a GMAST rise of
 $\Delta T_s = +2^\circ\text{C}$

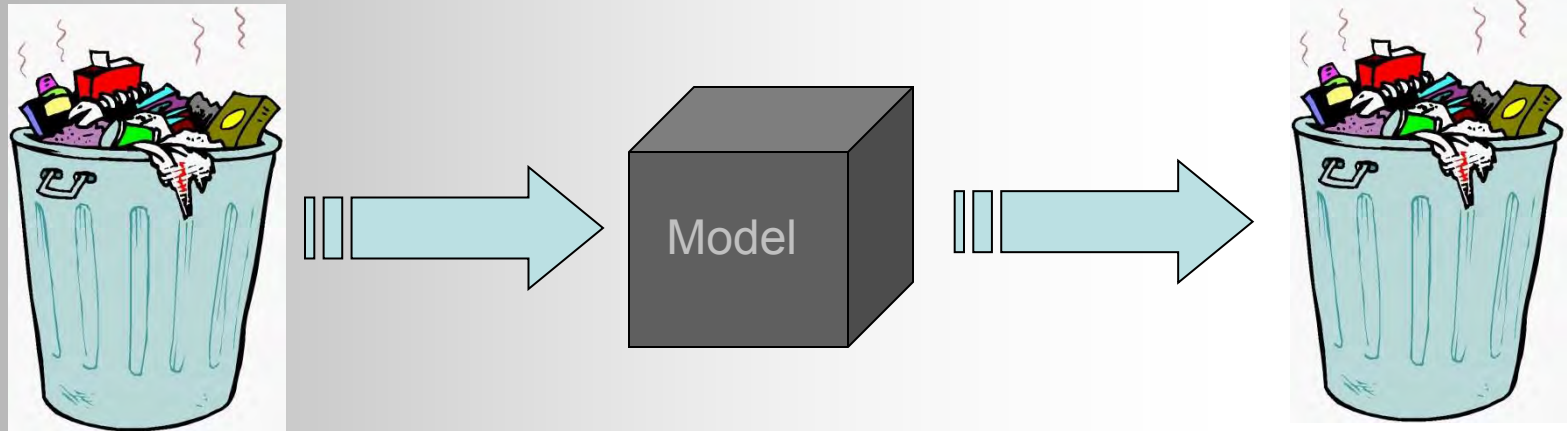


OBSERVED AST MAP
– NASA/GISS data for
1881-2008 (for which
measured GMAST rise
 $\approx 1.1^\circ\text{C}$)





A sceptical view of models



This is **always** true

- hard to evaluate without detailed knowledge of model and its application
- when different models say the same thing, we need to take them seriously
- and note that we can be irrationally selective about which models we chose to believe and disbelieve! (such selection is often needed – we must ensure we sue rational and objective selection)

The Greenhouse Effect



▶ First suggested by Svante Arrhenius (1896)



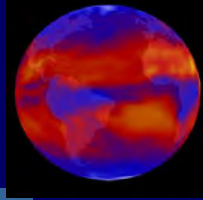
▶ CO₂ rise first linked to temperature rise by Guy Stewart Callendar (1939)



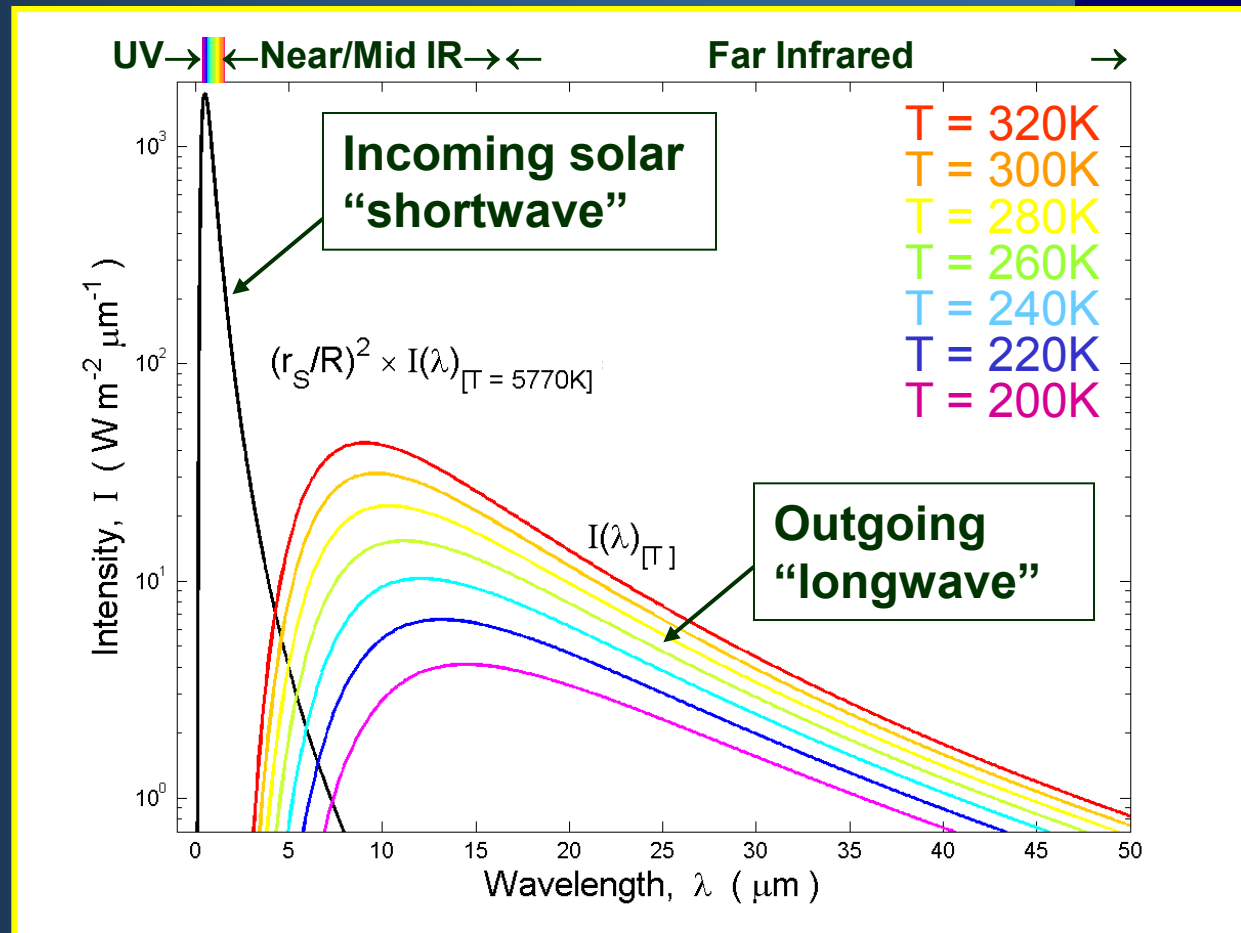
▶ Concern is that perturbations will cause runaway greenhouse effect suffered by Venus

▶ Venus was initially very similar to Earth but: (1) was closer to the Sun; (2) could not remove CO₂ by tectonic subduction and (3) never developed a biomass to keep CO₂ in its atmosphere in check

Spectra at the Heart of the Greenhouse Effect



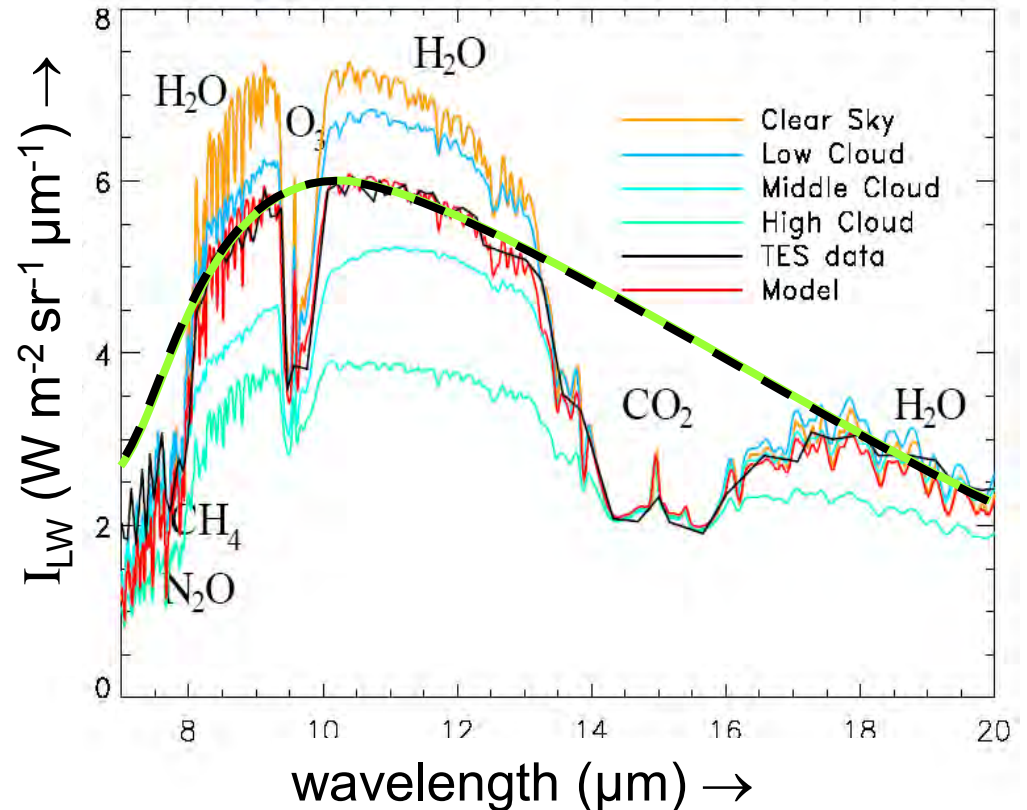
- A “blackbody” is an ideal radiator, that is often seen in nature
- The sun is close to a blackbody of temperature $T = 5770 \text{ K}$
- Different parts of Earth radiate with different T
- To show SW and LW on same plot we here use a logarithmic intensity scale





The greenhouse effect

Spectrum of outgoing longwave (infra red)

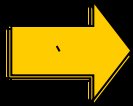


observations from
Mars Global Surveyor
(in black)

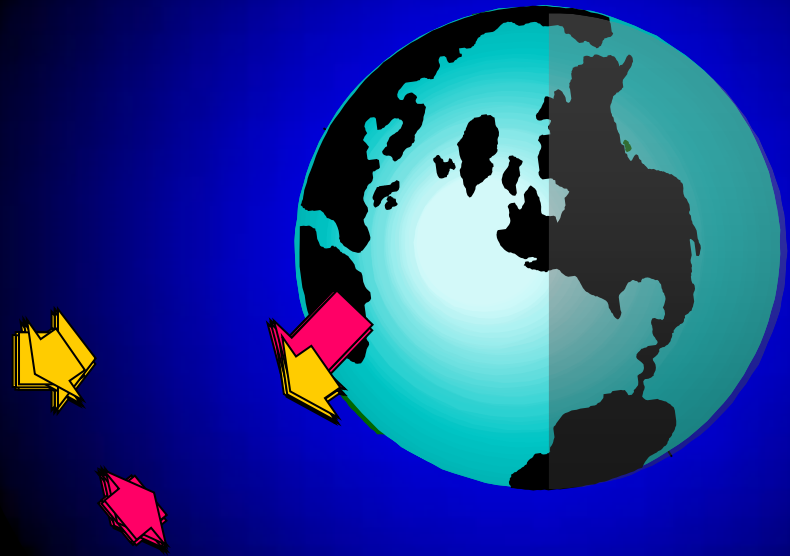
Model is the appropriate
mix of Earth "scene"
types (in red)

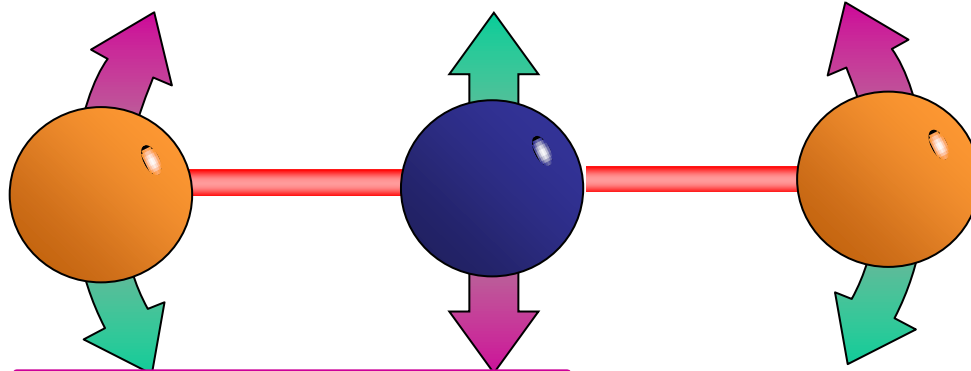
The Greenhouse Effect

- ▶ incoming solar power (called shortwave or SW)



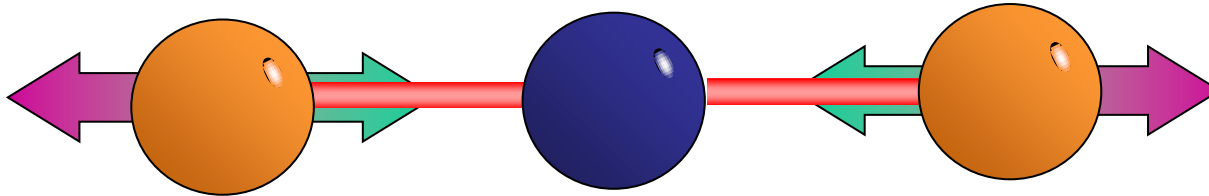
- ▶ about 1/3 reflected back into space (“albedo”)
- ▶ the rest heats Earth’s surface
- ▶ which re radiates thermal longwave (LW) radiation
- ▶ but the atmosphere traps in some of that re-radiated LW radiation – heats surface a bit more
- ▶ increasing the LW trapping causes T_{SE} to rise so that P_e rises enough to keep $P_{in} \approx P_{out}$






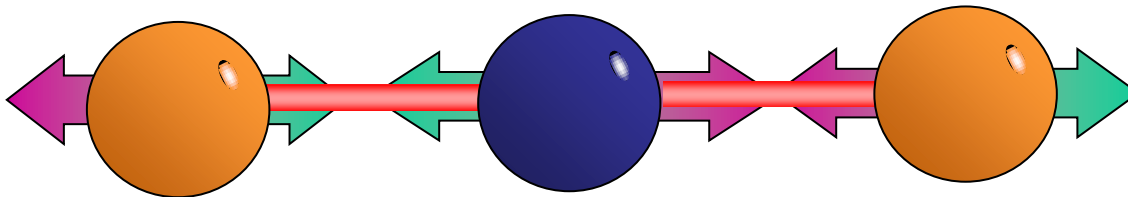
(a) Bending mode


Carbon



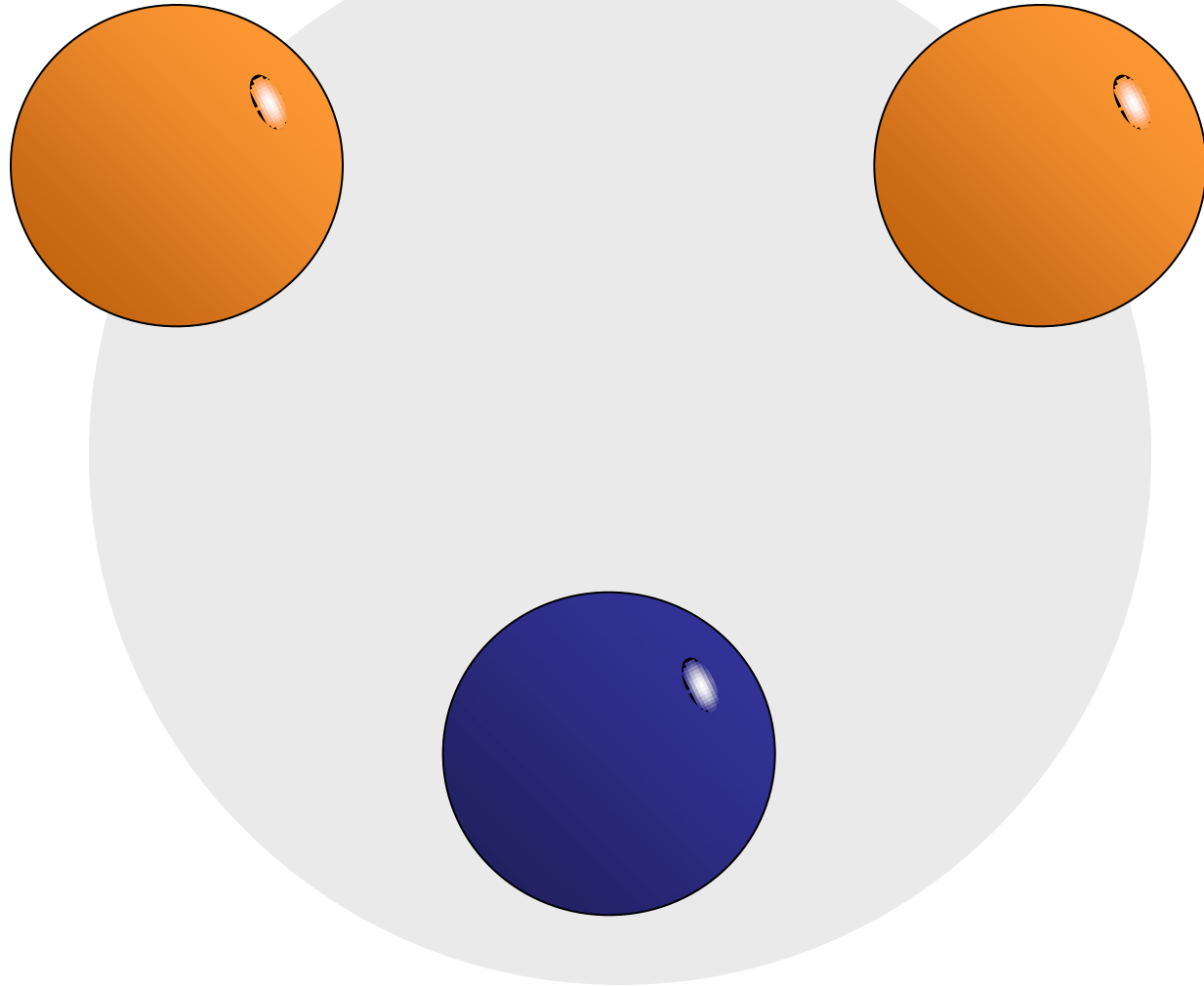
(b) symmetric stretch


Oxygen




(c) asymmetric stretch

a CO₂ molecule

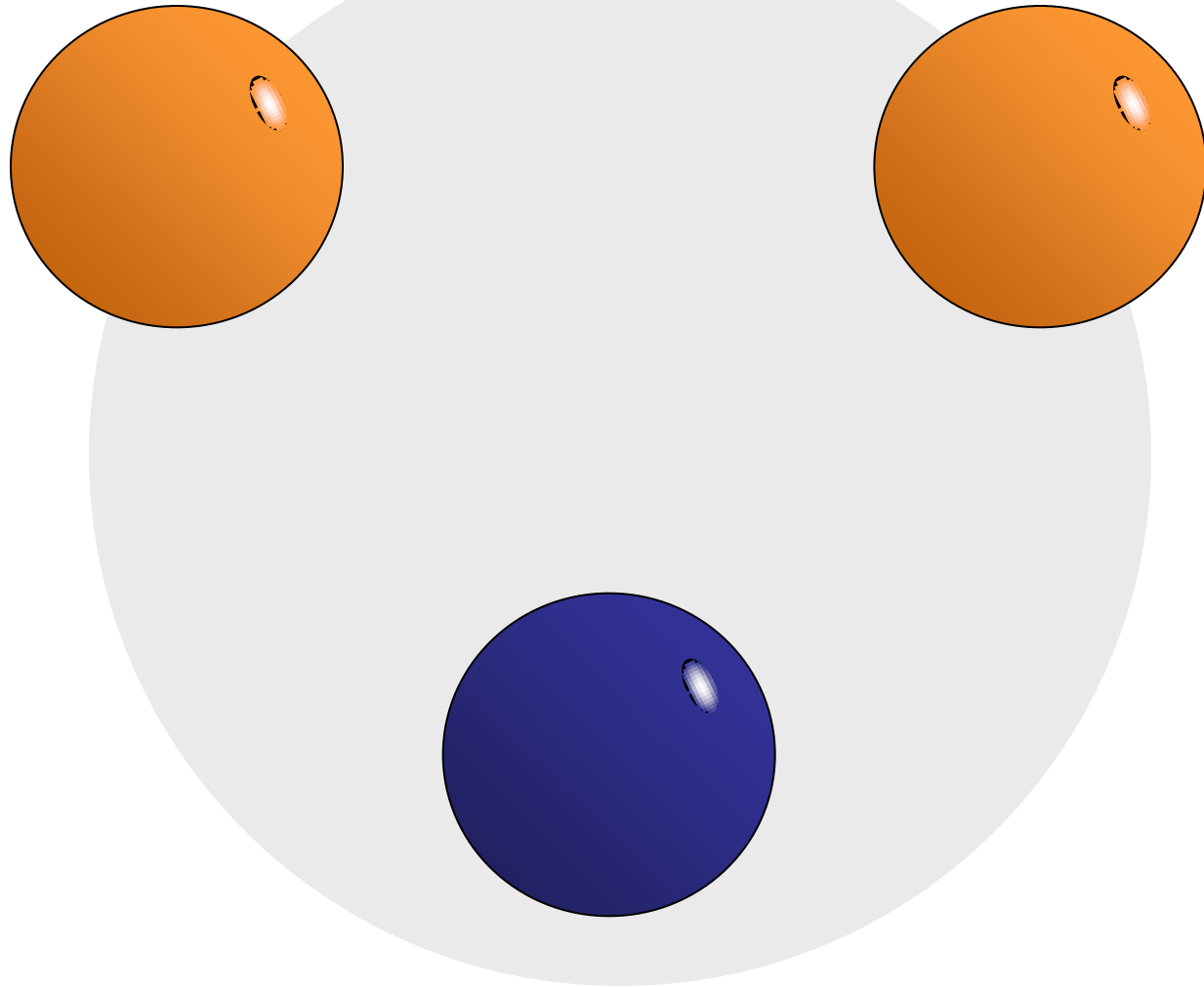


 Carbon

 Oxygen


SW Photon 

a CO₂ molecule

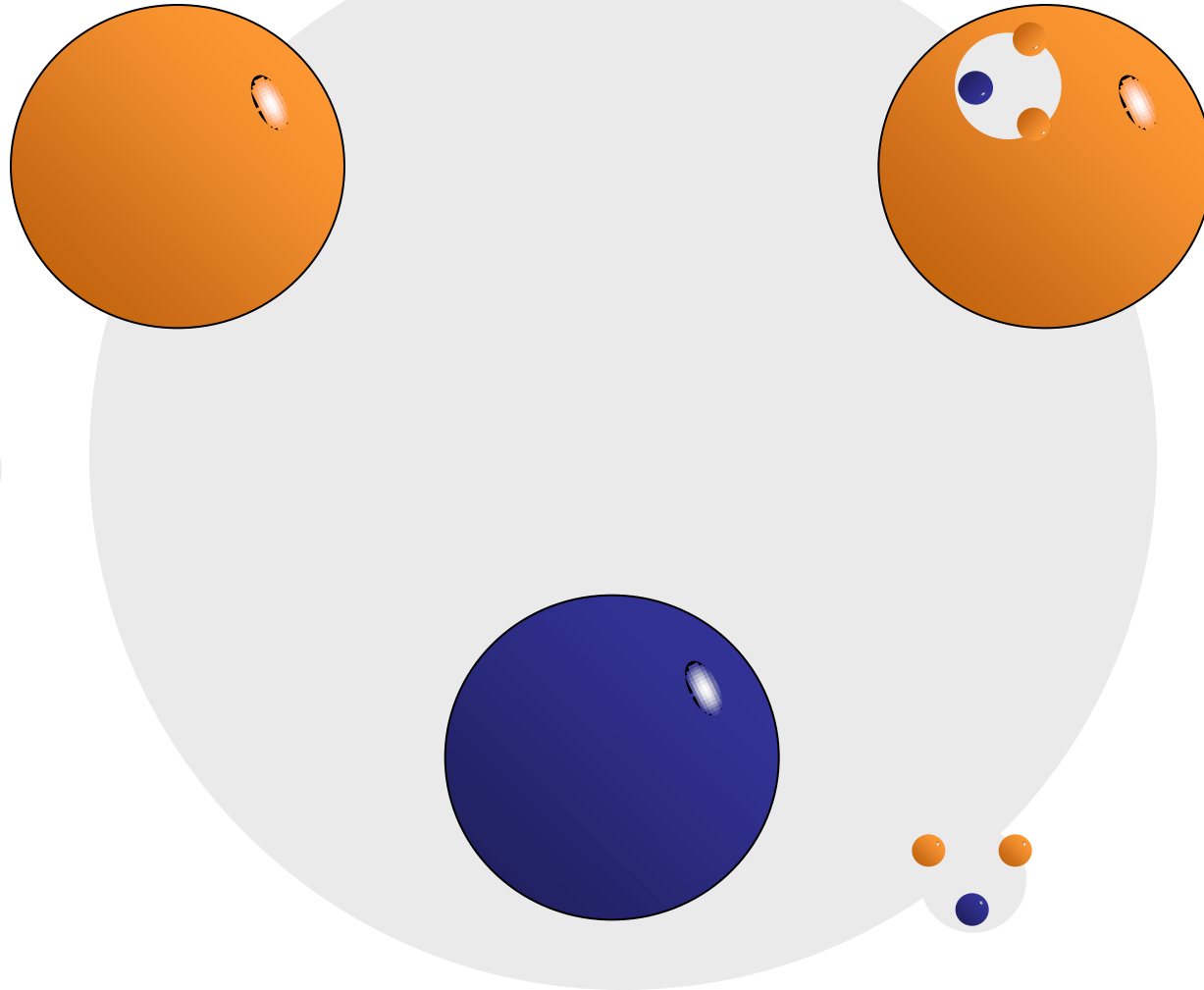


 Carbon

 Oxygen


LW Photon 

a CaCO_3 capsule

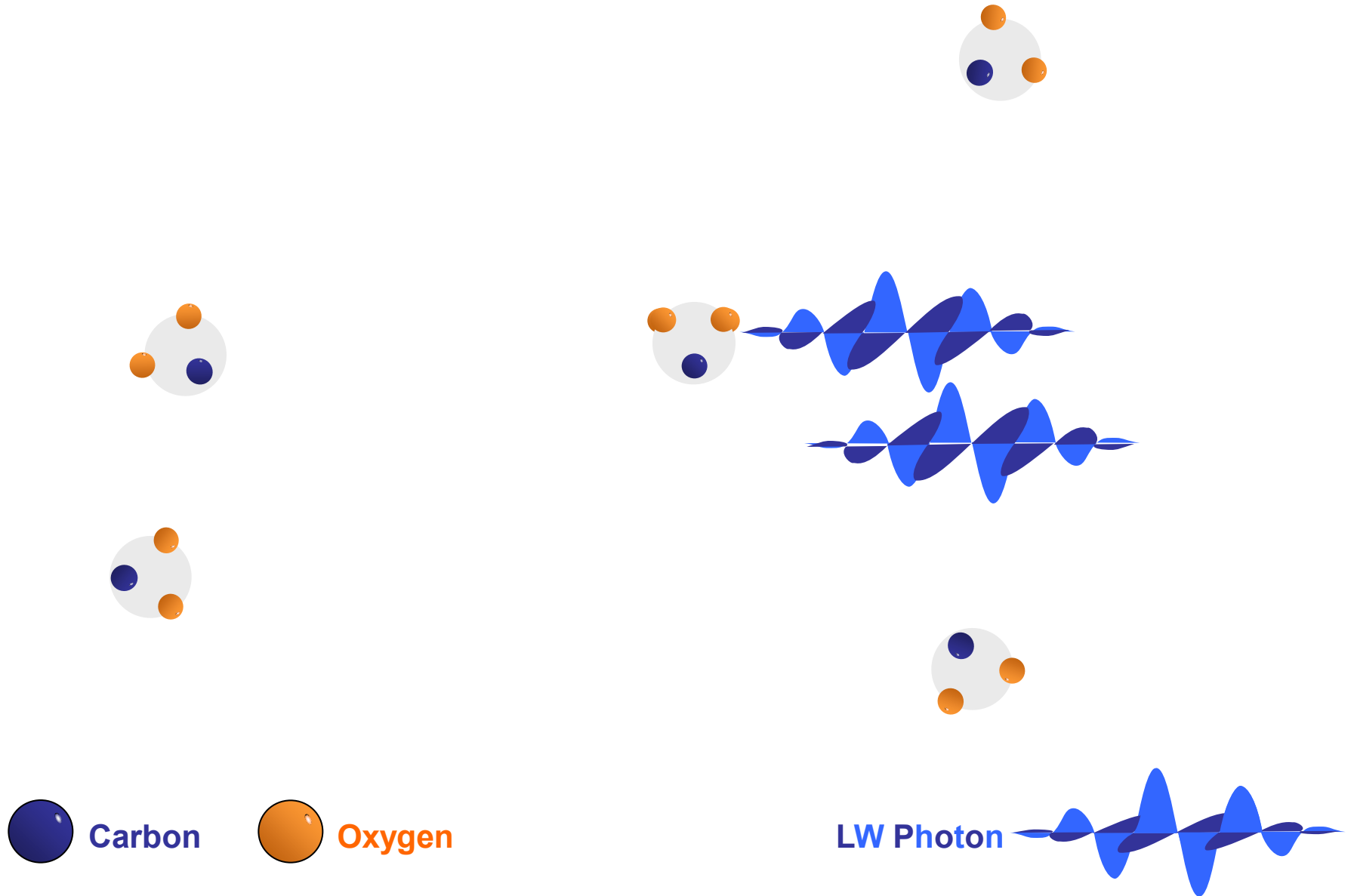


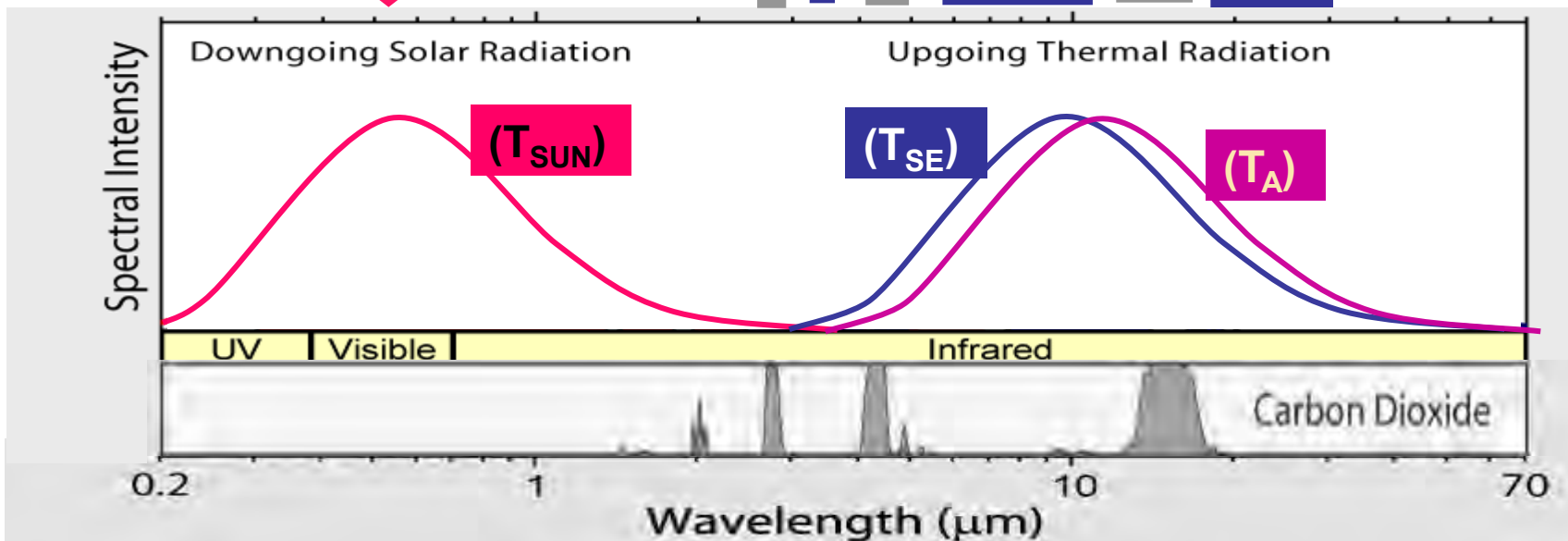
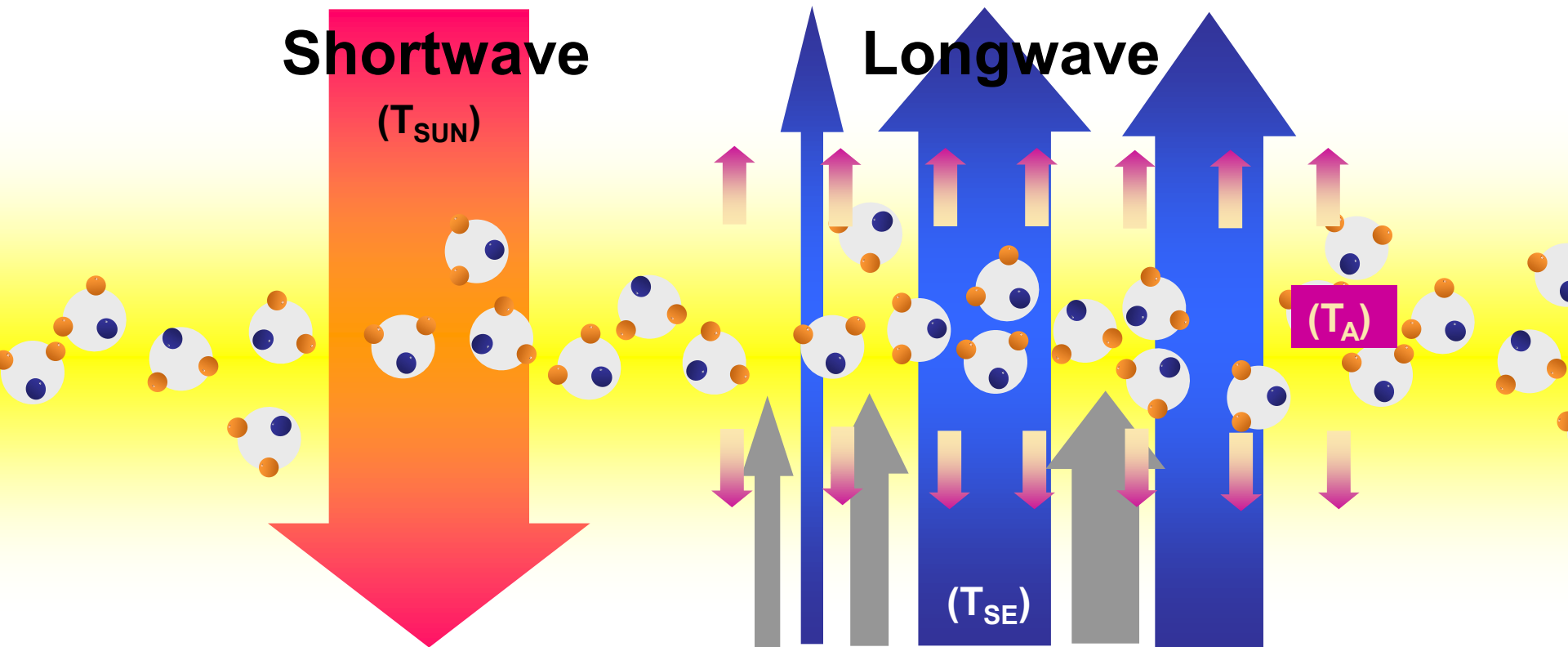
 Carbon

 Oxygen

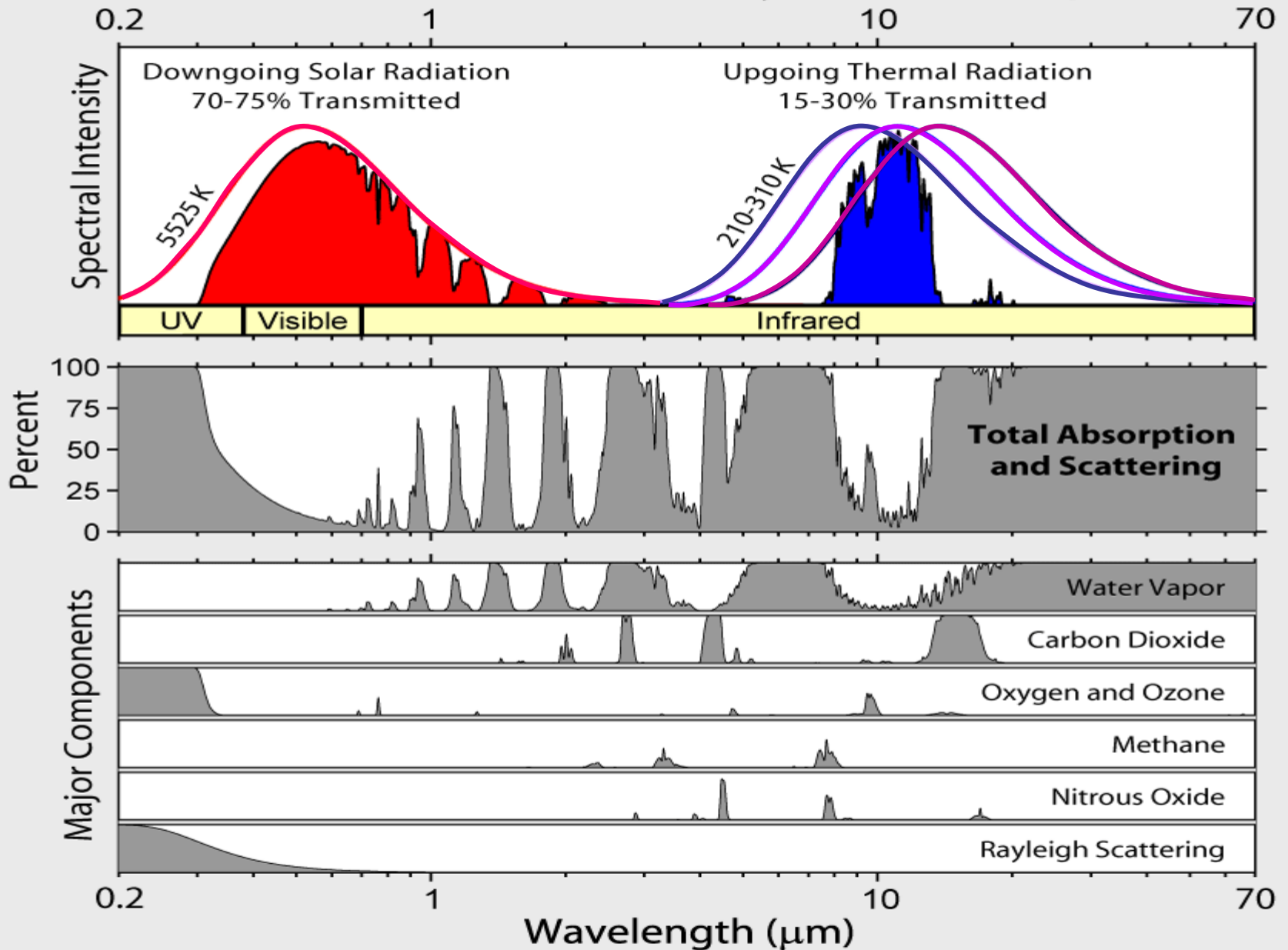
LW Photon 

a CO₂ gas

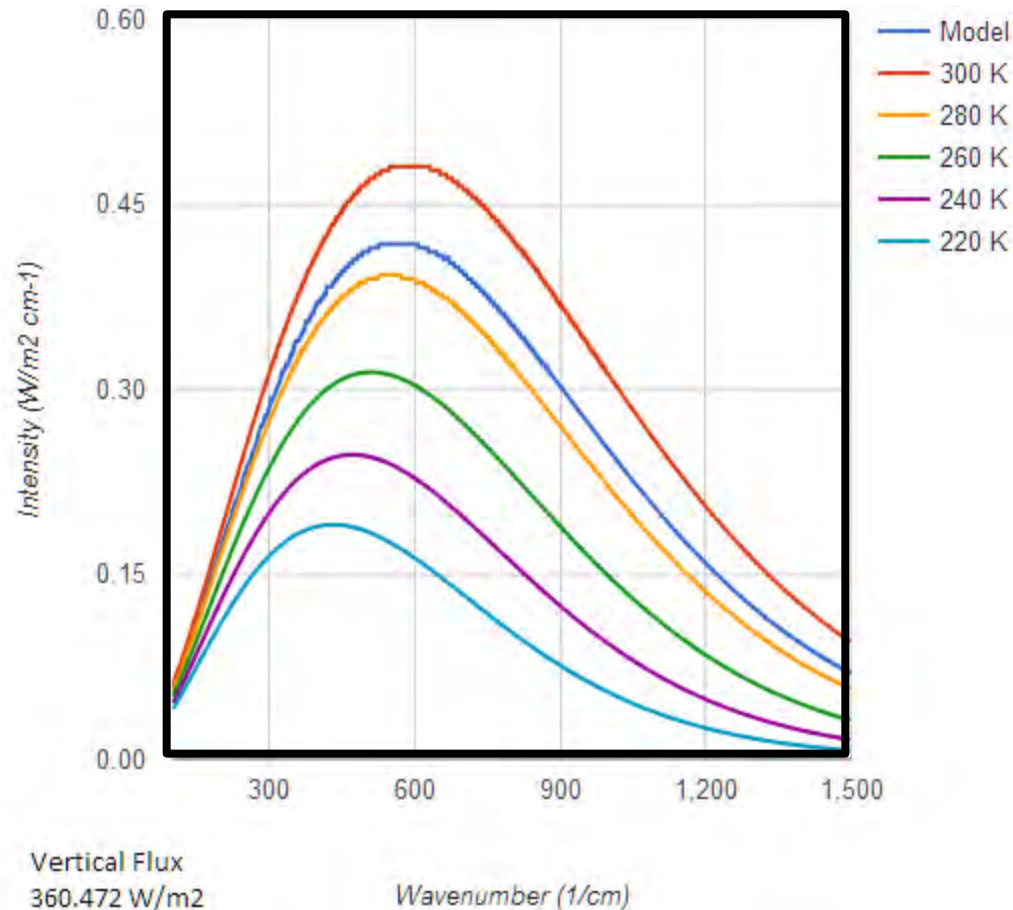




Radiation Transmitted by the Atmosphere



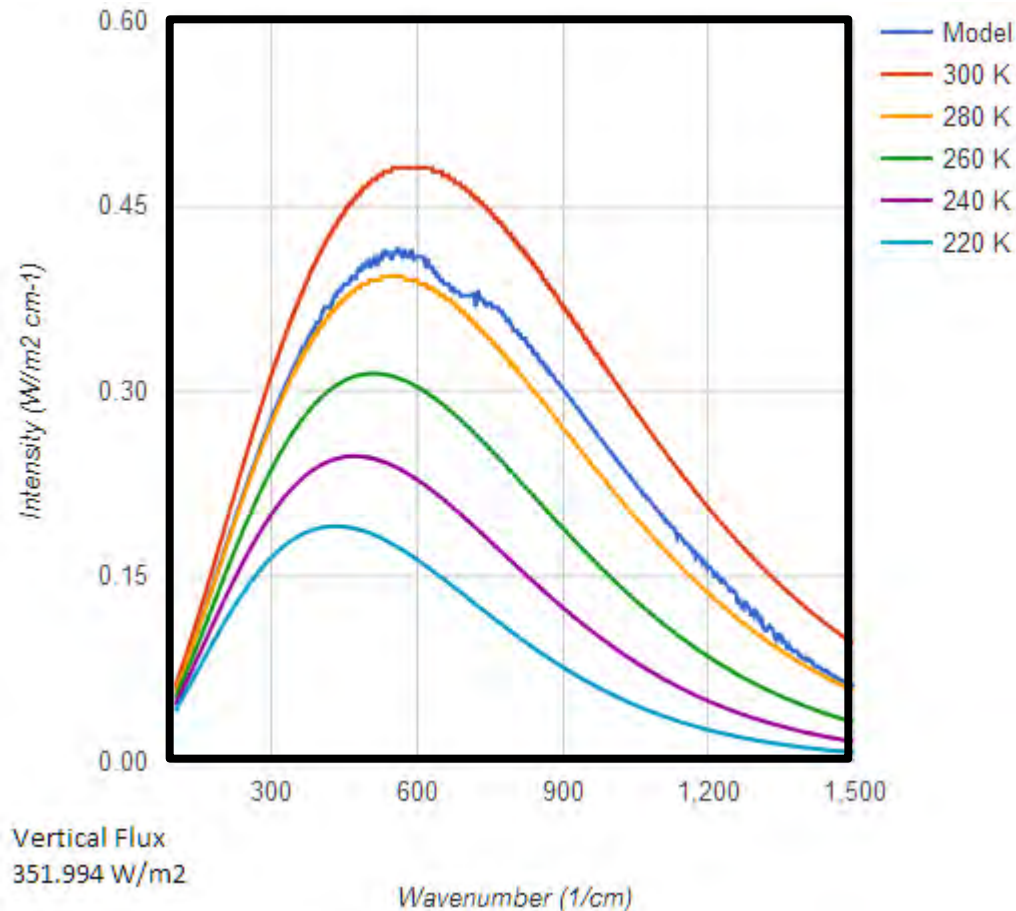
OLR spectrum looking down from $h = 0$ km



how does the Greenhouse effect work?

Modtran 3 v1.3 imulations with U.S. Standard Atmosphere

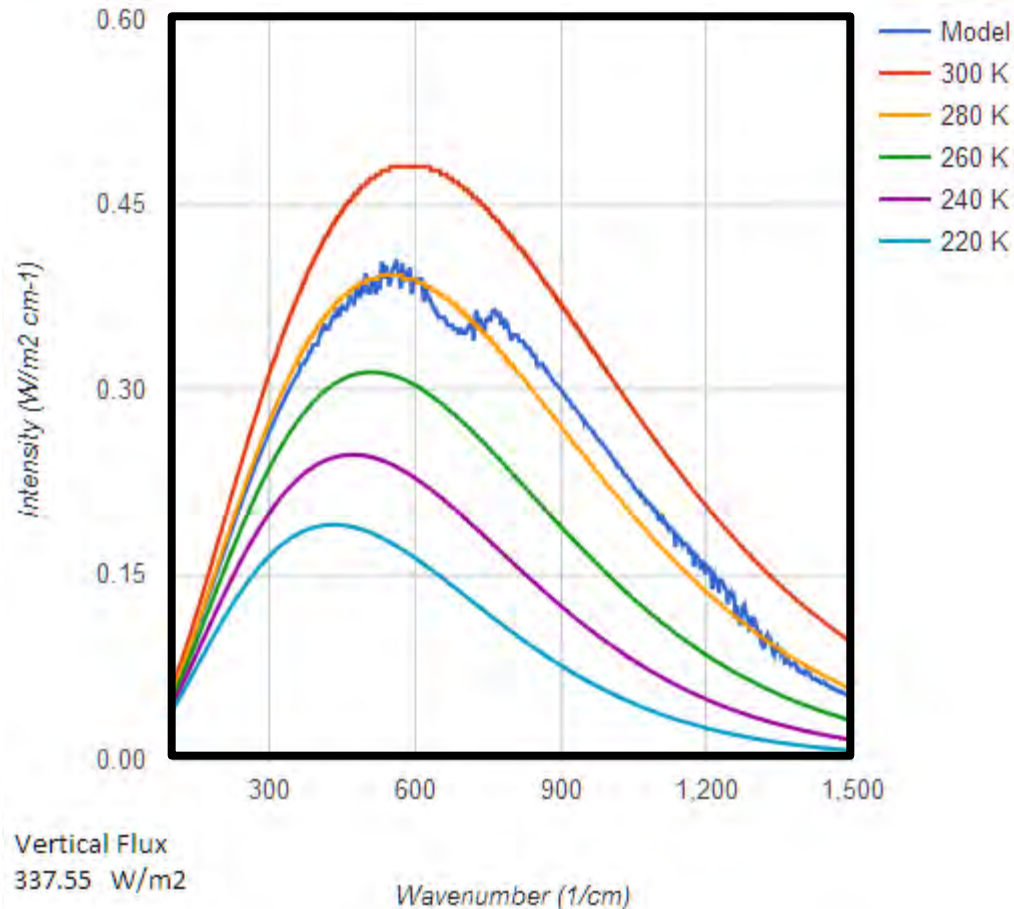
OLR spectrum looking down from $h = 1$ km



how does the Greenhouse effect work?

Modtran 3 v1.3 imulations with U.S. Standard Atmosphere

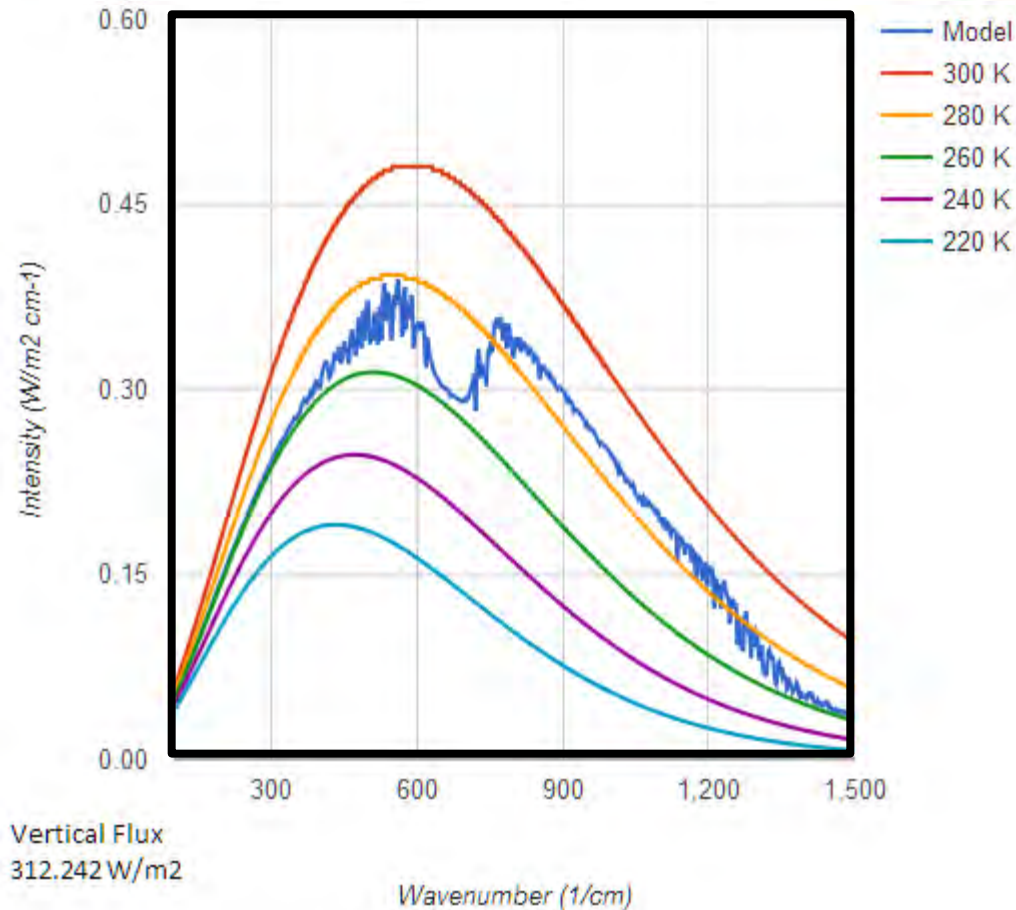
OLR spectrum looking down from $h = 2$ km



how does the Greenhouse effect work?

Modtran 3 v1.3 imulations with U.S. Standard Atmosphere

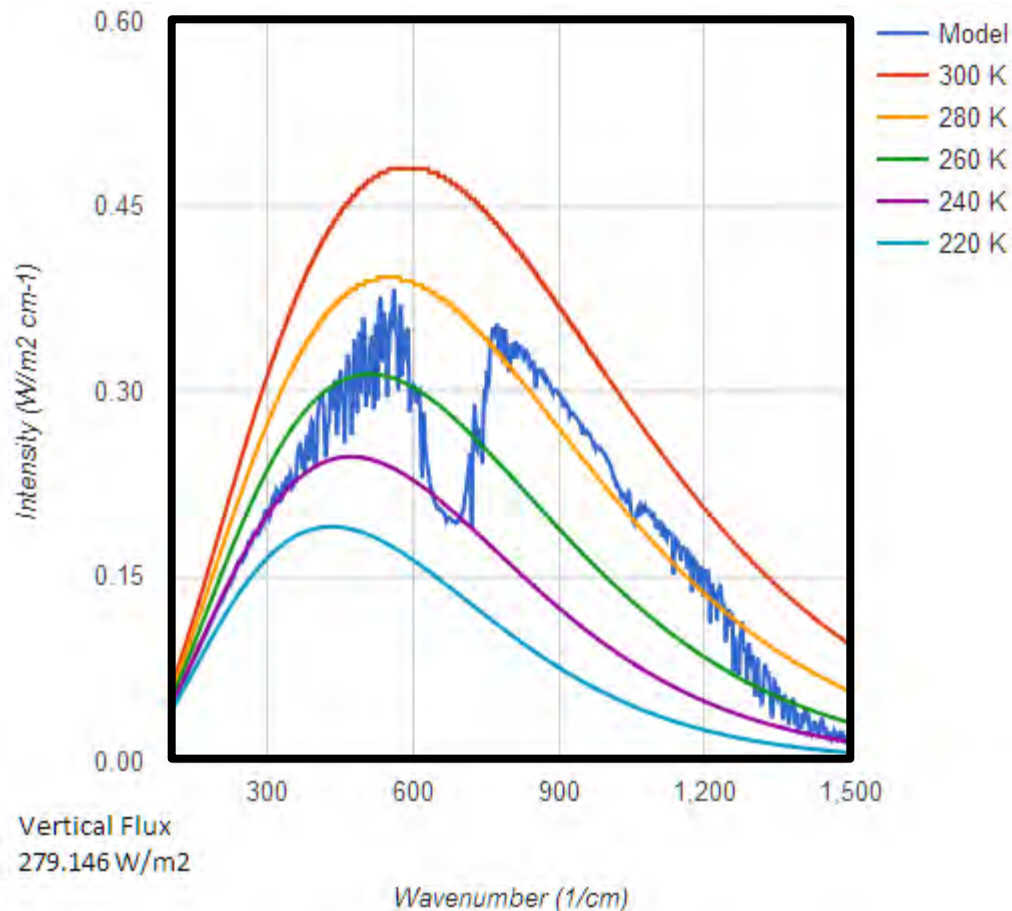
OLR spectrum looking down from $h = 4$ km



how does the Greenhouse effect work?

Modtran 3 v1.3 imulations with U.S. Standard Atmosphere

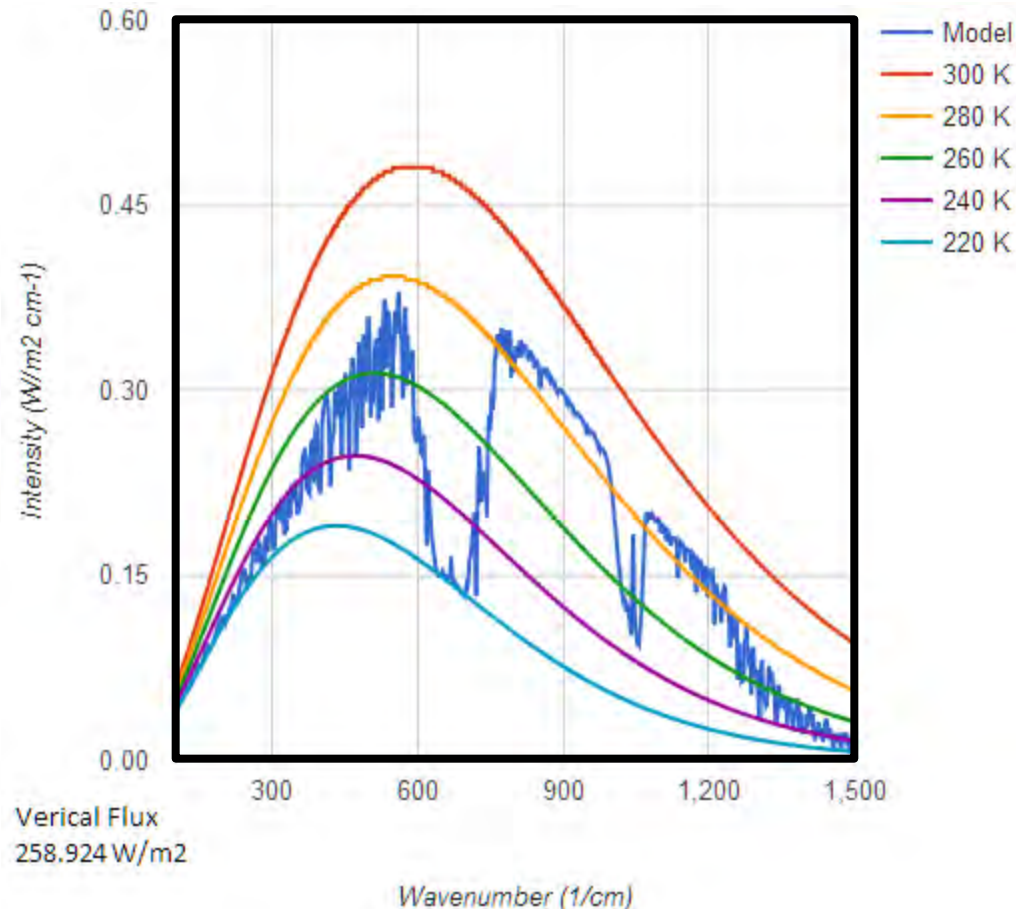
OLR spectrum looking down from $h = 8$ km



how does the Greenhouse effect work?

Modtran 3 v1.3 imulations with U.S. Standard Atmosphere

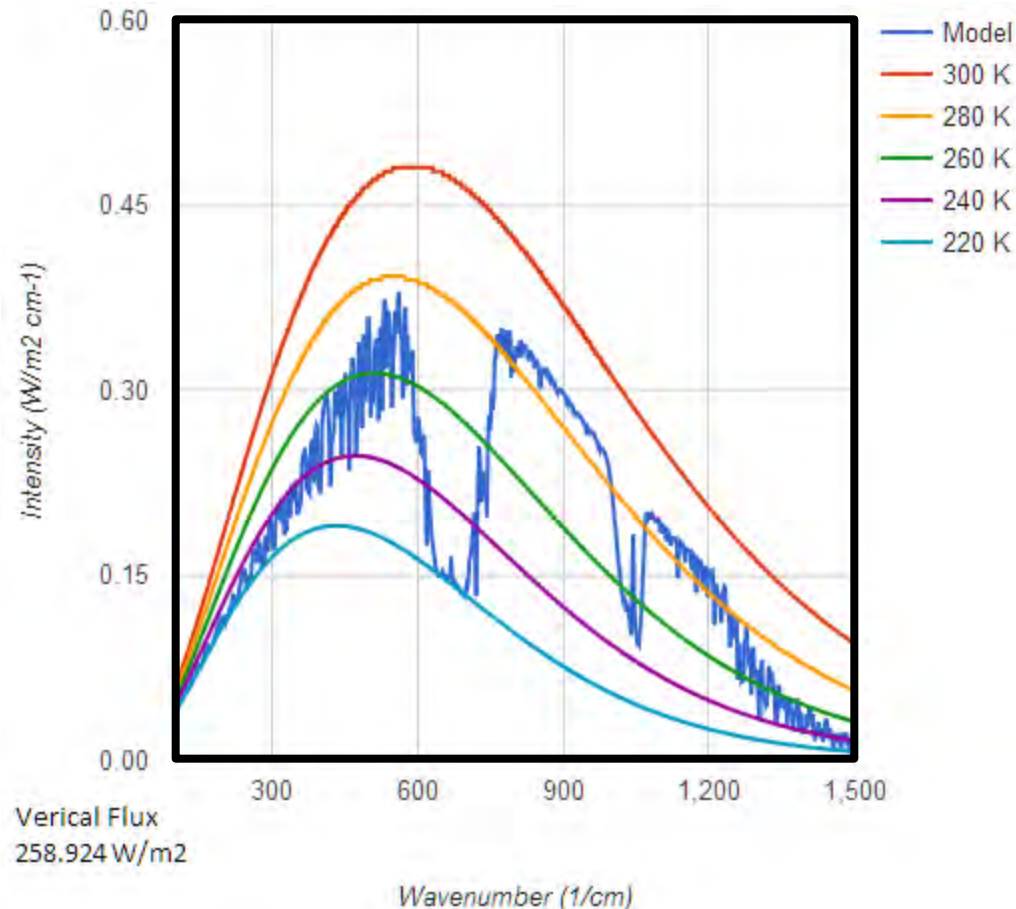
OLR spectrum looking down from $h = 16$ km



how does the Greenhouse effect work?

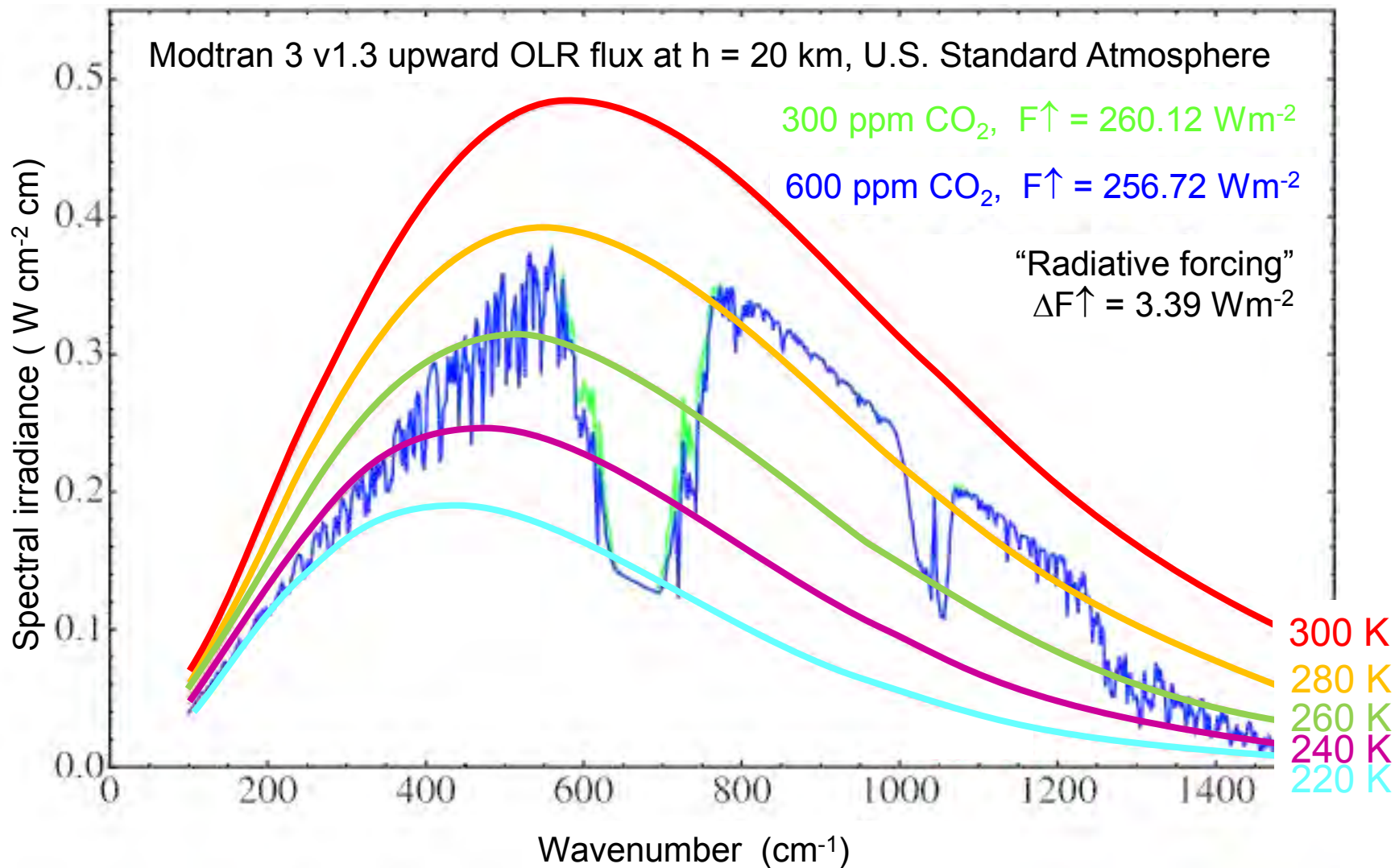
Modtran 3 v1.3 imulations with U.S. Standard Atmosphere

OLR spectrum looking down from $h = 32$ km



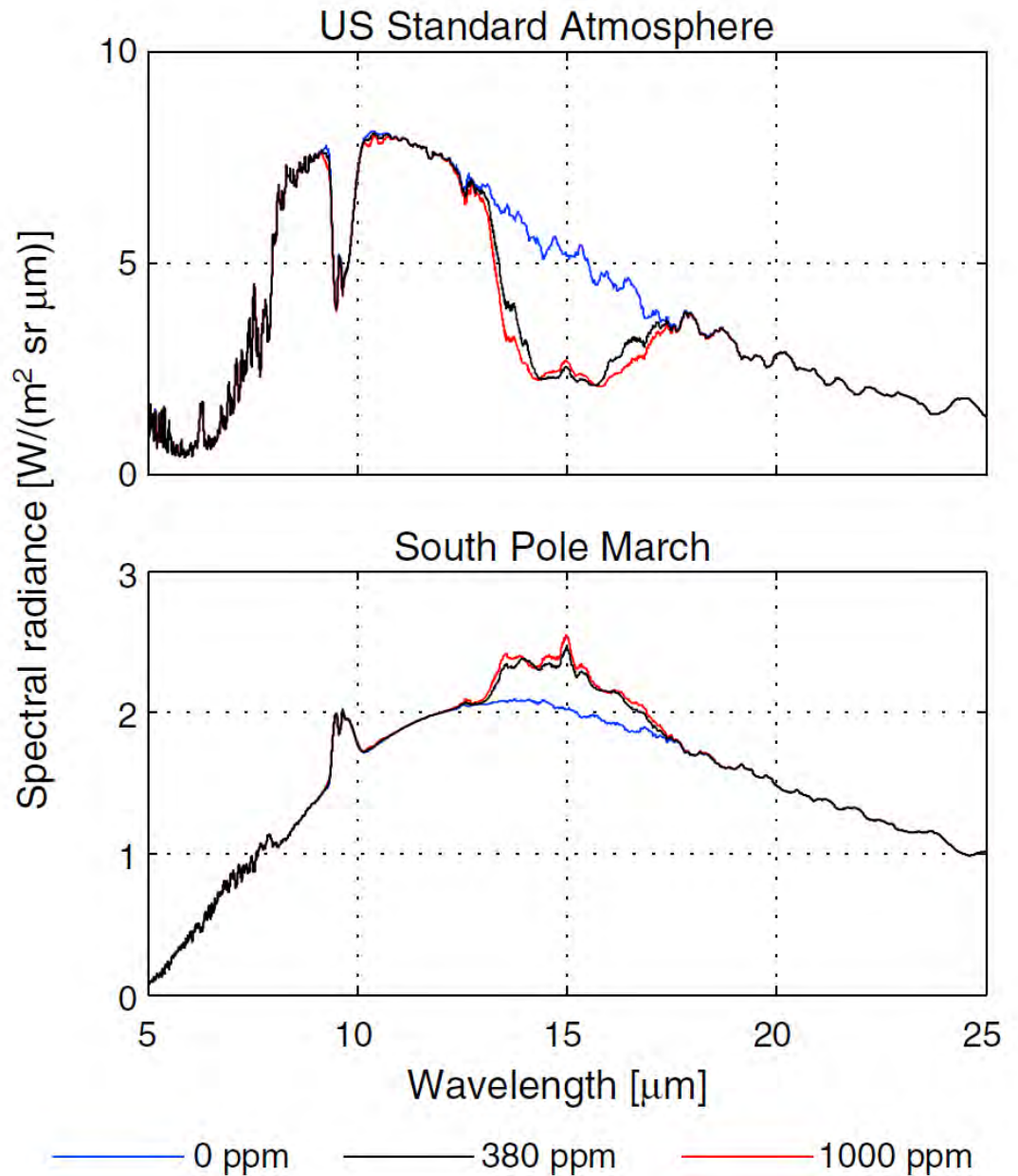
how does the Greenhouse effect work?

Modtran 3 v1.3 imulations with U.S. Standard Atmosphere



Negative greenhouse effect (observed to sometimes happen in Antarctica when atmosphere at 20-30 km is warmer than at surface)

NB. Plotted against wavelength, λ not wavenumber, $k = 1/\lambda$, so main CO₂ line around $k = 675 \text{ cm}^{-1}$ appears at $\lambda = 15 \text{ }\mu\text{m}$



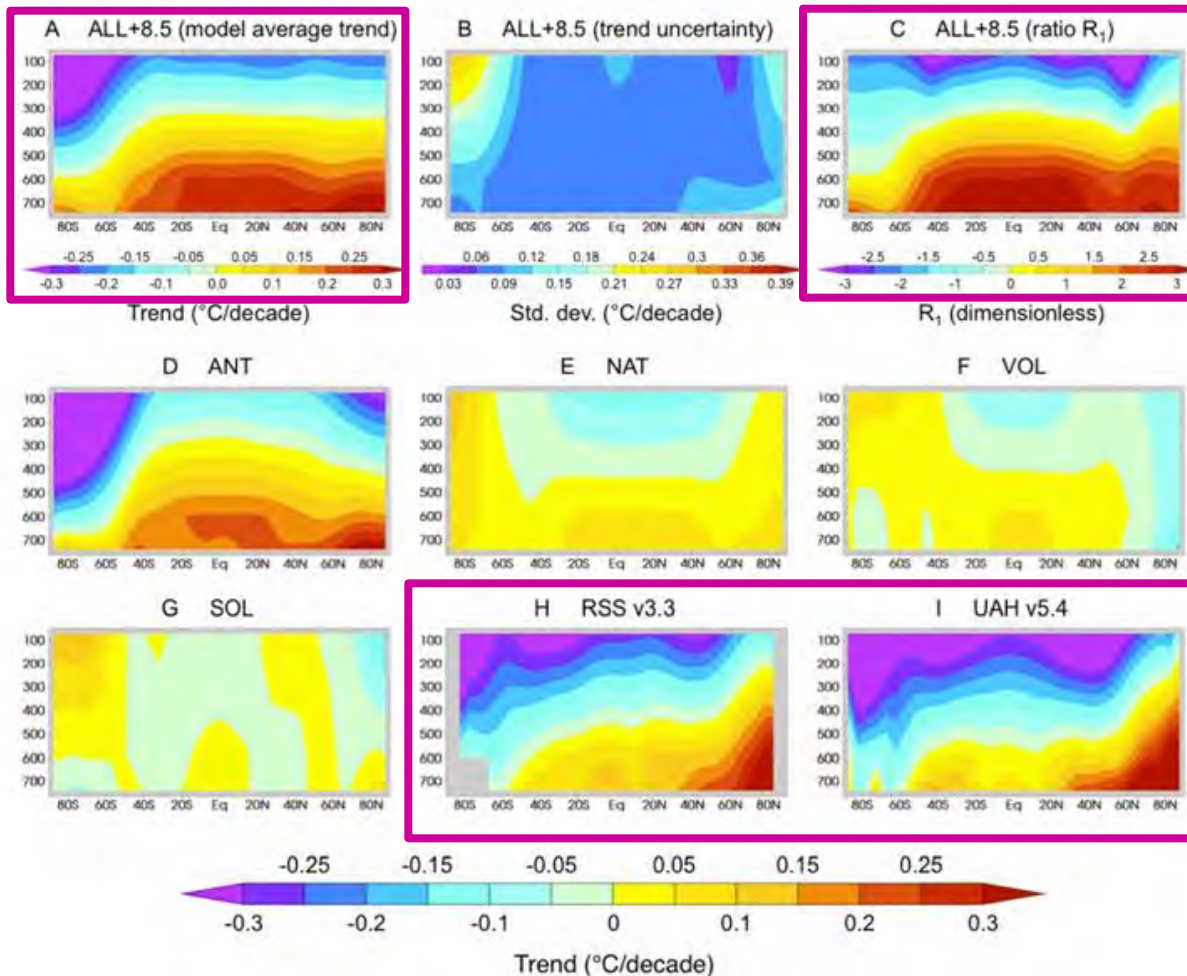
(Schmithüsen et al, GRL, 2015)



Altitude Variations (observed and modelled zonal mean trends in latitude-altitude plots for 1979-2012)

(*Santer et al., 2013*)

Zonal-Mean Atmospheric Temperature Trends in CMIP-5 Models and Observations



Modelled. Forcings:
ANT = anthropogenic
NAT = natural
VOL = Volcanic
SOL = Solar

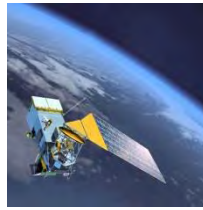
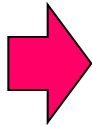
ALL = ANT+NAT
NAT = VOL+SOL

Observed
(RSS and UAH
analysis of
satellite data)

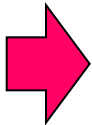
Babies and Bathwater



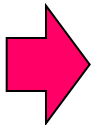
Solar Variability: Effects on Climate?



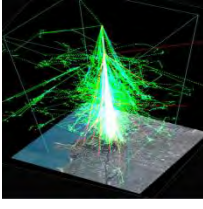
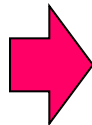
Solar Outputs



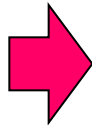
Solar Variability



Global Effects

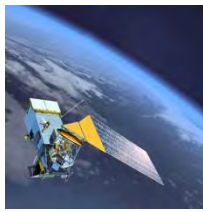
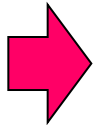


Regional & Seasonal Effects

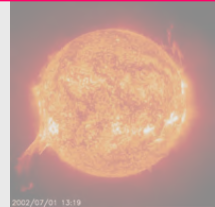


The Future

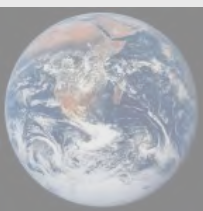
Solar Variability: Effects on Climate?



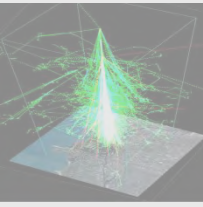
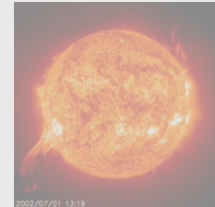
Solar Outputs



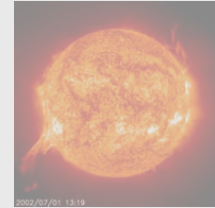
Solar Variability



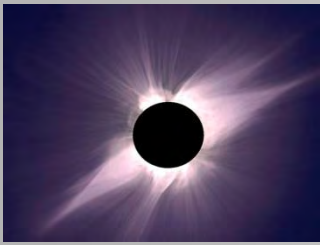
Global Effects



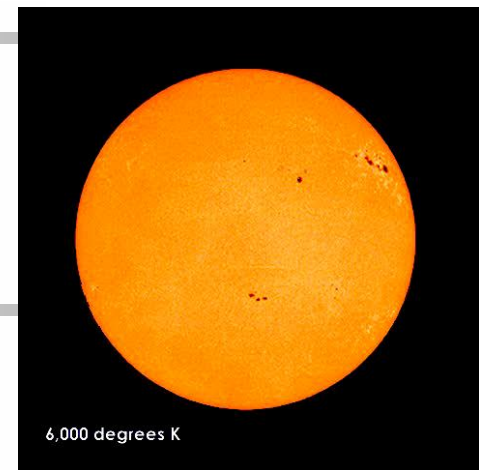
Regional & Seasonal Effects



The Future



Solar Outputs



Visible/IR

weakly modulated ($\sim 0.1\%$) by magnetic field in photosphere

UV

modulated ($\sim 1\%$) by magnetic fields threading the lowest solar atmosphere (chromosphere)

EUV

strongly modulated ($\sim 50\%$) by magnetic fields in the solar atmosphere (corona)

X-Rays

fully dependent on (modulated $\sim 90\%$) by magnetic fields in the solar atmosphere (corona)

Solar wind

$\sim 65\%$ modulated over the solar magnetic cycle

Cosmic Rays

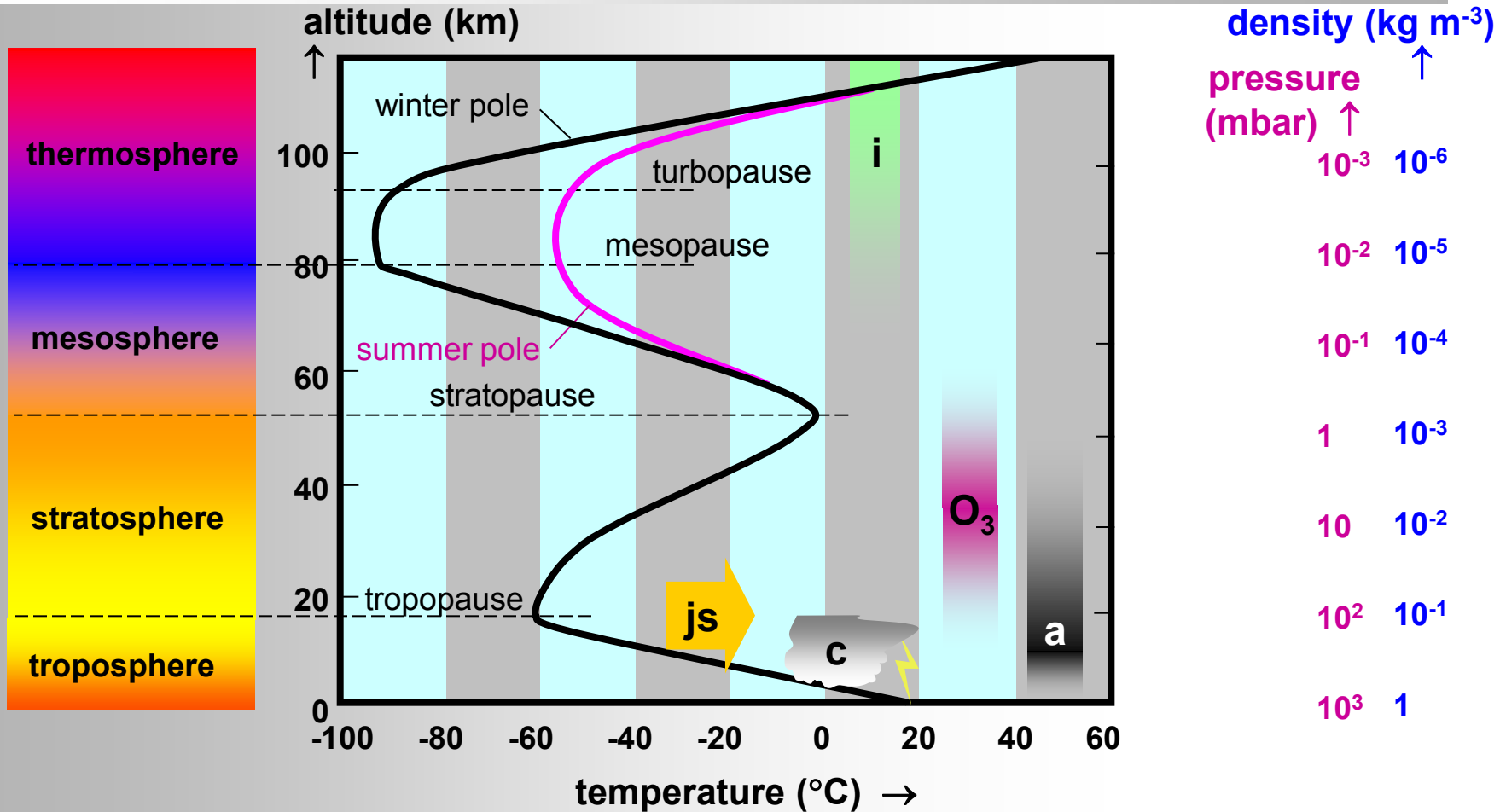
$\sim 20\% - 40\%$ modulated (at 10 - 1 GeV) by solar magnetic field irregularities in heliosphere

SEPs

$\sim 100\%$ modulated by transient magnetic fields in solar flares & ahead of interplanetary coronal mass ejections



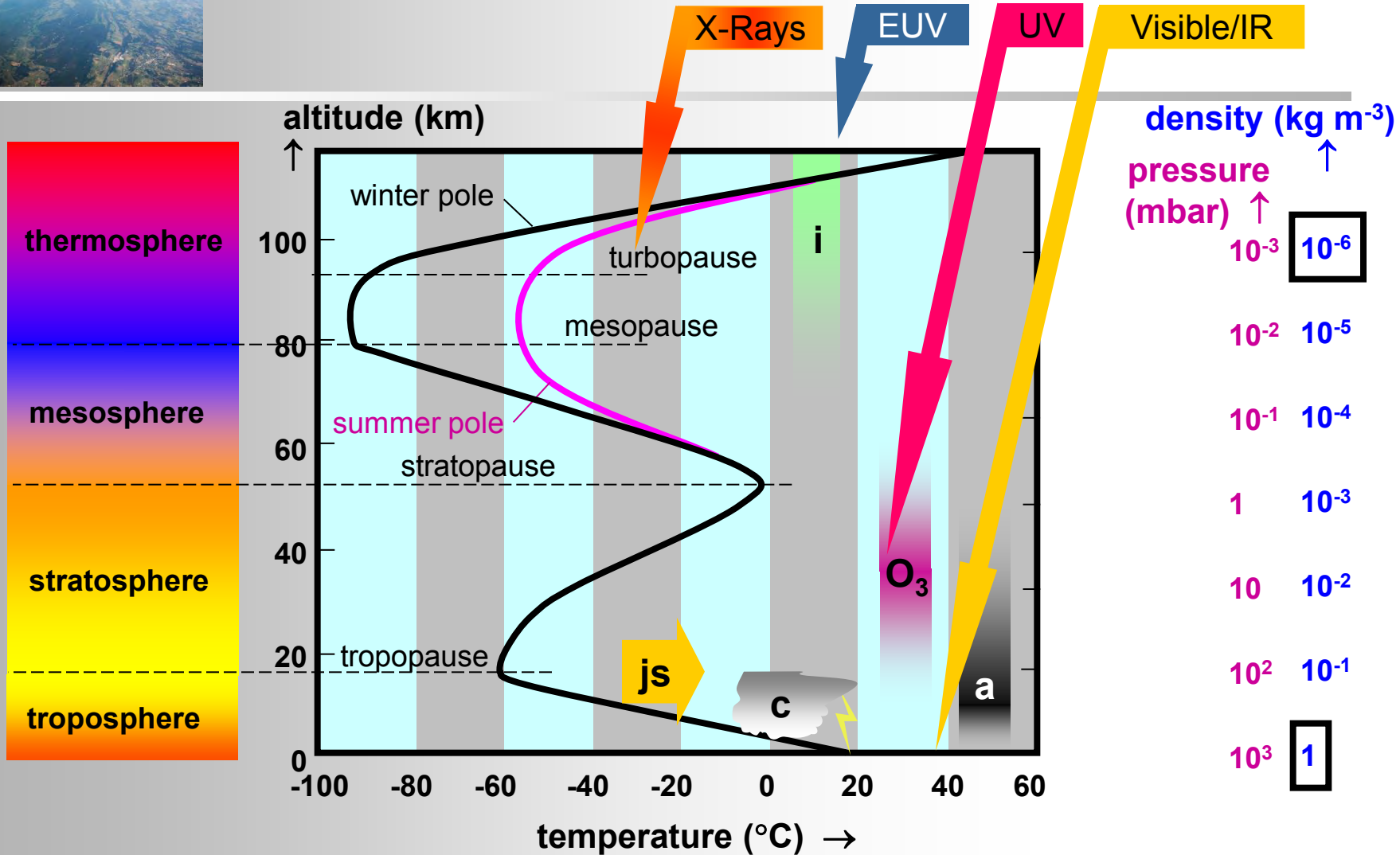
Earth's atmosphere



O_3 = ozone layer a = aerosols js = jet stream
 c = cloud i = ionosphere



Electromagnetic solar inputs



O₃ = ozone layer
c = cloud

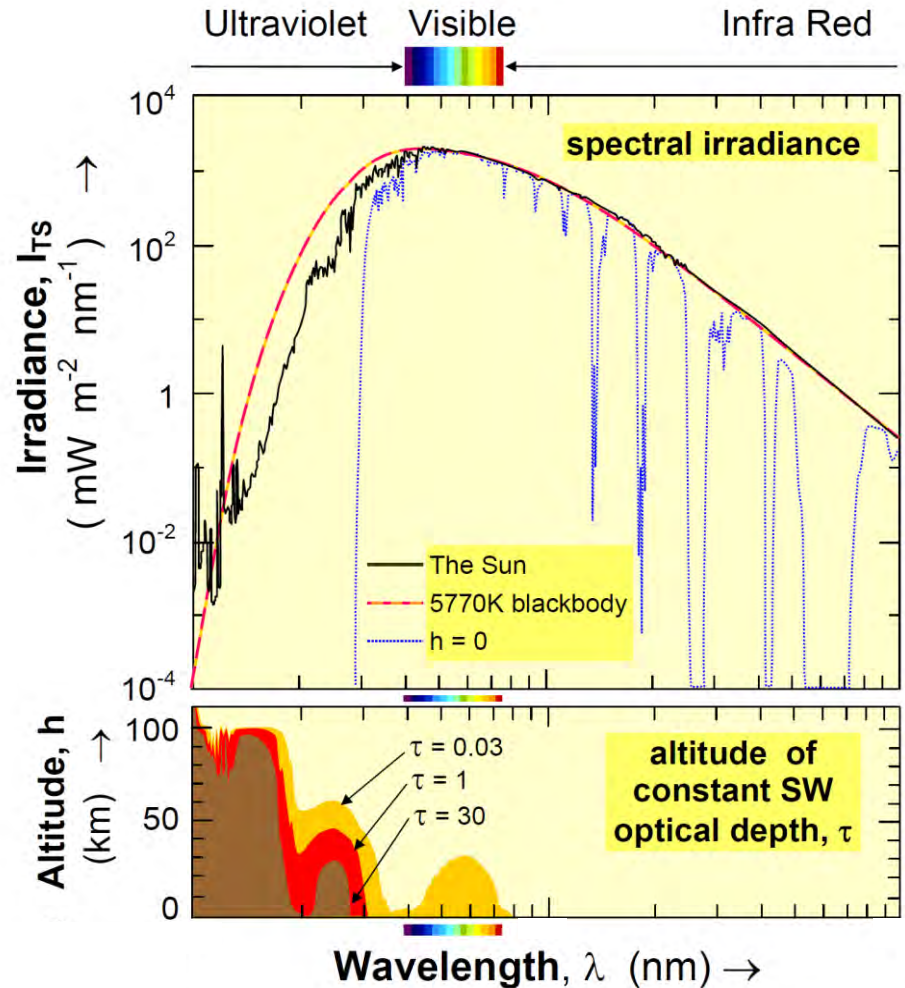
a = aerosols
i = ionosphere

js = jet stream

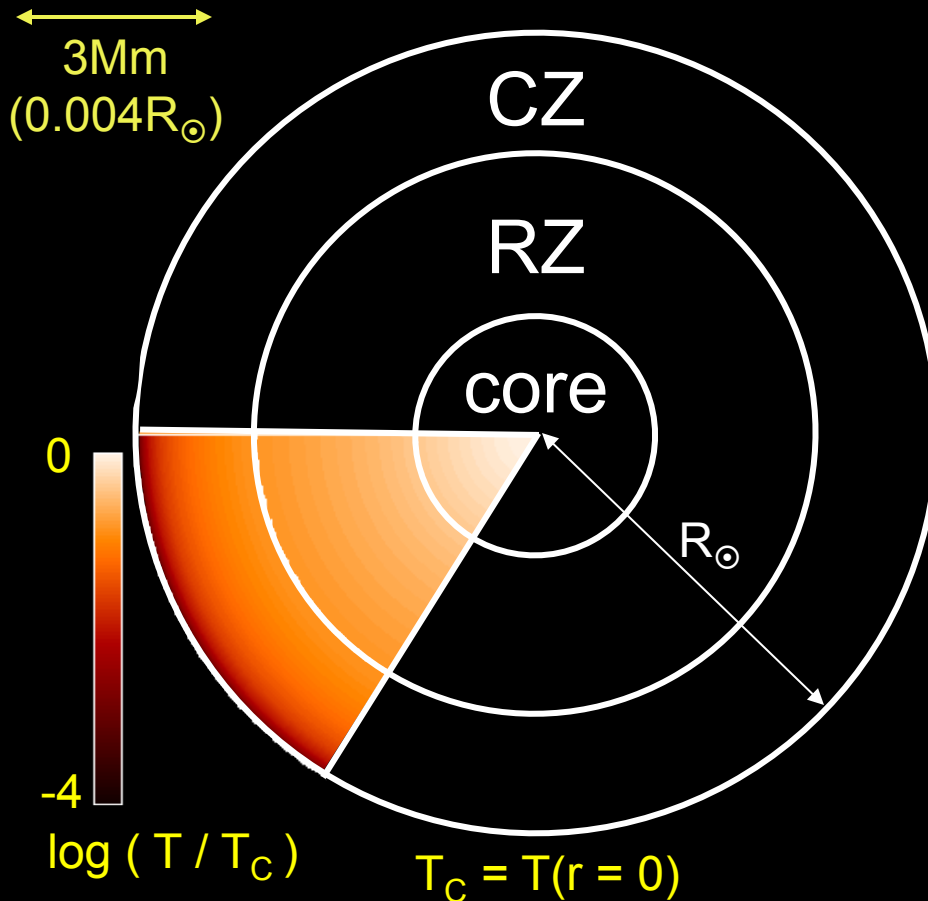
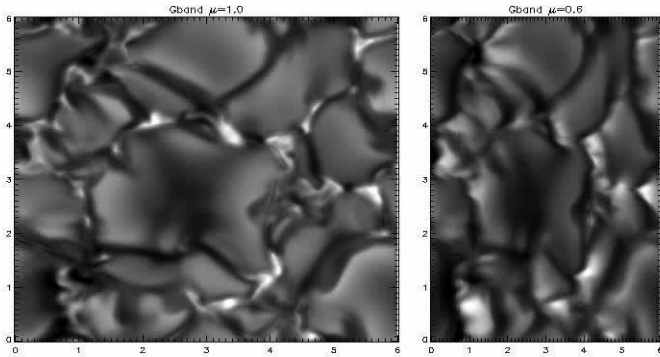


- The Sun's e-m radiation spectrum
- Close to a 5770K blackbody radiator
- Emitted flux

$$F = \varepsilon \sigma T_{\text{sun}}^4$$
- $\varepsilon \approx 1$ and surface temperature of Sun
 $T_S = 5770\text{K}$



Implications of high CZ mass



CZ contains $\sim 3 \times 10^{28} \text{ kg}$ ($M_{\odot}/60$)

thermal timescale of the CZ as a whole = timescale for its warming or cooling, $\tau \approx 10^5 \text{ yr}$

Switch off source at base of CZ and in $t = 100 \text{ yr}$, T_{sun} changes by $1 - \exp(t/\tau) = 0.001$

$$F = \varepsilon \sigma T_{\text{sun}}^4 \quad \text{so that}$$

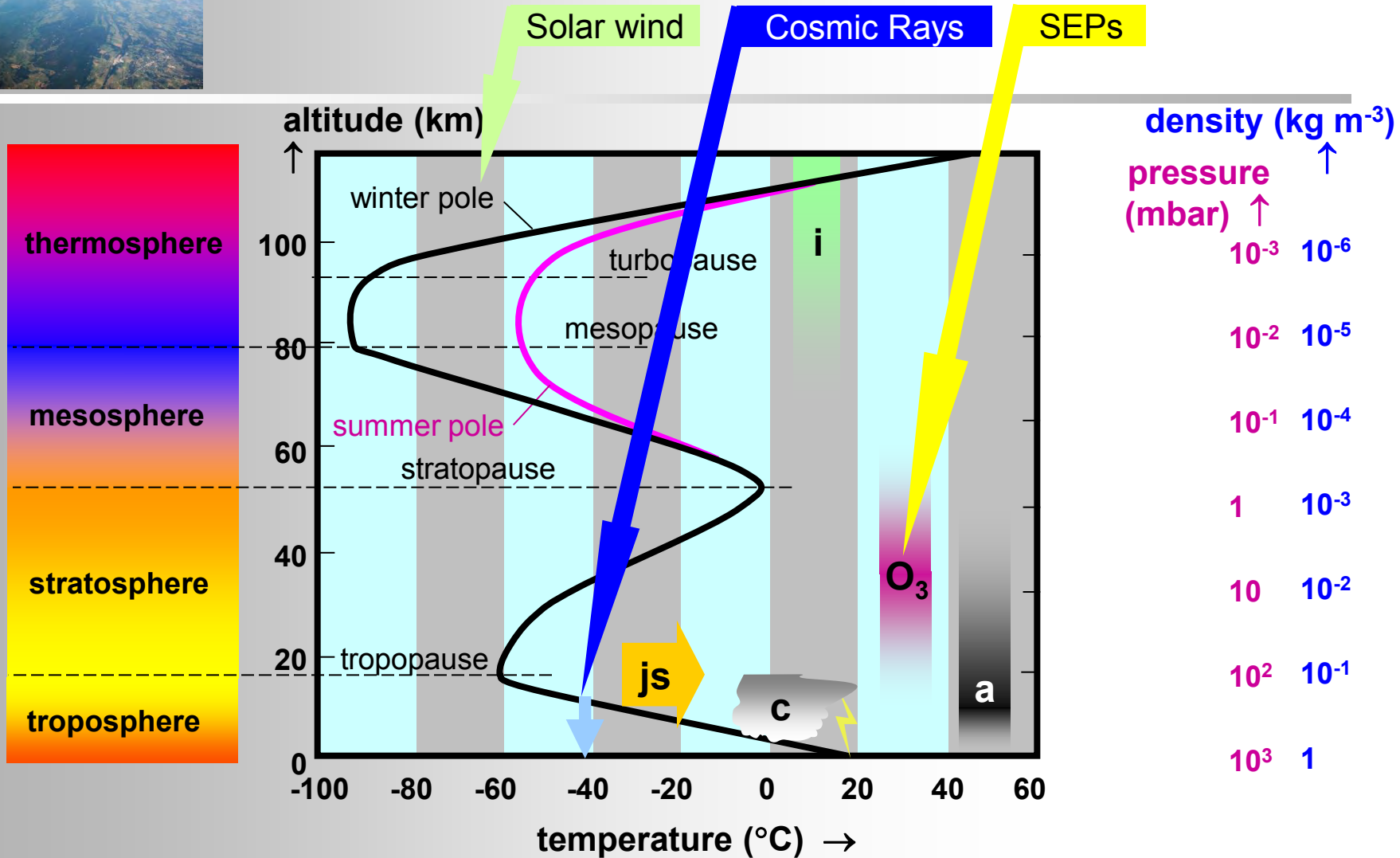
$$F'/F = (T_{\text{sun}}'/T_{\text{sun}})^4$$

$$= 0.999^4 = 0.996$$

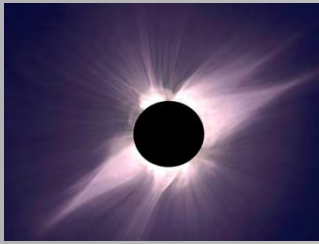
i.e. F changes by just 0.4%



Corpuscular solar inputs



O ₃ = ozone layer	a = aerosols	js = jet stream
c = cloud	i = ionosphere	



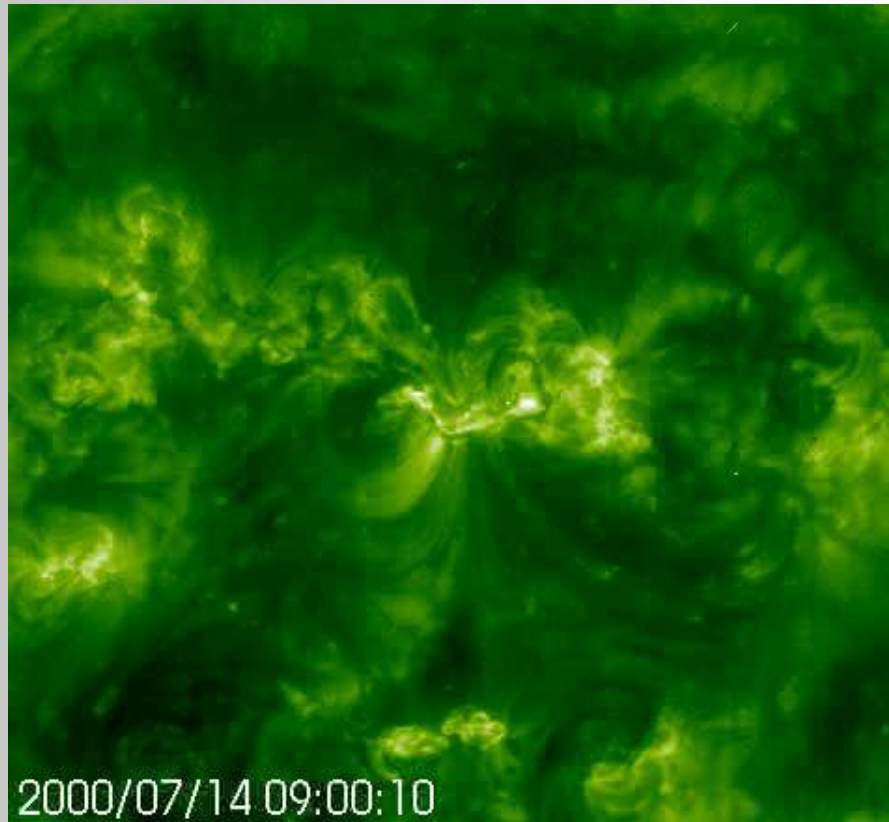
The Bastille Day Storm

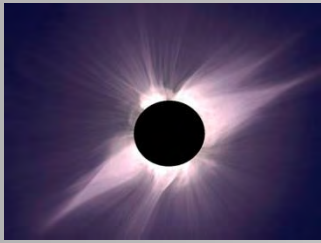
Flare and SEPs



Solar Terrestrial Physics
Summer School

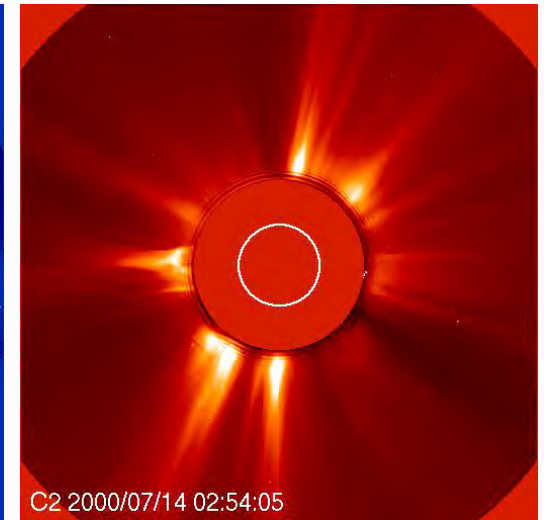
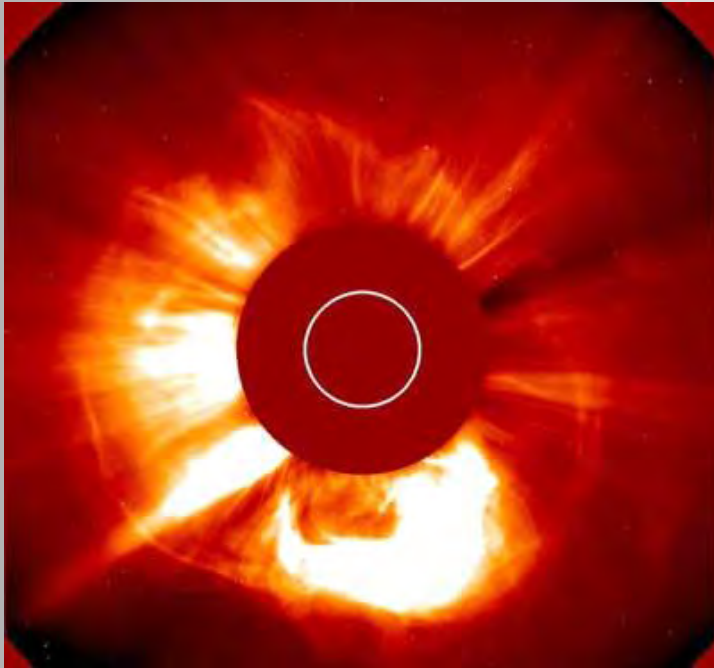
- CME hit Earth on 14th July 2000
- Start of the Story: the associated flare



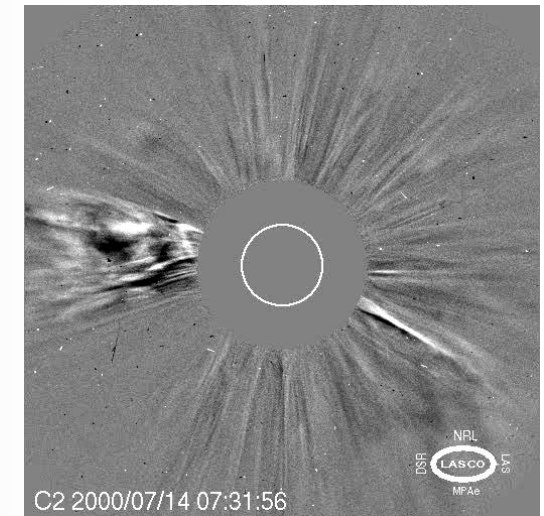


The Bastille Day Storm CME

seen by SoHO/Lasco C2 and C3 Coronagraphs



- “Halo”
(Earthbound)
form most
easily seen in
C2 difference
movie ▶

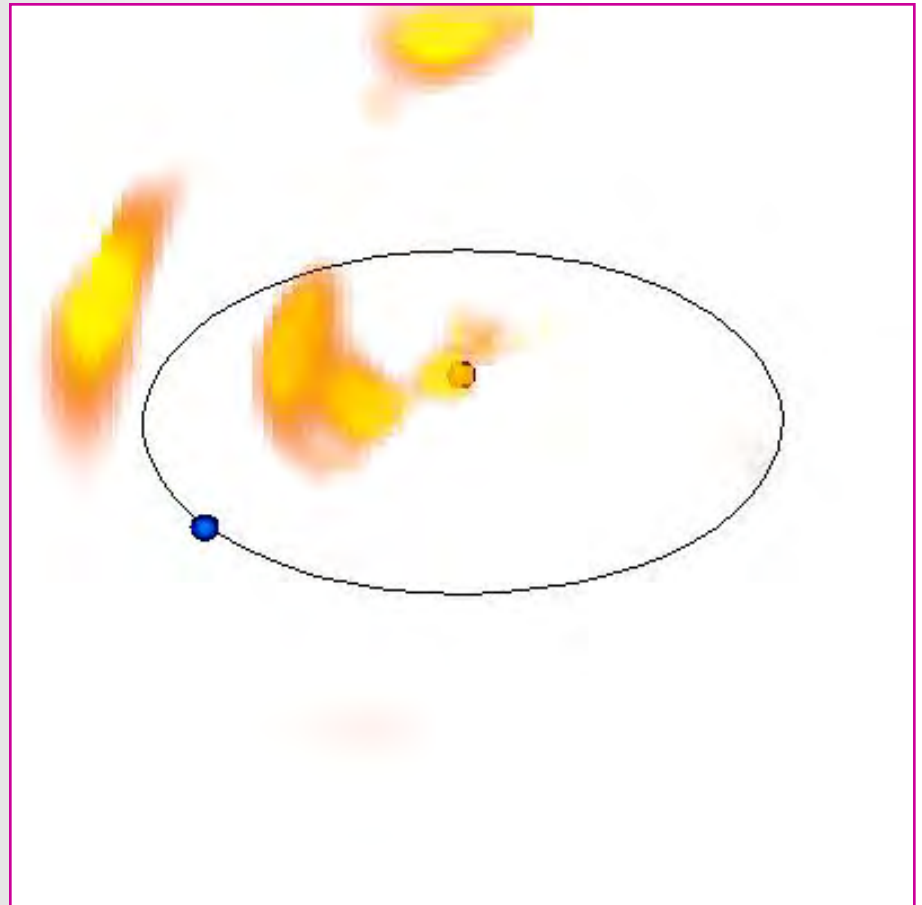


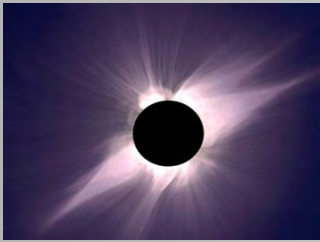


The Bastille Day Storm

CMEs seen by IPS

- Tomographic reconstruction from interplanetary scintillations





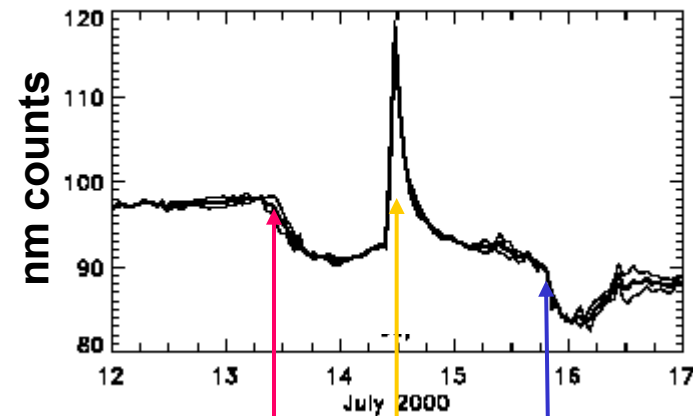
The Bastille Day Storm

GCRs and SEPs



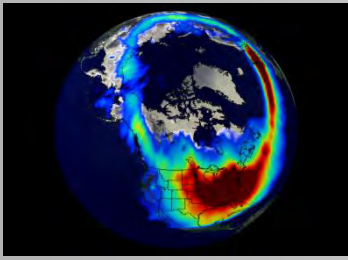
- Ground-level enhancement (GLE) of solar energetic particles seen between Forbush decreases of galactic cosmic rays caused by shielding by the two CMEs
- Here seen at stations in both poles (McMurdo and Thule)

Neutron Monitor counts



Forbush decrease caused by 1st CME

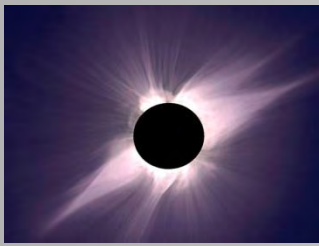
Forbush decrease caused by CME associated with GLE



The Bastille Day Storm

SEP Proton Aurora – seen by Image FUV-SI12



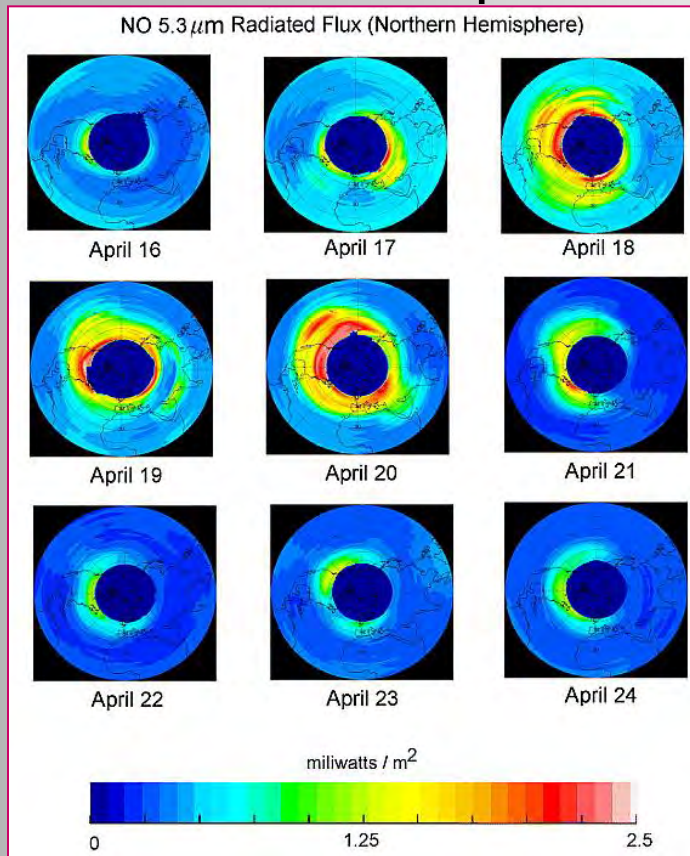


Polar Cap NO

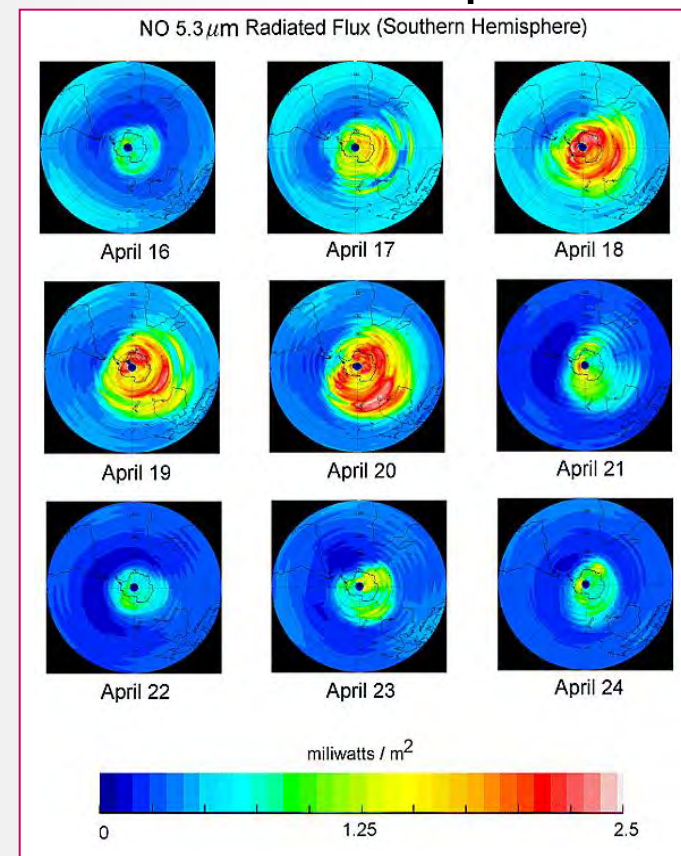
From SEP event of April 2002



▶ Northern hemisphere



▶ Southern hemisphere



TIMED observations of 5.3 μm NO radiative fluxes (Wm^2)
(*Mlynczak et al., 2003*)

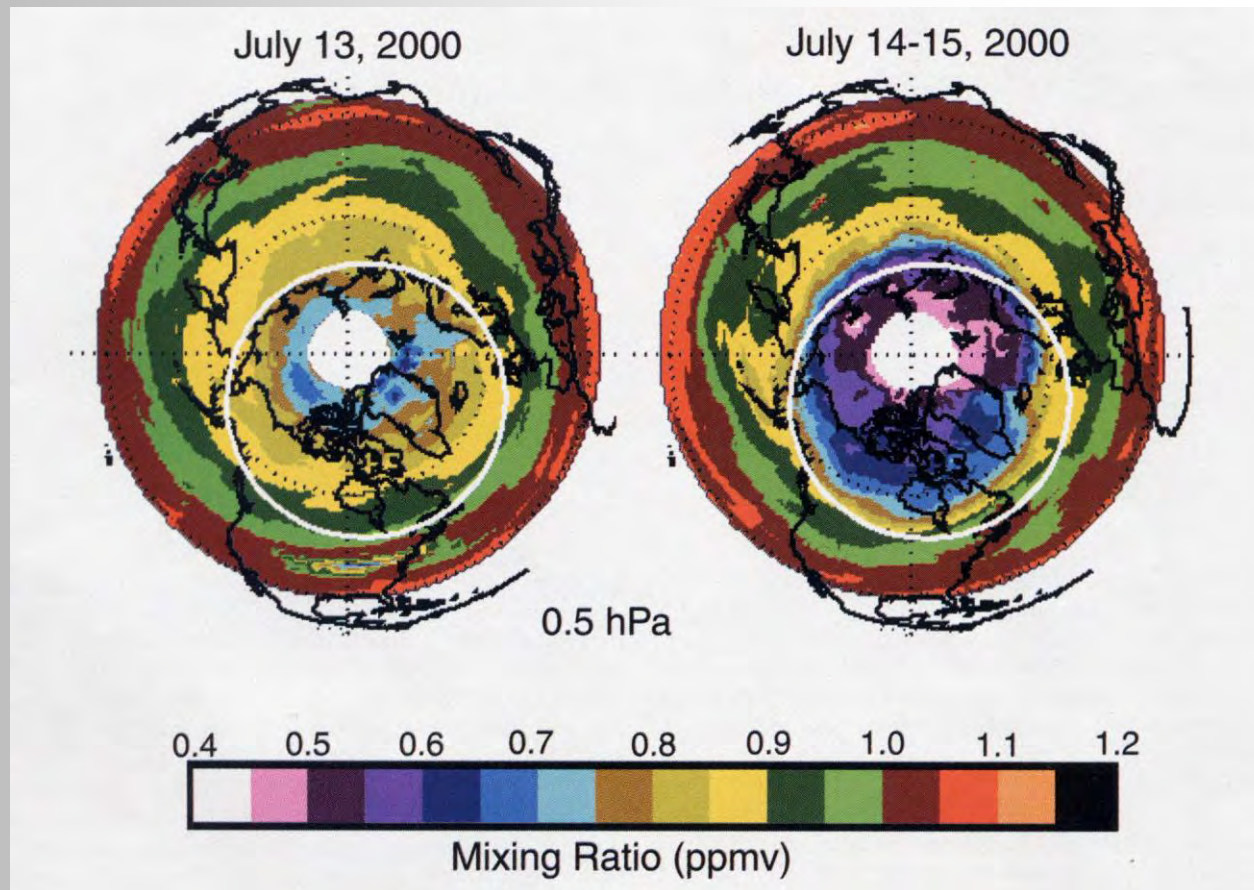


The Bastille Day Storm

Ozone Depletion (TOMS)



Storm Event – SEP Ozone Depletion



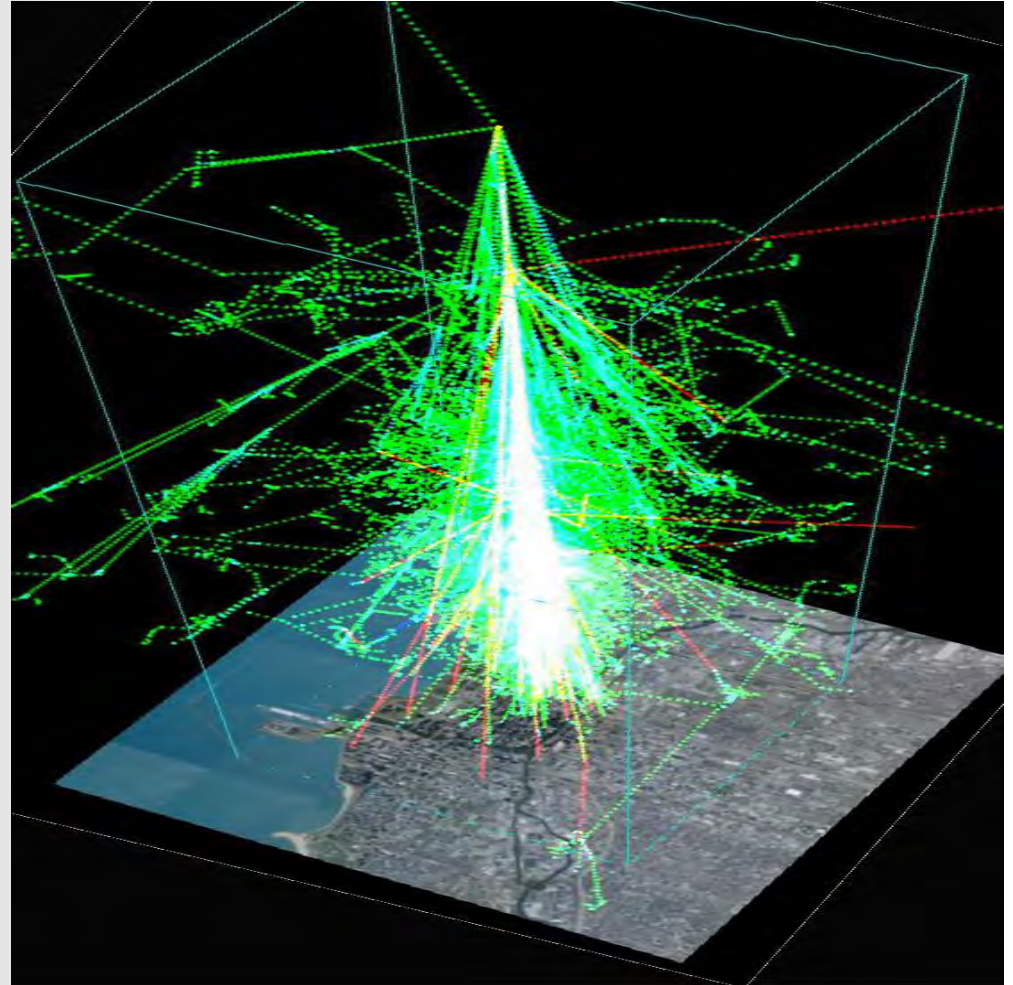


Energetic Particles

Galactic Cosmic Rays



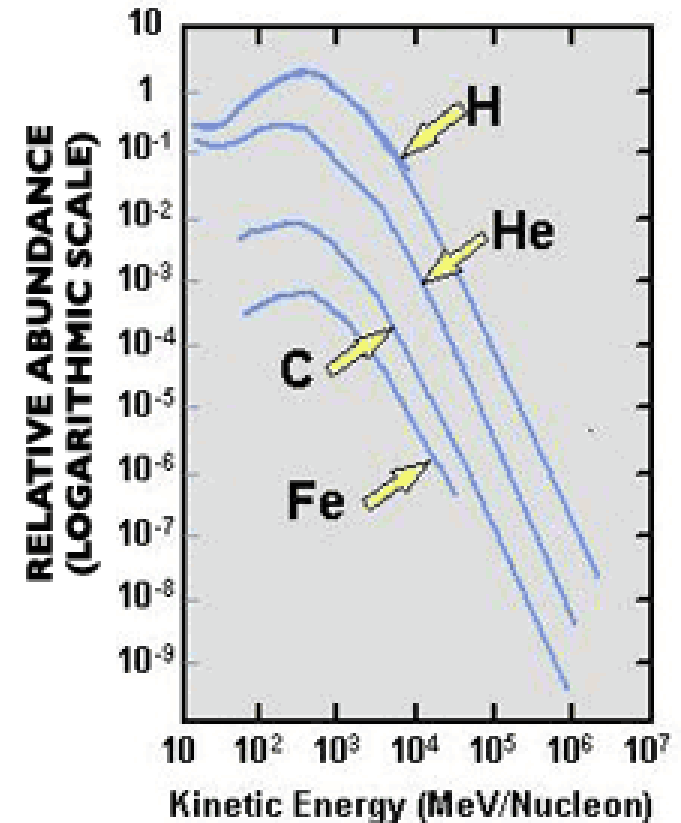
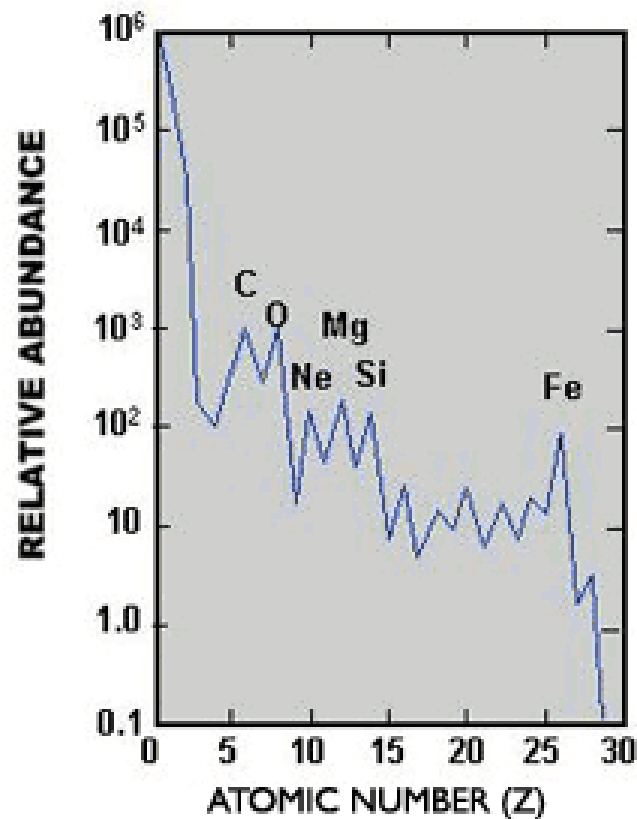
- Generated at the shock fronts ahead of supernovae
- Protons up to iron ions, travelling at close to speed of light
- Three shields protect us on Earth's surface:
 - The heliospheric field
 - Earth's magnetic field
 - Earth's atmosphere

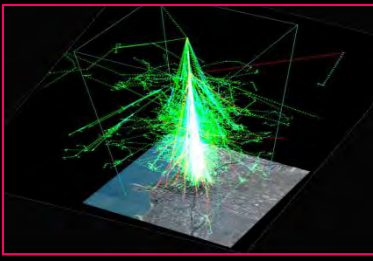


Galactic Cosmic Ray Spectra

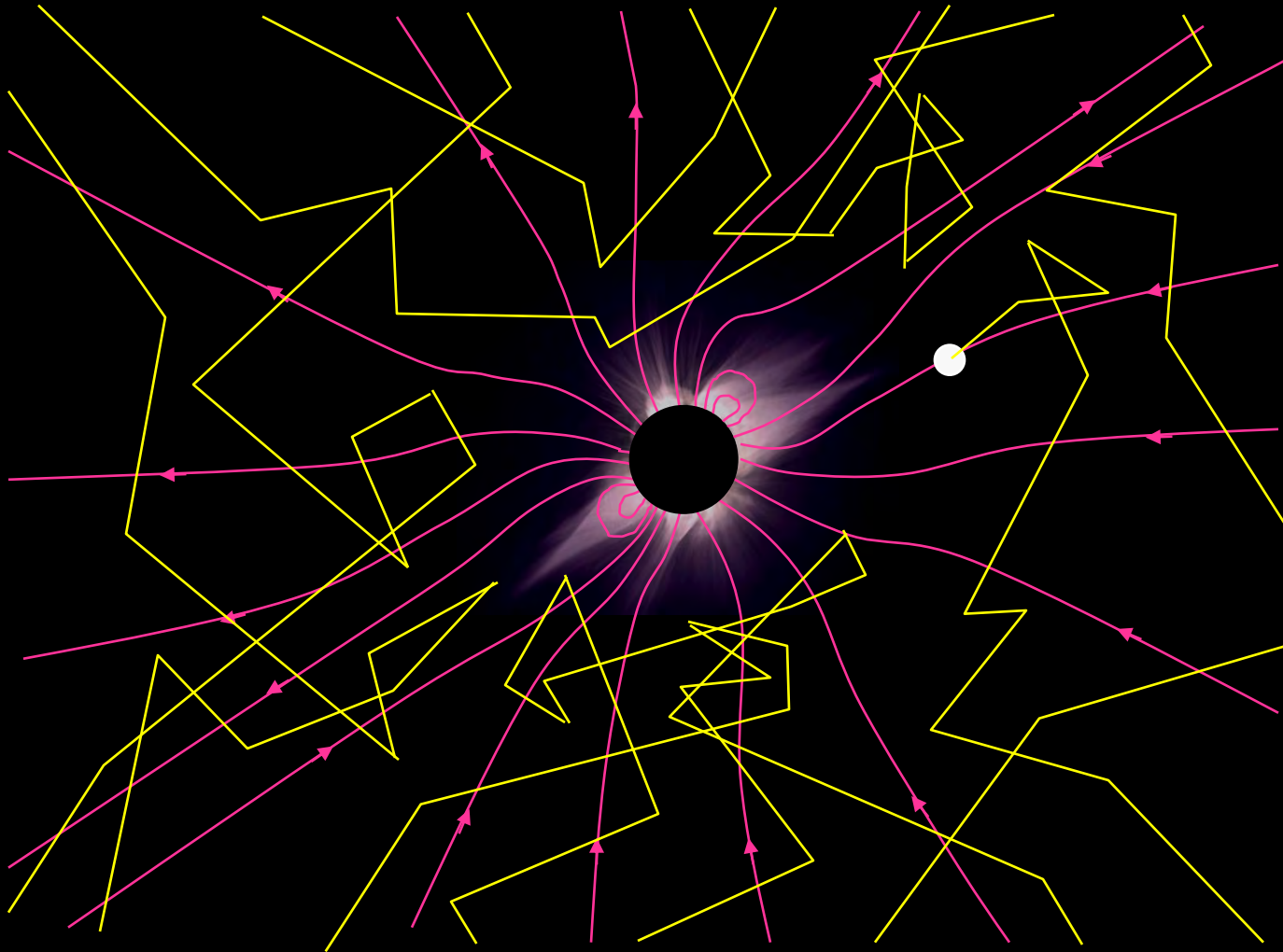


HZE PARTICLE ABUNDANCE AND ENERGY DISTRIBUTIONS





Galactic Cosmic Rays

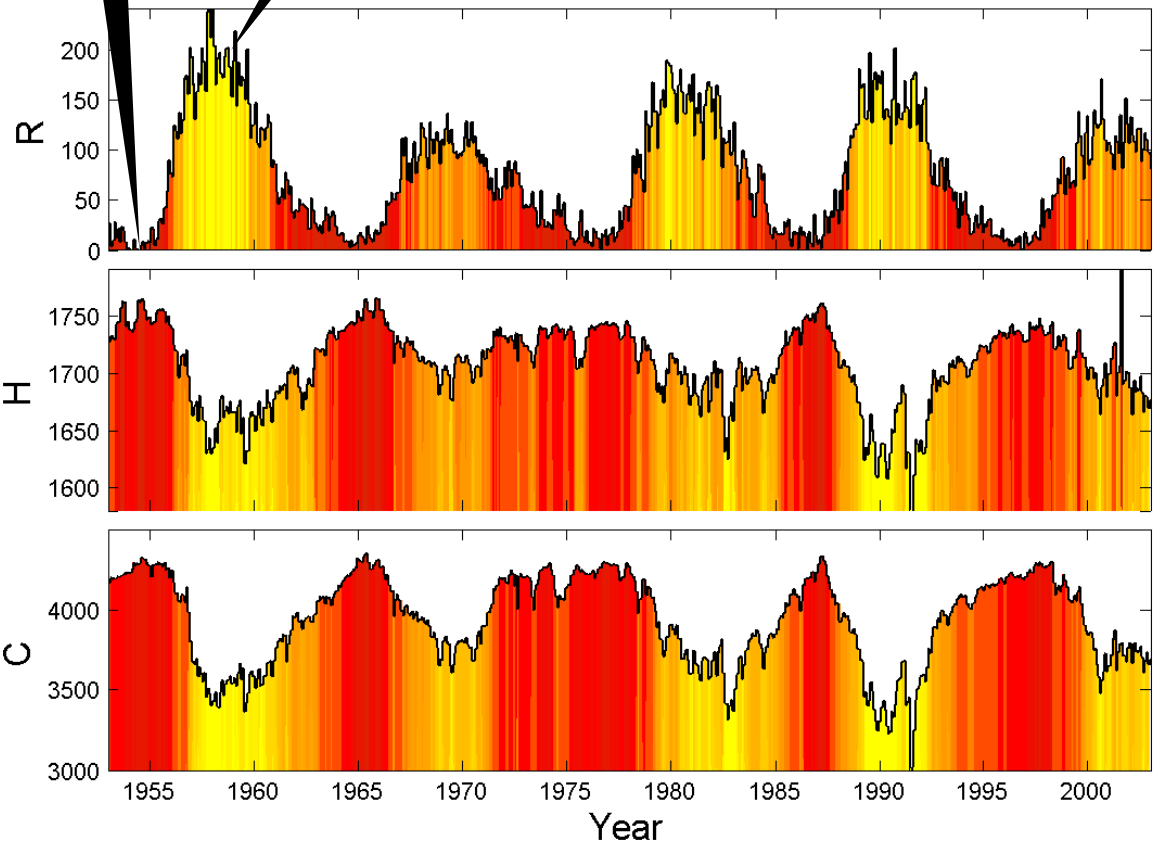


The coronal source flux is dragged out by the solar wind flow to give the heliospheric field which shields Earth from galactic cosmic rays



Cosmic Rays

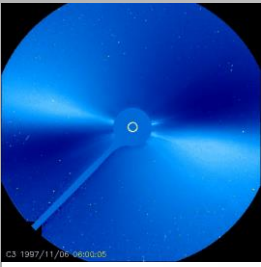
Anticorrelation with sunspot numbers



Sunspot Number

**Huancauyo –
Hawaii neutron
monitor counts
(>13GV)**

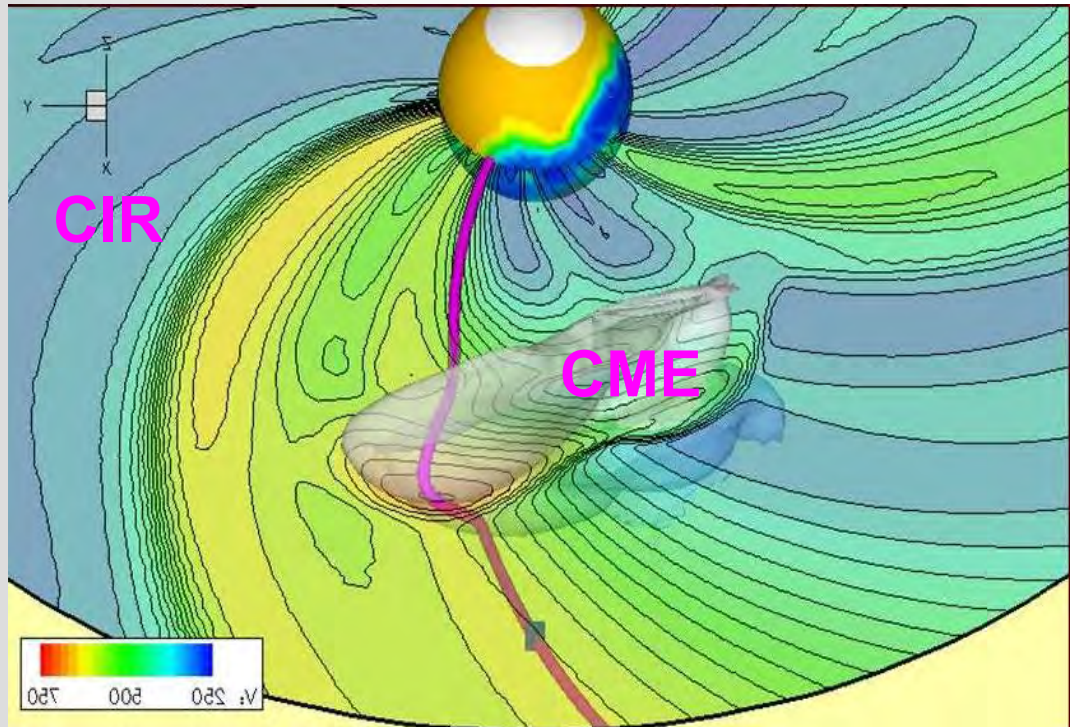
**Climax neutron
monitor counts
(>3GV)**

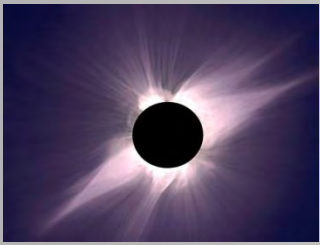


CMEs, CIRs, GCRs and SEPs



- Both CME fronts and CIRs shield Earth from Galactic Cosmic Rays by scattering
- Both CME fronts and CIRs generate SEPs
- Both CMEs and CIRs are more common and more extensive at sunspot maximum

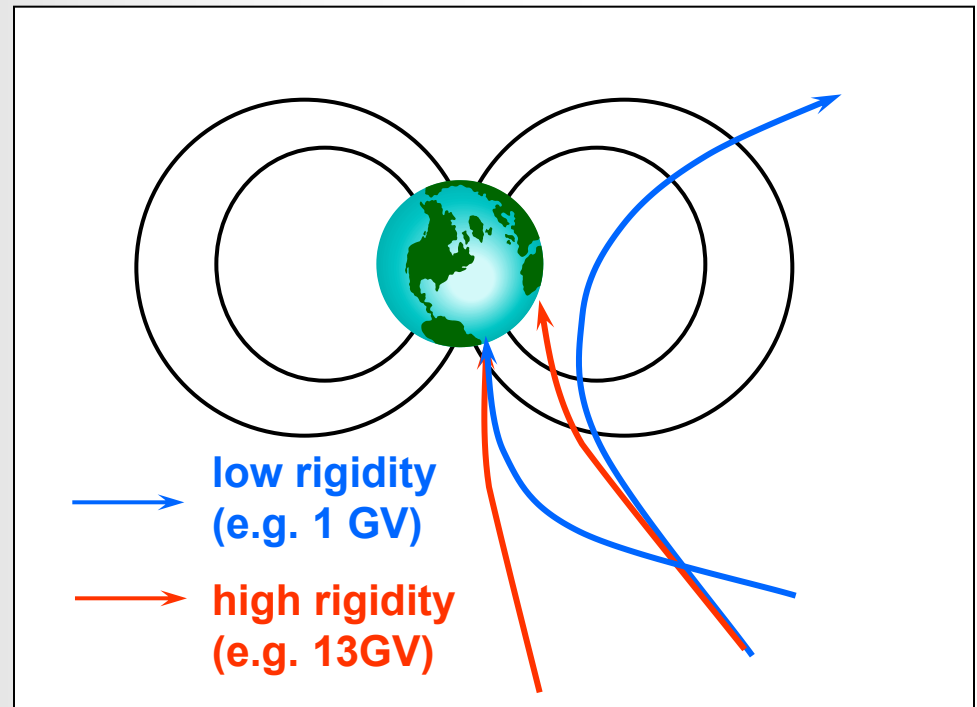




Geomagnetic Shielding of GCRs

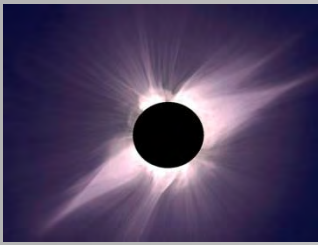
(Cut-off rigidity)

- **Rigidity** is a measure of the extent to which cosmic rays maintain their direction of motion
- It is measured in GV ($v \approx c$, nGV rigidity \equiv energy \approx nGeV)
- Higher rigidity GCRs can penetrate to lower geomagnetic latitudes
- minimum rigidity that can be seen at a magnetic latitude called the “rigidity cut-off” (e.g.) for Hawaii and Huancayo \approx 13GV for Climax (Boulder) \approx 3GV
- At highest latitudes rigidity cut-off set by atmosphere at \approx 1GV

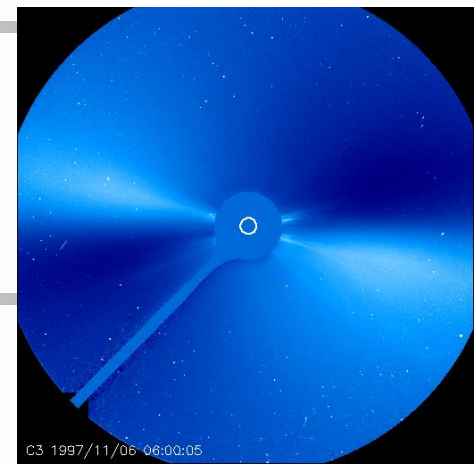


Cosmic ray tracks in a bubble chamber





Solar Output Signals in Troposphere



Visible/IR

at most, very small “bottom up” signals reported in troposphere

UV

clear heating effects in stratosphere (ozone layer) – may have subtle “top down” effects on troposphere

EUV

dominates thermosphere, no evidence nor credible mechanism for coupling to the troposphere

X-Rays

major effects in thermosphere, no evidence or credible mechanism for coupling to the troposphere

Solar wind

same as for EUV and X-rays

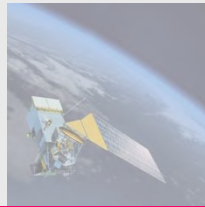
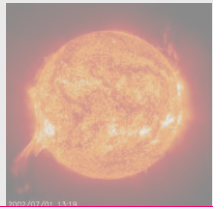
Cosmic Rays

proposed modulation of cloud cover: effect on surface temperatures depends critically on cloud height

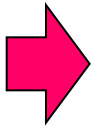
SEPs

destroy ozone so may have similar effects to UV

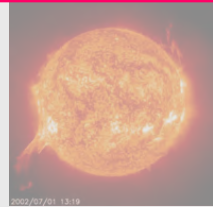
Solar Variability: Effects on Climate?



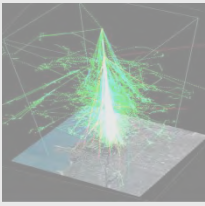
Solar Outputs



Solar Variability



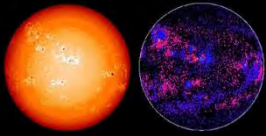
Global Effects



Regional & Seasonal Effects

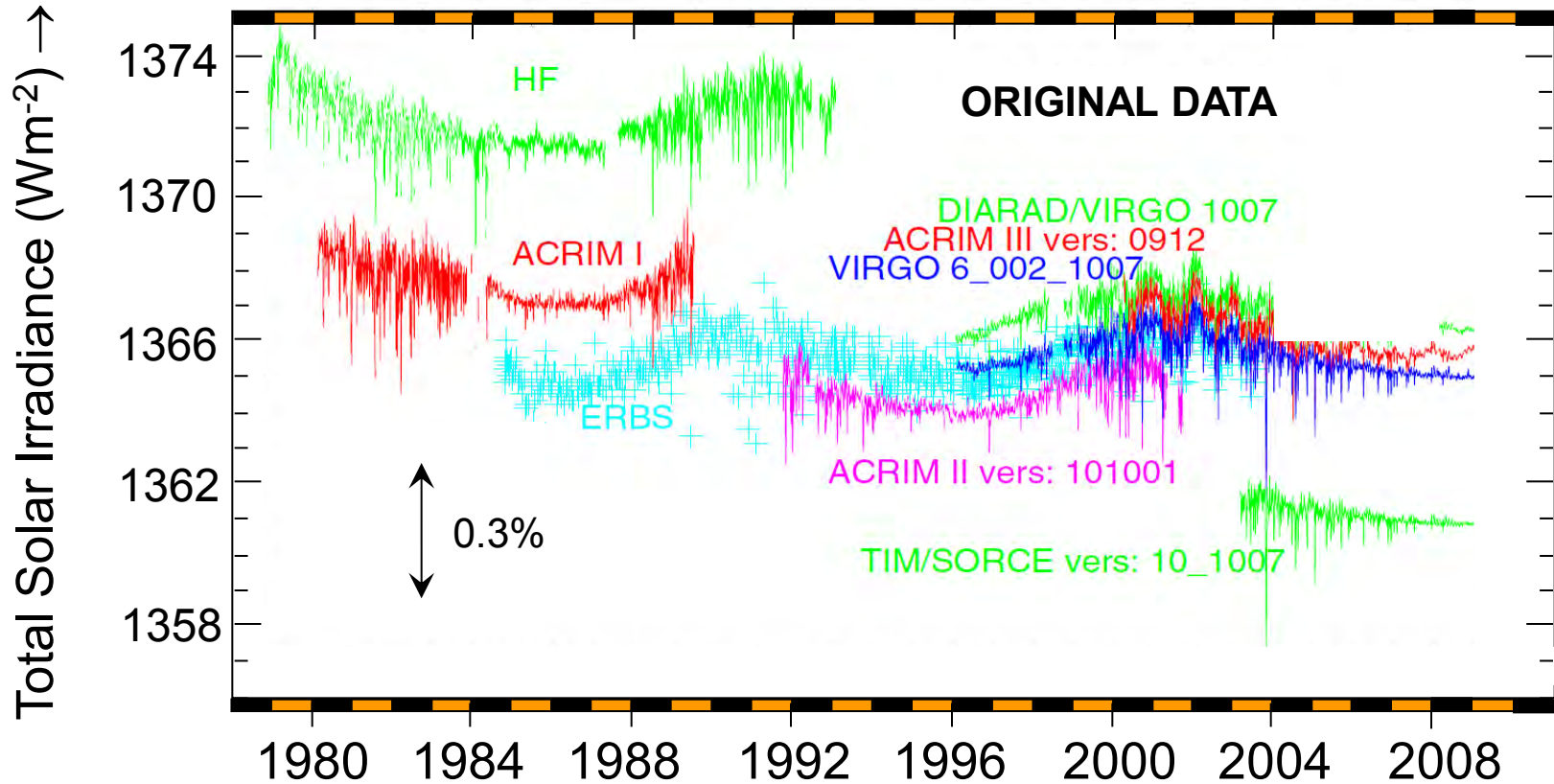


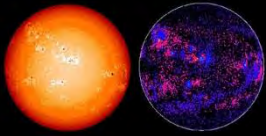
The Future



Total Solar Irradiance Observations

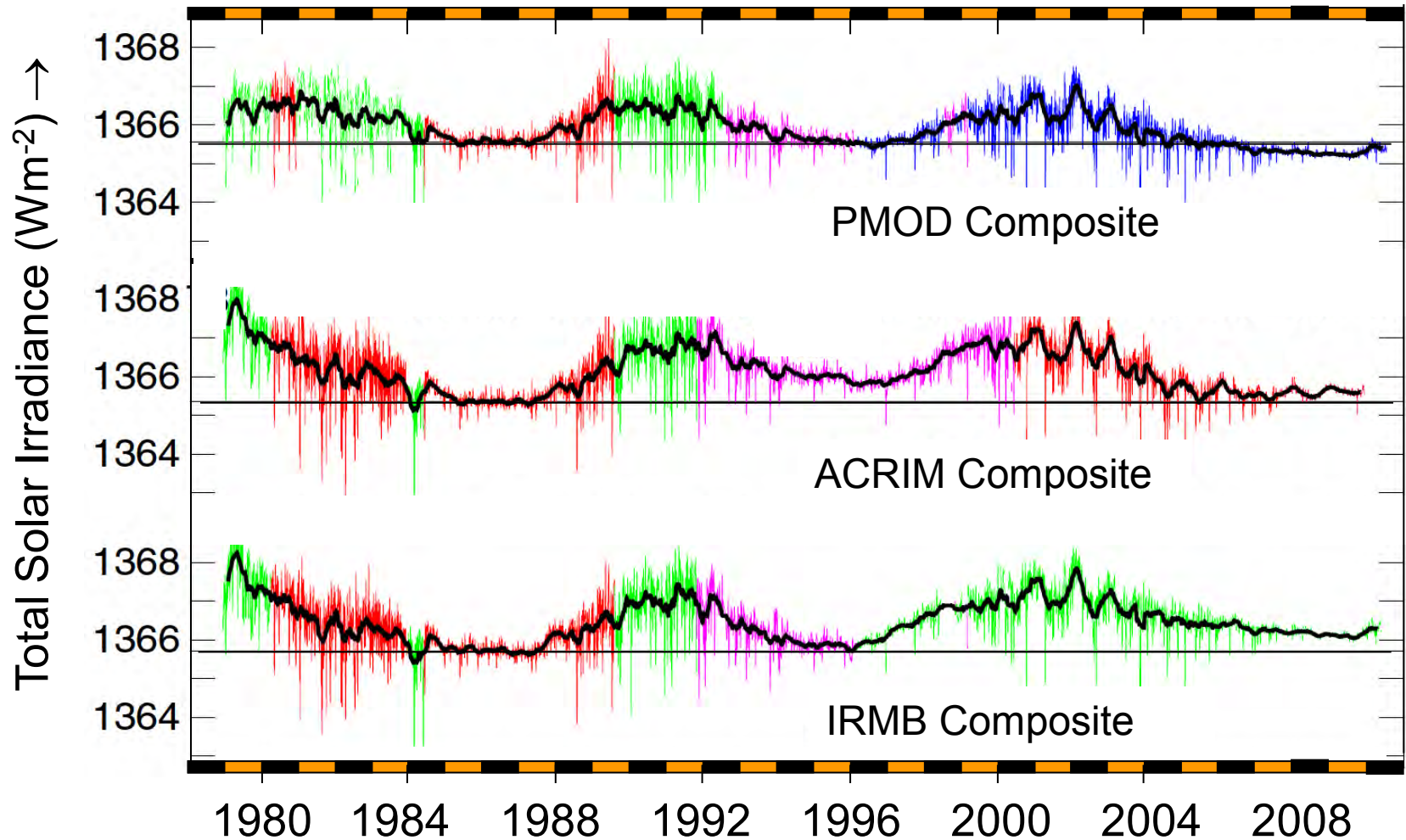
Systematic errors and drifts
due to instrument degradation





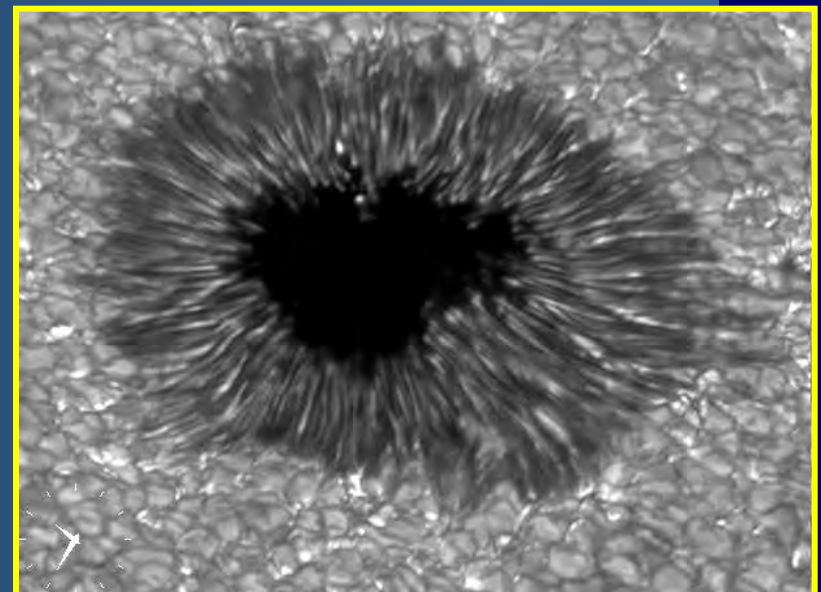
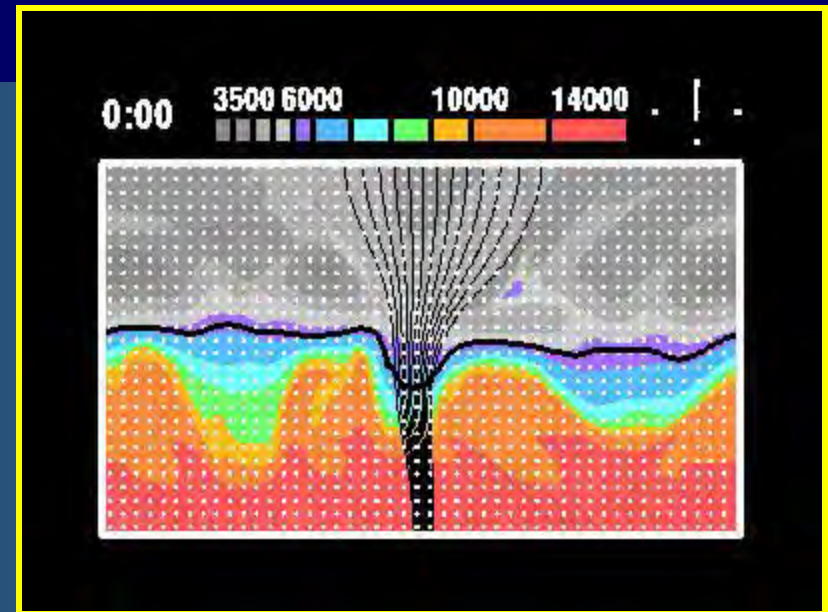
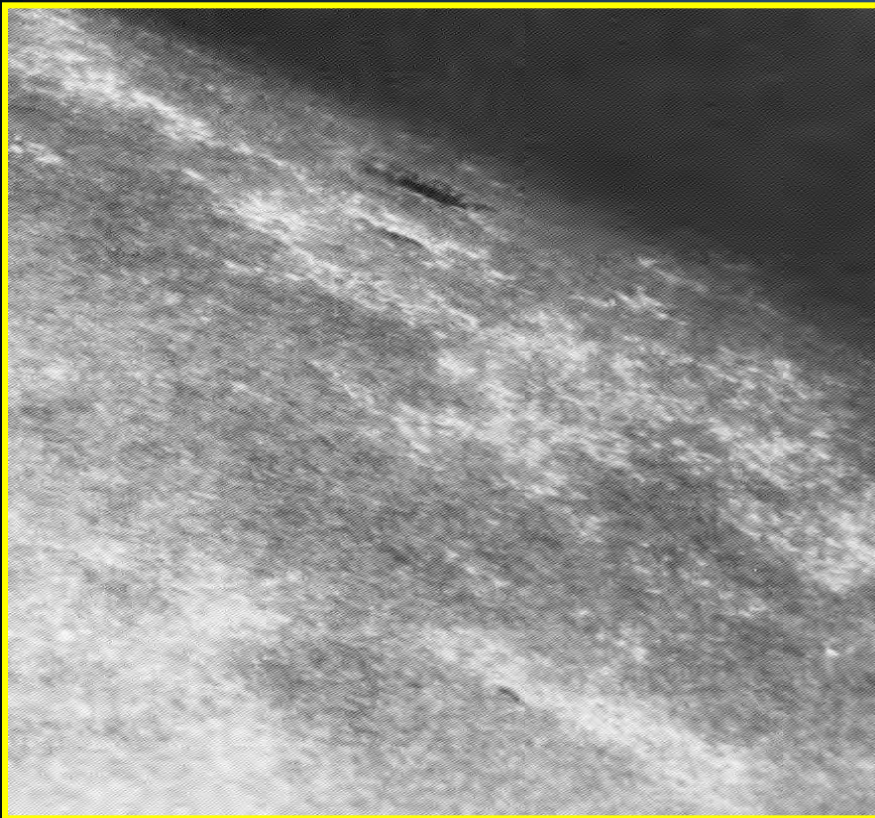
Solar Irradiance Composites

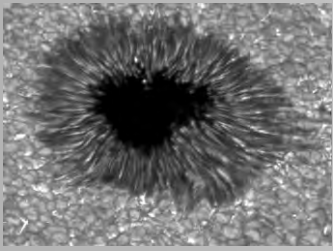
Errors and drifts corrected
by intercalibration



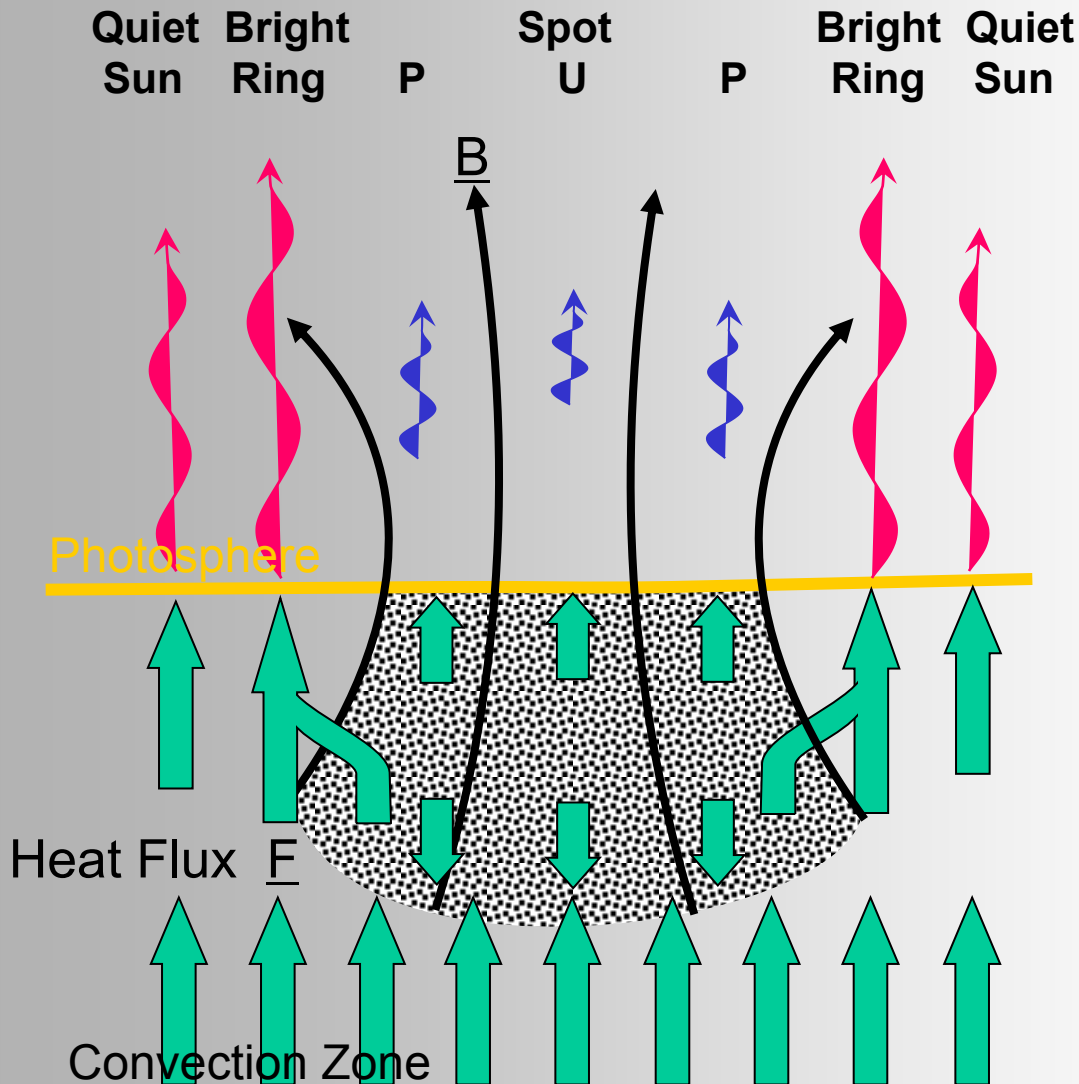
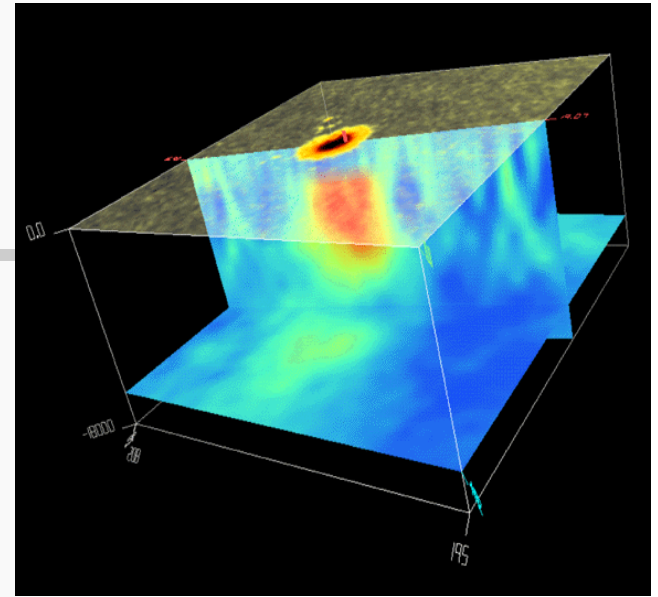
Total solar irradiance changes and magnetic field emergence

- Dark sunspots and bright faculae are where magnetic field threads the solar surface

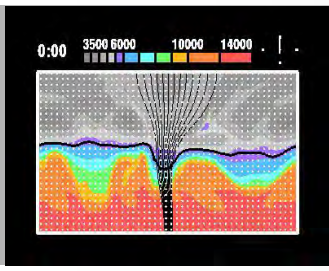




Sunspot Darkening



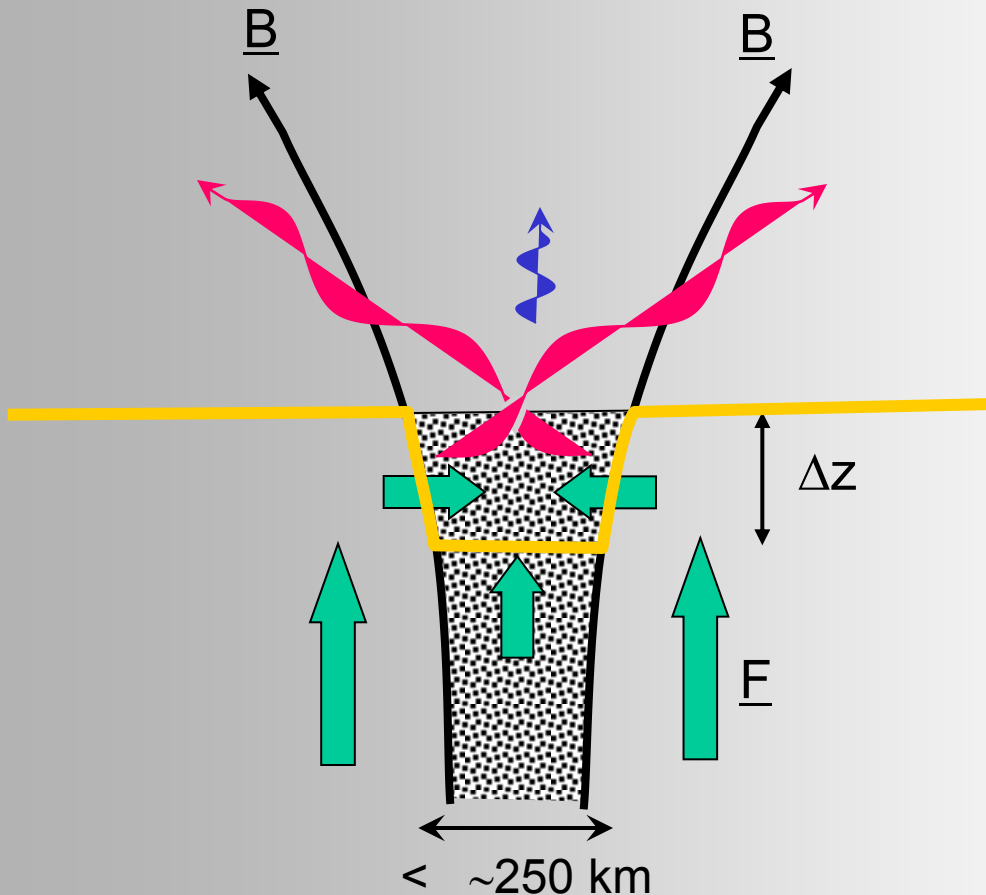
- Enhanced field \underline{B} blocks upward heat flux \underline{F}
- Gives temperatures:
 - Quiet Sun $T_{QS} \approx 6050\text{K}$
 - Bright ring $T_{BR} \approx 6065\text{K}$
 - Penumbra $T_P \approx 5680\text{K}$
 - Umbra $T_U \approx 4240\text{K}$



Facular Brightening

The Bright Wall Model

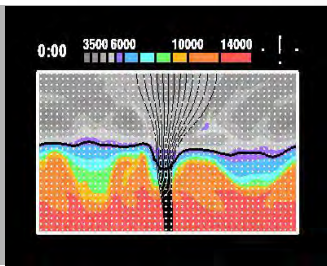
- Enhanced field raises magnetic pressure and depresses thermal pressure $Nk_B T$



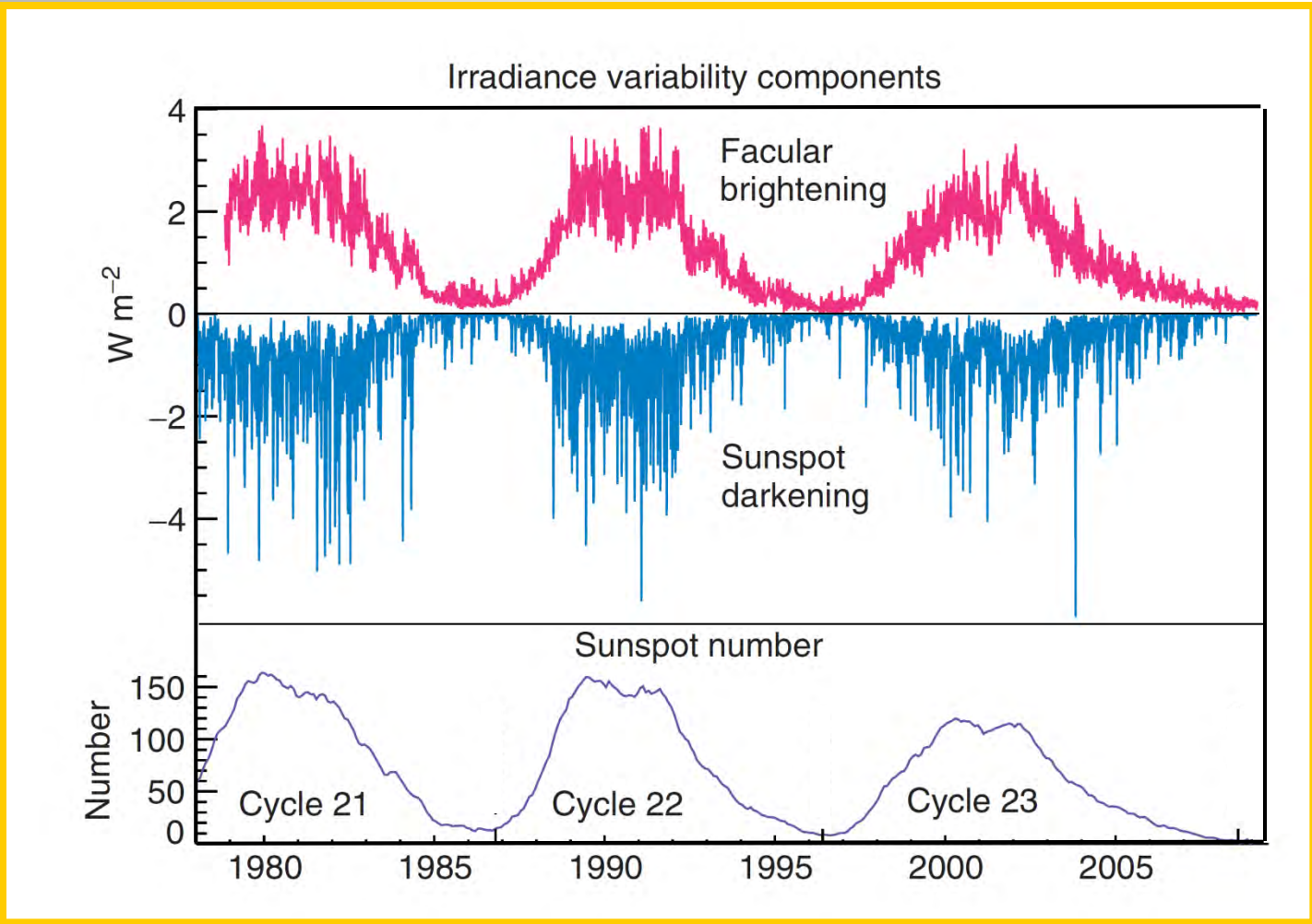
- flux tube small enough for radiation from walls to maintain internal temperature T

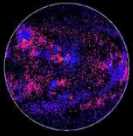
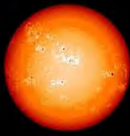
- N falls & the $\tau_{0.5} = 2/3$ contour is depressed by $\Delta z \approx 50 \text{ km}$

- bright walls most visible at small μ for which $T_f \approx 6200 \text{ K}$

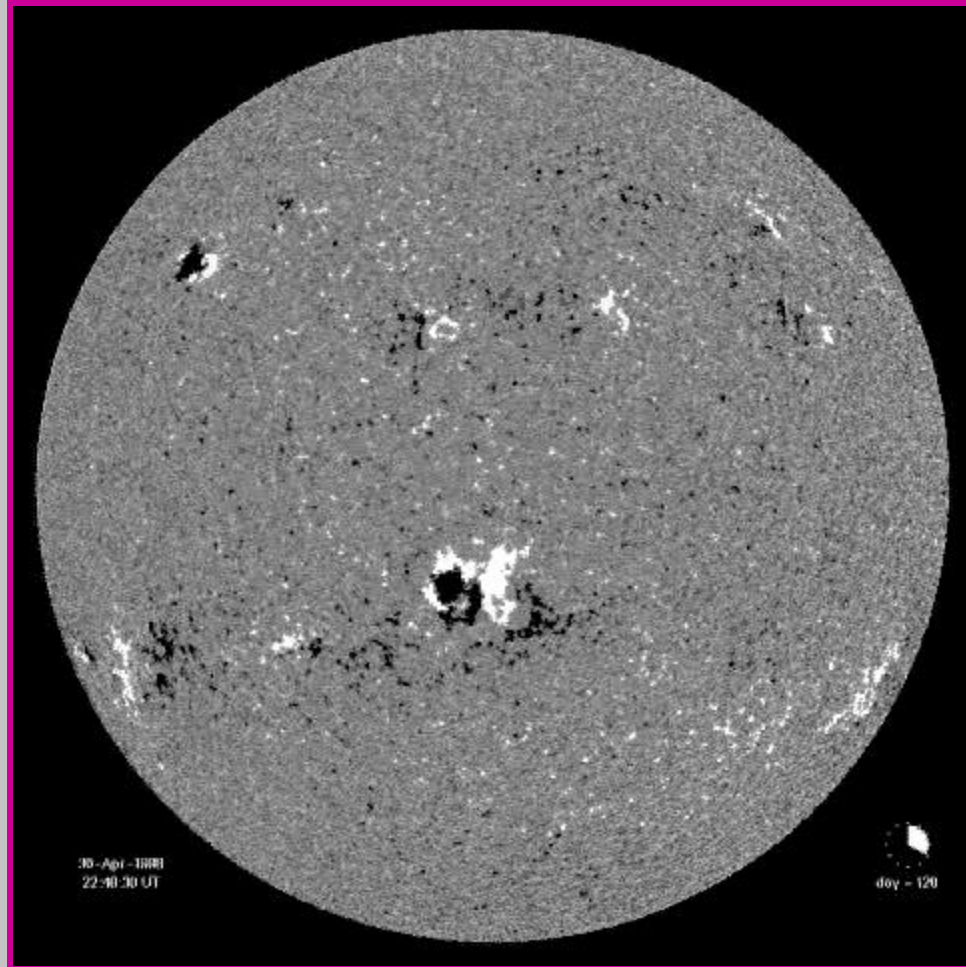


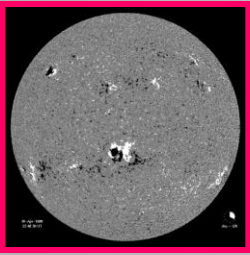
Sunspot Darkening & Facular Brightening





Photospheric magnetic field magnetogram data



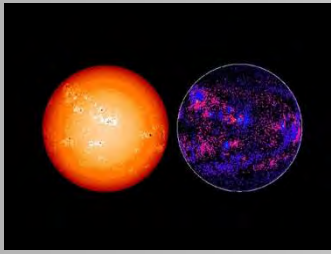


3-component TSI model

using magnetogram data

- Use model contrasts of umbrae, penumbrae and faculae C_U , C_P , and C_F (>0 for brightenings) as a function of position on disc μ and wavelength λ (w.r.t quiet Sun, so $C_{QS}(\mu, \lambda) = 0$)
- Contrasts independent of time t – the time dependence is all due to that in the filling factors α which are functions of μ and t , but not λ .
- Every pixel in the magnetogram for time t that falls on the visible disc is then classified as either umbra, penumbra, facula or quiet Sun to derive α_U , α_P , α_F . Limb darkening function is $L_D(\mu, \lambda)$ and the quiet-Sun intensity (free of all magnetic features) of the disc centre is I_O

$$I_{TS}(\lambda, t) = (\pi R_s^2 / R_1^2) I_O \int_0^1 L_D(\mu, \lambda) \left[\underbrace{\alpha_U(\mu, t)\{C_U(\mu, \lambda)+1\}}_{\text{umbrae}} + \underbrace{\alpha_P(\mu, t)\{C_P(\mu, \lambda)+1\}}_{\text{penumbrae}} + \underbrace{\alpha_F(\mu, t)\{C_F(\mu, \lambda)+1\}}_{\text{faculae}} + \underbrace{\{1-\alpha_P(\mu, t)-\alpha_U(\mu, t)-\alpha_F(\mu, t)\}}_{\text{quiet Sun}} \right] \mu d\mu$$

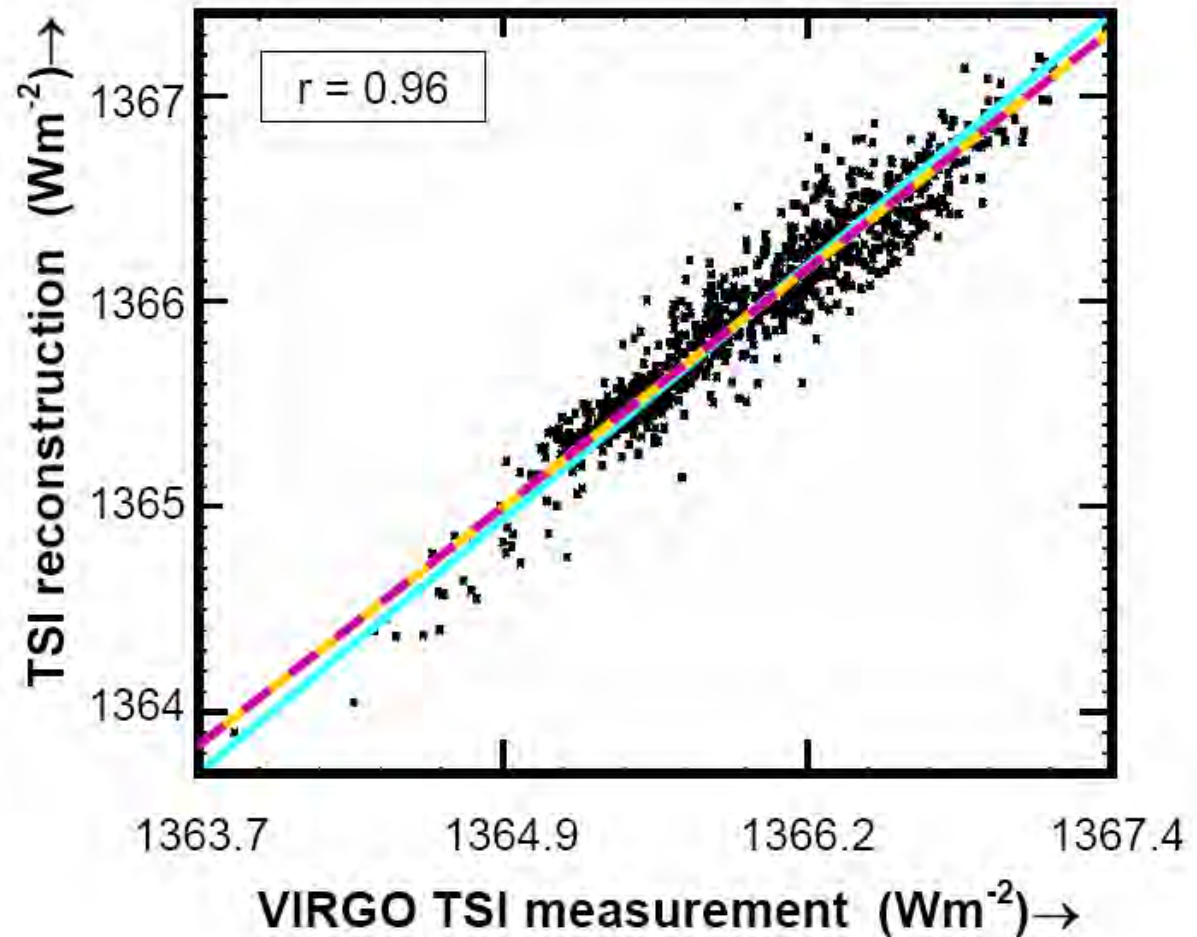


4-component model

(*Solanki et al., 2003*)



- Total Solar Irradiance reconstructions using 4 component model (“SATIRE”) with magnetograms for 1996-2002 from the MDI satellite, compared with SoHO TSI data

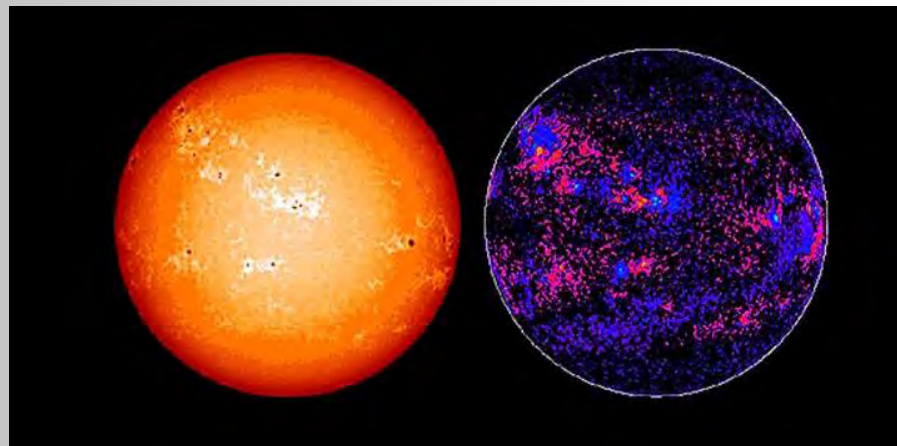




Stellar Analogues: The use of the S index



- ▶ S index is a measure of stellar flux in the Ca I H and K lines (chromospheric emissions associated with magnetic field threading the solar surface)
- ▶ related to facular brightening term in TSI by Lean et al. (1992)





Stellar Analogues:

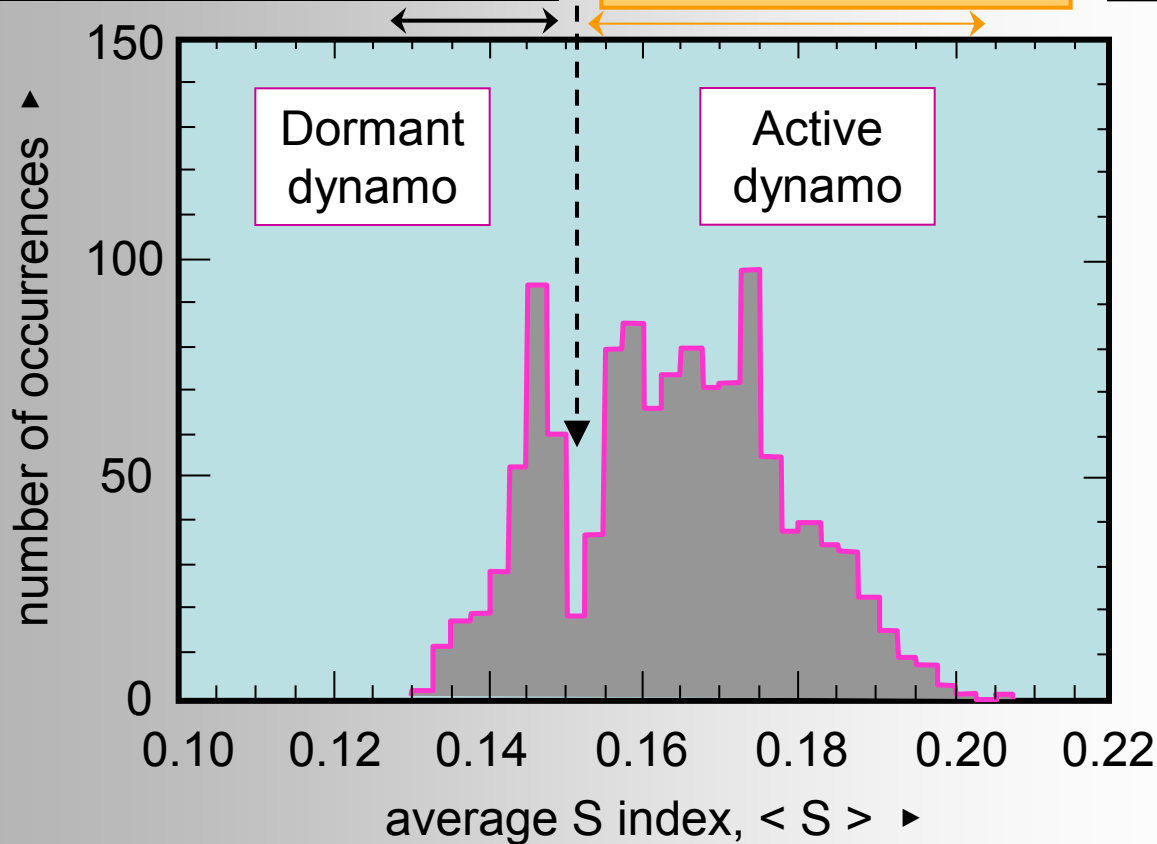
The distribution of S index values



1/3 non-cyclic stars
≡ Sun's Maunder minimum?

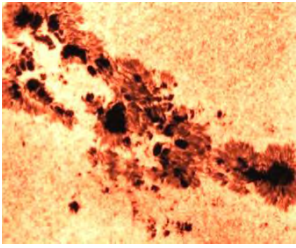
2/3 cyclic stars
≡ present-day Sun?

← from 13 of the 74
(so a third is just 4)

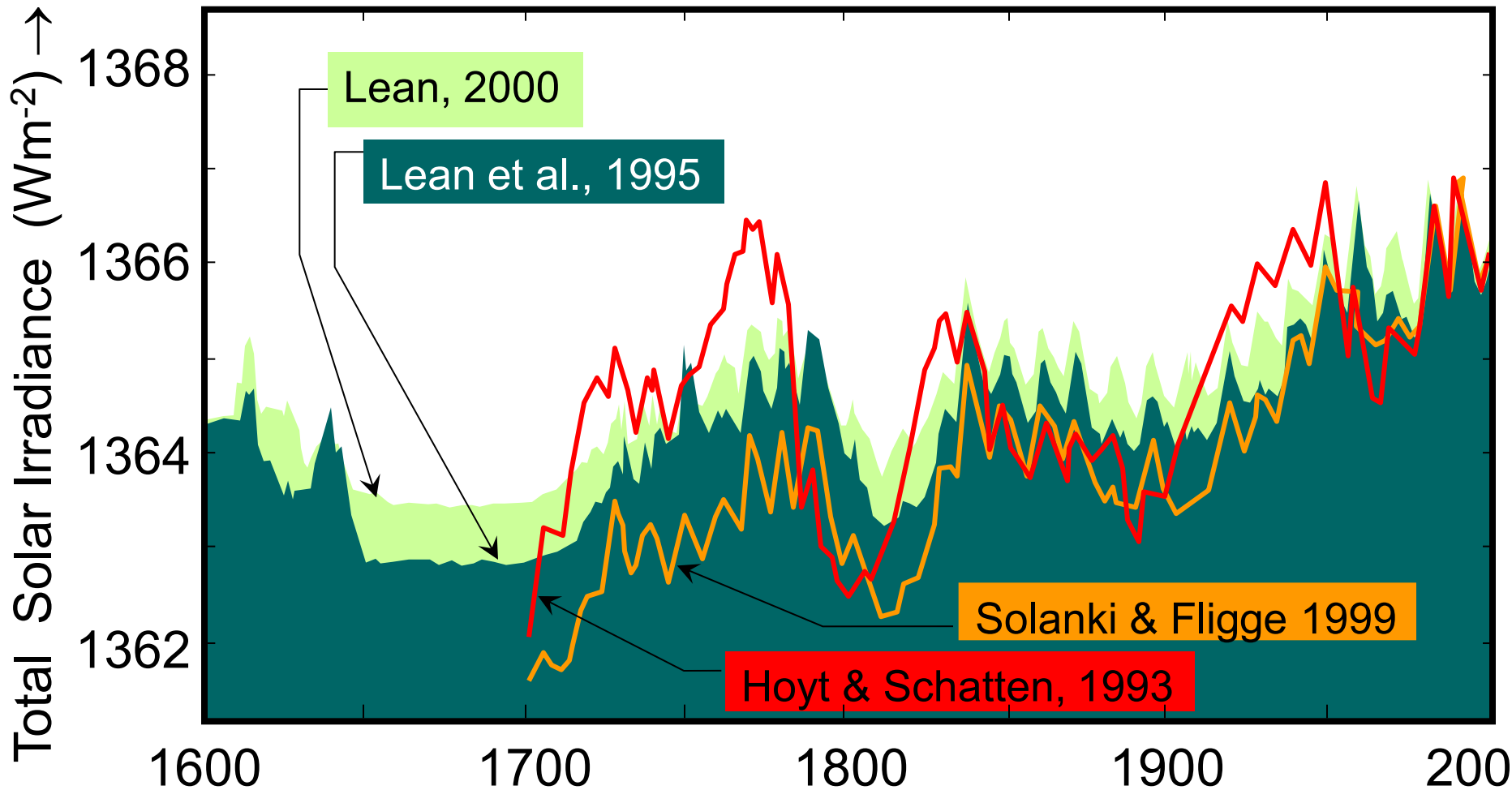


► Baliunas & Jastrow (1990). Data from the Mt. Wilson survey of Sun-like stars.

► 74 “solar-type” stars with B - V colours in range 0.60–0.76 (0.95–1.10 M_{\odot}).



TSI Reconstructions



Hoyt and Schatten used solar cycle length, L , Lean et al. and Lean used a combination of sunspot number R and R_{11} , Solanki and Fligge use a combination of R and L , Lockwood and



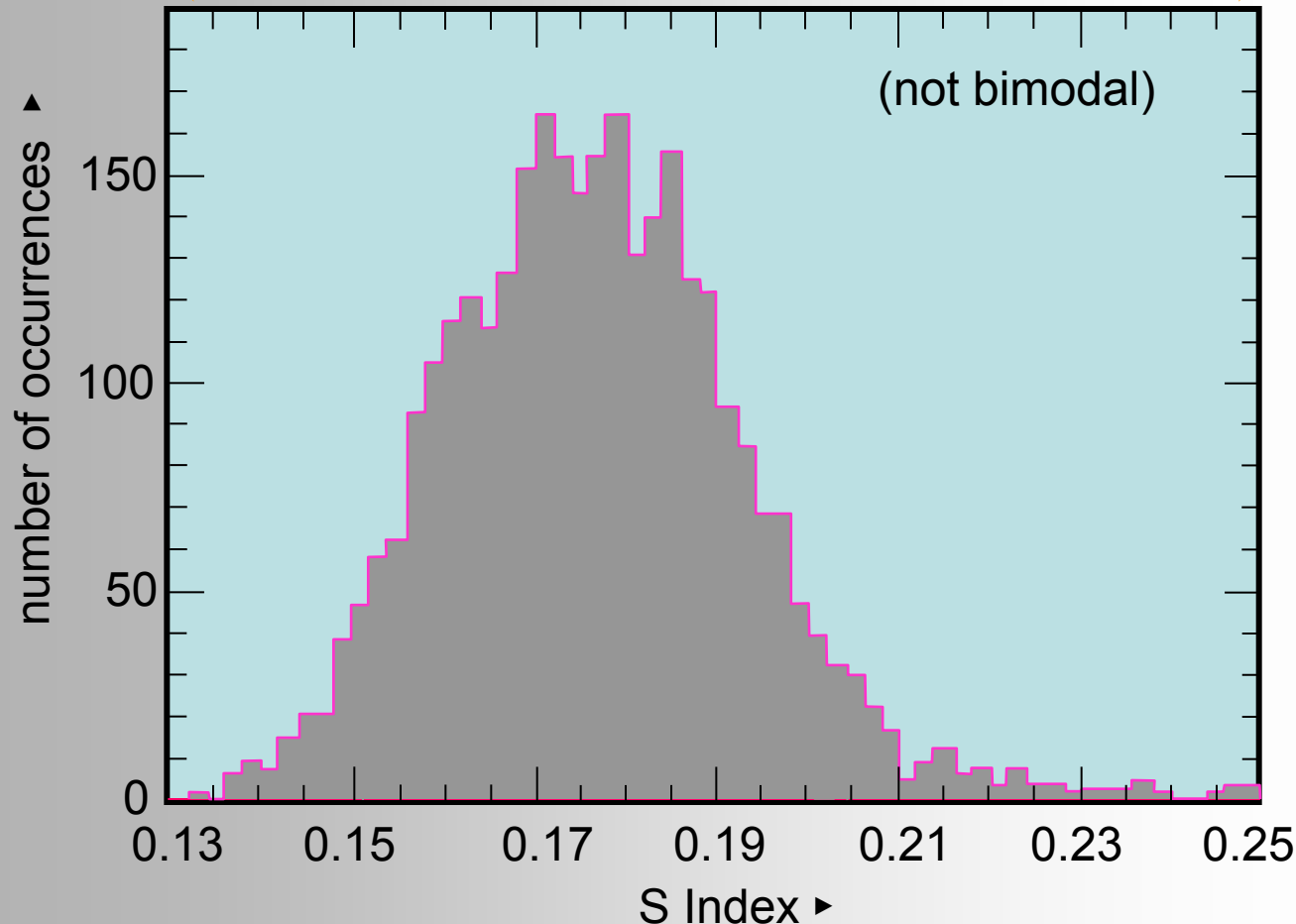
Stellar Analogues:

Recent re-evaluation of distribution



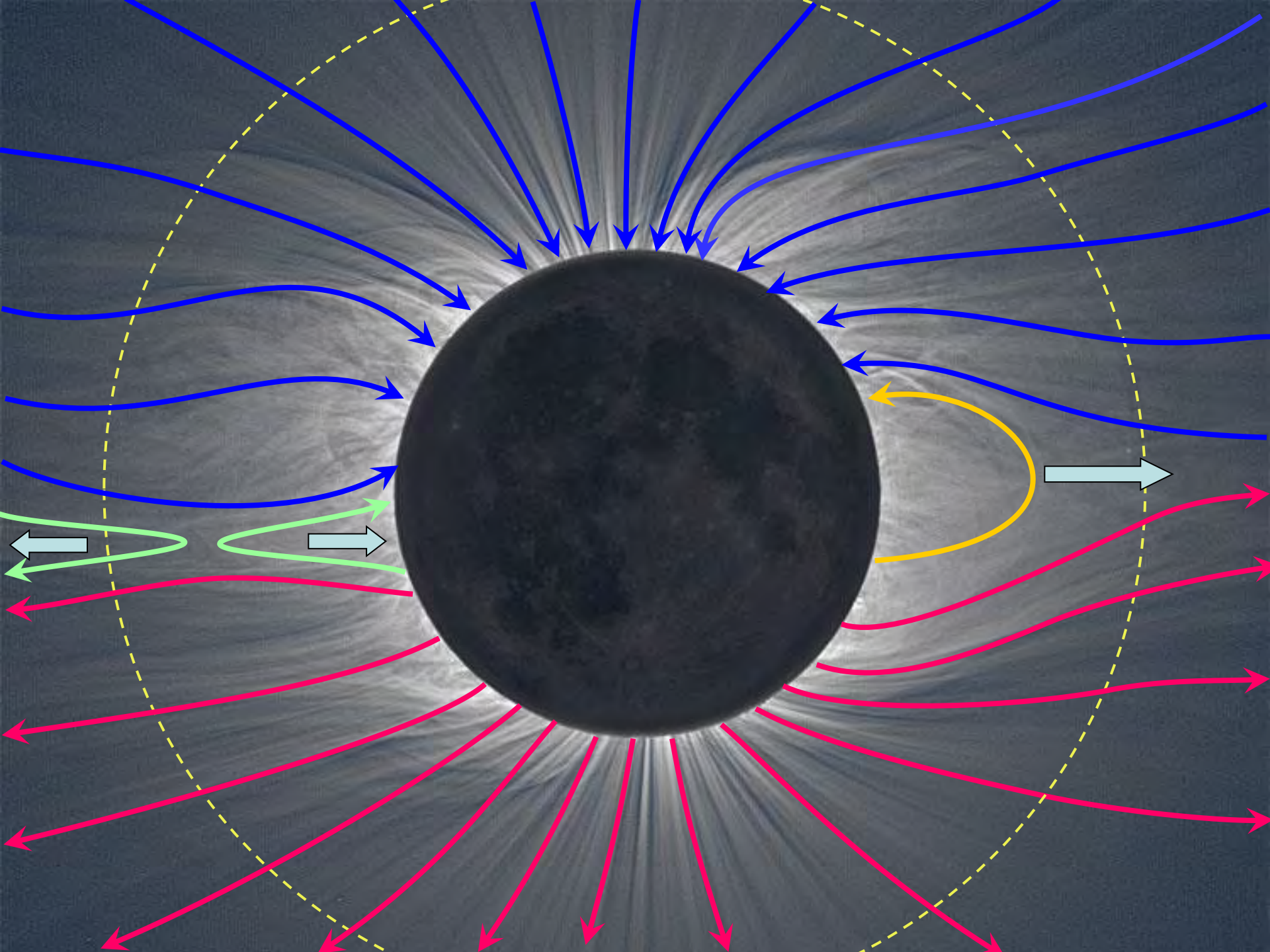
1/3 non-cyclic stars

2/3 cyclic stars



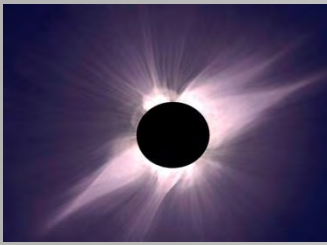
► Hall and Lockwood (2004) Lowell survey of 300 stars with colours in the same range as adopted by B&J ($0.60 \leq B-V \leq 0.76$)





Analogy: the spacing of birds on a wire!



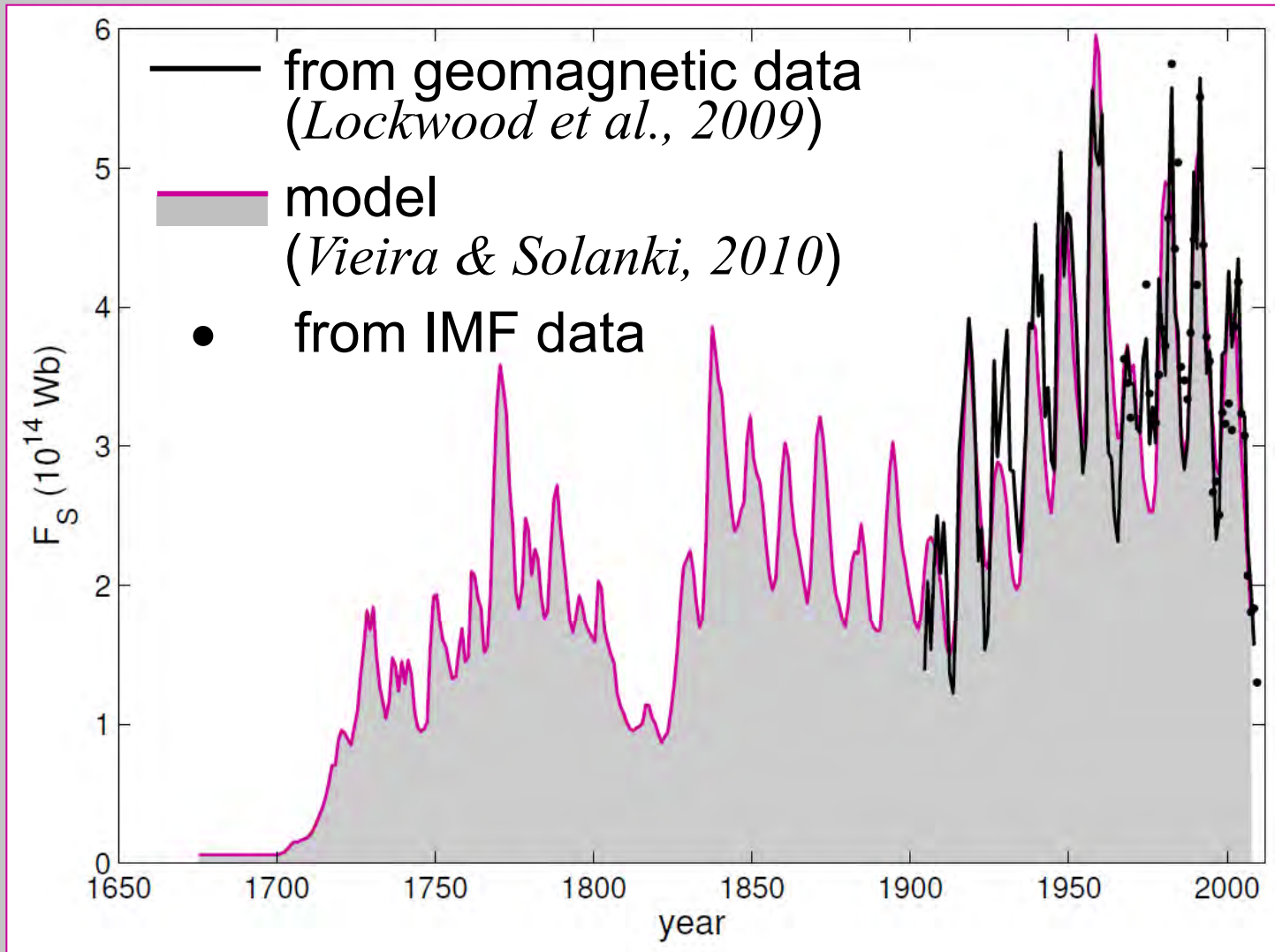


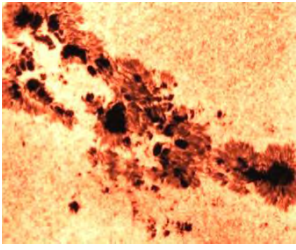
Open Solar Flux, F_S

(allowing for longitudinal structure in solar wind)

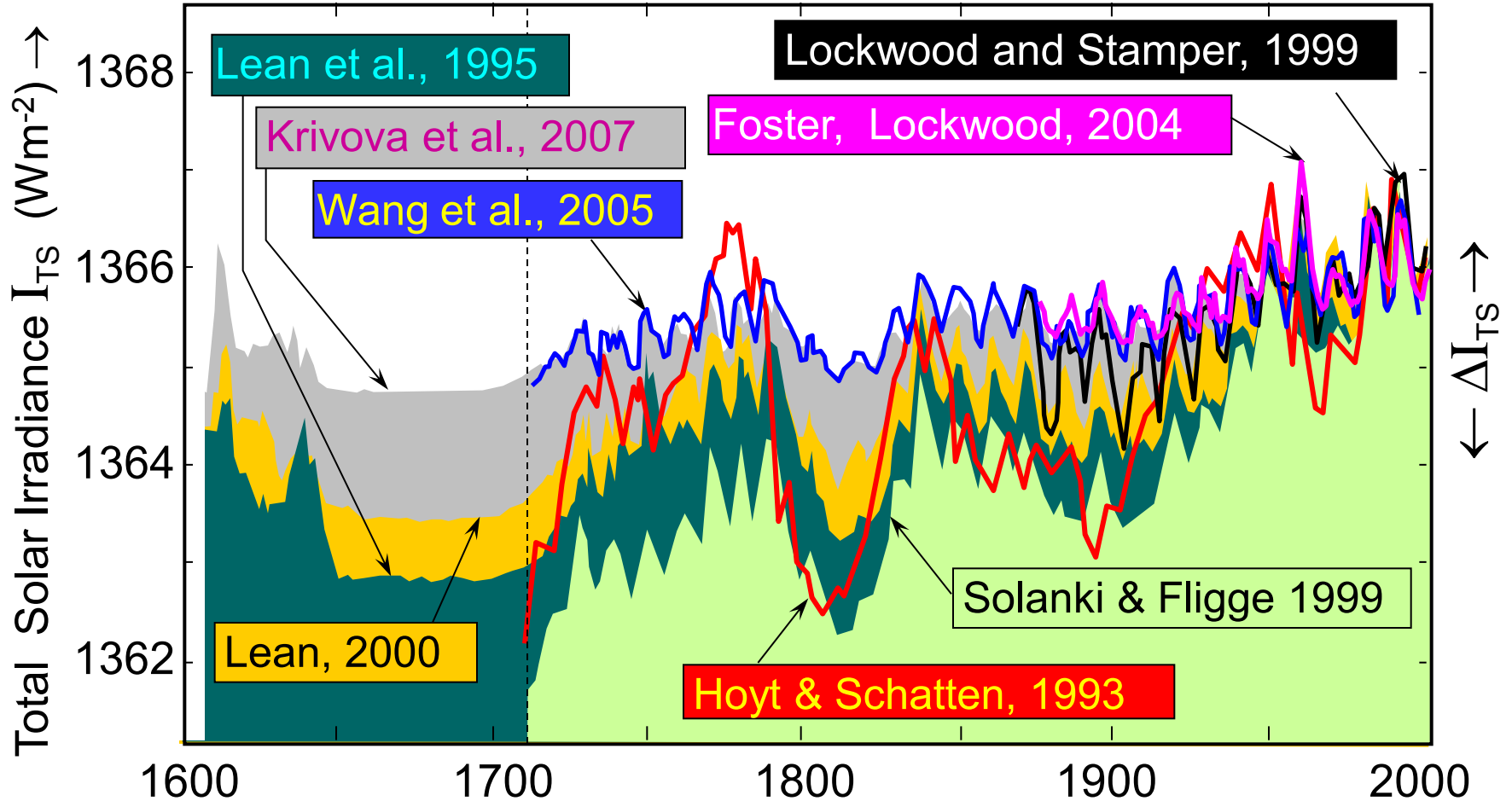
► use both range and hourly mean geomagnetic data

► model emergence from sunspot number with two time constants for decay of open flux





TSI Reconstructions



Most recent best estimates are $\Delta I_{TS} \approx 1 Wm^{-2}$ since MM

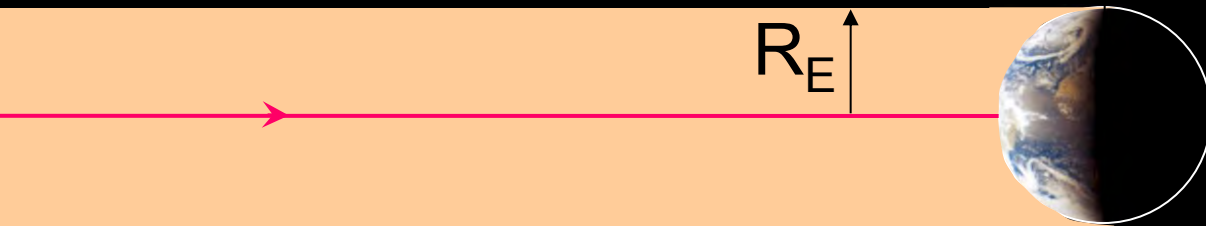
Outgoing Longwave (LW) Radiation

- ▶ infra-red (Longwave, LW) emission = heat
- ▶ Earth is close to a “Blackbody” radiator of effective temperature T_E
- ▶ emitted power by unit area of Earth = σT_E^4 where σ is the Stefan – Boltzmann constant
- ▶ **Define** $T_E^4 = (1-g)T_S^4$, where g is the greenhouse term
- ▶ surface area of $4\pi R_E^2$, so total LW power output,



$$P_{\text{out}} = 4\pi R_E^2 \sigma T_E^4 = 4\pi R_E^2 \sigma (1 - g)T_S^4$$

Incoming short wave (SW) radiation



- ▶ power density in sunlight = I_{TS} ($W m^{-2}$)
- ▶ called the “total solar irradiance” (TSI)
- ▶ the area of target presented by Earth = πR_E^2 (m^2)
where R_E is the mean Earth radius
- ▶ of the incident power a fraction A is reflected back into space, where A is called Earth’s “albedo”
- ▶ of the incident power a fraction $(1-A)$ is not reflected back into space,
- ▶ Input SW Power $P_{in} = I_{TS} \pi R_E^2 (1 - A)$

Terrestrial Energy Budget



$$\text{Input SW Power } P_{\text{in}} = I_{\text{TS}} \pi R_{\text{E}}^2 (1 - A)$$

$$\text{Output LW Power } P_{\text{out}} = 4\pi R_{\text{E}}^2 \sigma T_{\text{E}}^4 = 4\pi R_{\text{E}}^2 \sigma (1 - g)T_{\text{S}}^4$$

σ = Stefan-Boltzmann constant

T_{E} = effective temperature of Earth / atmosphere $\approx 255\text{K}$

T_{S} = surface temperature of Earth

g = normalised greenhouse effect

Also need to consider power q (per unit area) surface gives to sub-surface layers (particularly the deep oceans)

$$P_{\text{in}} = P_{\text{out}} + 4\pi R_{\text{E}}^2 q$$

→

$$I_{\text{TS}}(1 - A)/4 = \sigma (1 - g)T_{\text{S}}^4 + q$$

Terrestrial Energy Budget



$$I_{TS}(1 - A)/4 - \sigma T_S^4 + \sigma g T_S^4 - q = 0$$

$$I_{TS}(1 - A)/4 - \sigma T_S^4 + G - q = 0$$

Differentiate w.r.t. time

$$\Delta T_S = [\Delta I_{TS}/4 - I_{TS}\Delta A/4 + \Delta G - \Delta q] / (4\sigma T_S^3)$$

σ = Stefan-Boltzmann constant

T_E = effective temperature of Earth / atmosphere $\approx 255\text{K}$

T_S = surface temperature of Earth

g = normalised greenhouse effect, N.B., $g = G / (\sigma T_S^4)$

G = greenhouse radiative forcing (in Wm^{-2})

Gives the concept of “radiative forcing” where we can add together the changes in the powers per unit surface area due to different effects in the term in square brackets

A little greenhouse gas is a good thing!



$$T_S = \left[I_{TS} \frac{(1-A) - 4q}{4\sigma(1-g)} \right]^{1/4}$$

If no greenhouse gases, $g = 0$ and surface in equilibrium with oceans ($q = 0$):

$$I_{TS} = 1366.5 \text{ Wm}^{-2}, \quad \text{Albedo, } A = 1/3$$

$$\sigma = 5.669 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$$

$$(\text{if } g = 0 \quad T_E = T_S)$$

$$\text{Gives } T_S = 251.8 \text{ K} = -21.2 \text{ }^\circ\text{C}$$

TOO COLD FOR ALMOST ALL LIFEFORMS!

Terrestrial Energy Budget



$$T_S = \left[\frac{I_{TS} (1 - A) - 4q}{4\sigma (1 - g)} \right]^{1/4}$$

Typical values

$I_{TS} = 1366.5 \text{ Wm}^{-2}$, Albedo, $A = 1/3$,

$q = 1 \text{ Wm}^{-2}$ (*Hansen et al., Science, 2005*)

$\sigma = 5.669 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$

T_E = effective temperature of Earth & its atmos. $\approx 253\text{K}$

$g = 1 - (T_E / T_S)^4$

Above eqn. for $g = 0.410$ gives $T_S = 286.9 \text{ K} = 13.9 \text{ }^\circ\text{C}$

Increase g to 0.416 (a 1.5% rise & \approx the value or 2000)
gives $T_S = 14.7 \text{ }^\circ\text{C}$ i.e. it gives a rise in T_S of $\Delta T_S = 0.8 \text{ }^\circ\text{C}$

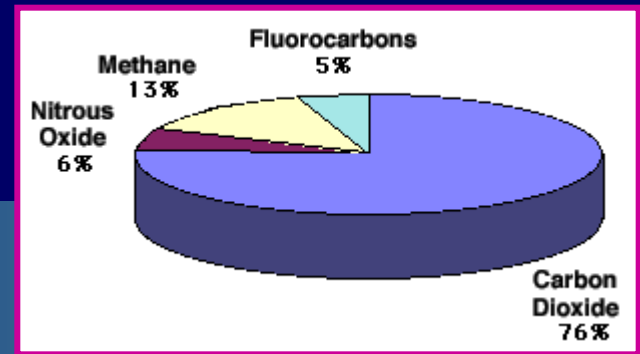
Terrestrial Energy Budget



Typical values from before:

- $g = 0.410$ gives $T_S = 286.9 \text{ K}$ ($= 13.2 \text{ }^\circ\text{C}$)
corresponds to $G = g \sigma T_S^4 = 157.5 \text{ Wm}^{-2}$
- Increasing g to 0.416 (\approx value for 2000) gives
 $T_S = 287.7 \text{ K}$ ($= 14.7 \text{ }^\circ\text{C}$)
(the observed rise in T_S , $\Delta T_S = 0.8 \text{ }^\circ\text{C}$)
corresponds to $G = g \sigma T_S^4 = 161.6 \text{ Wm}^{-2}$
- Thus a radiative forcing anomaly of $\Delta G = 4.1 \text{ Wm}^{-2}$
gives a surface temperature rise $\Delta T_S = 0.8 \text{ K}$
- The “climate sensitivity” = $\Delta T_S / \Delta G$
 $\approx 0.2 \text{ K W}^{-1} \text{ m}^2$

Do greenhouse gases alone explain the observed warming?



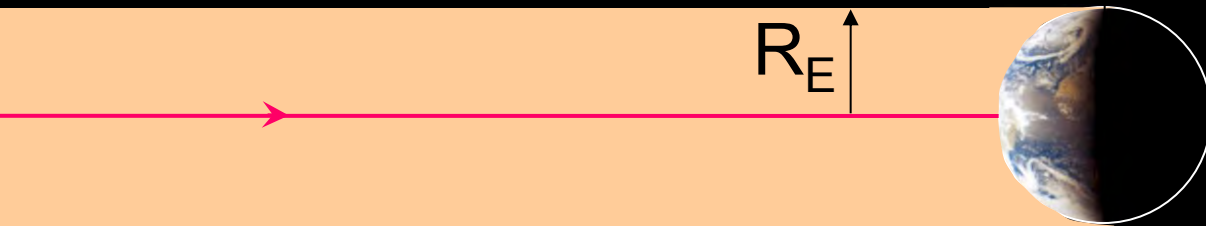
(from before) for 1900-2000:

● radiative forcing anomaly of $\Delta G = 4 \text{ Wm}^{-2}$ gives a surface temperature rise $\Delta T_s = 0.8 \text{ K}$

	ppmv			$\delta G \text{ (Wm}^{-2}\text{)}$		$\Delta G \text{ (Wm}^{-2}\text{)}$
	1700	1900	2000	1900	2000	1900-2000
CO ₂	278	295.2	362.5	0.3	1.4	1.1
NH ₄	700	898	1800	0.2	0.5	0.3
Others				0.1	0.5	0.4
Total				0.6	2.4	1.8

● direct effects not enough: but there are **feedback effects**

Solar radiative forcing



▶ Input SW Power $P_{in} = I_{TS} \pi R_E^2 (1 - A)$

▶ SW Power per unit surface area of earth

$$P_{SW} = I_{TS} \pi R_E^2 (1 - A) / (4 \pi R_E^2) = I_{TS} (1 - A) / 4$$

Solar radiative forcing = $\Delta P_{SW} = \Delta I_{TS} (1 - A) / 4$

Since pre-industrial times $\Delta I_{TS} \approx 1 \text{ Wm}^{-2}$

Gives $\Delta P_{SW} \approx 1/6 = 0.167 \text{ Wm}^{-2}$ (for $A = 1/3$)

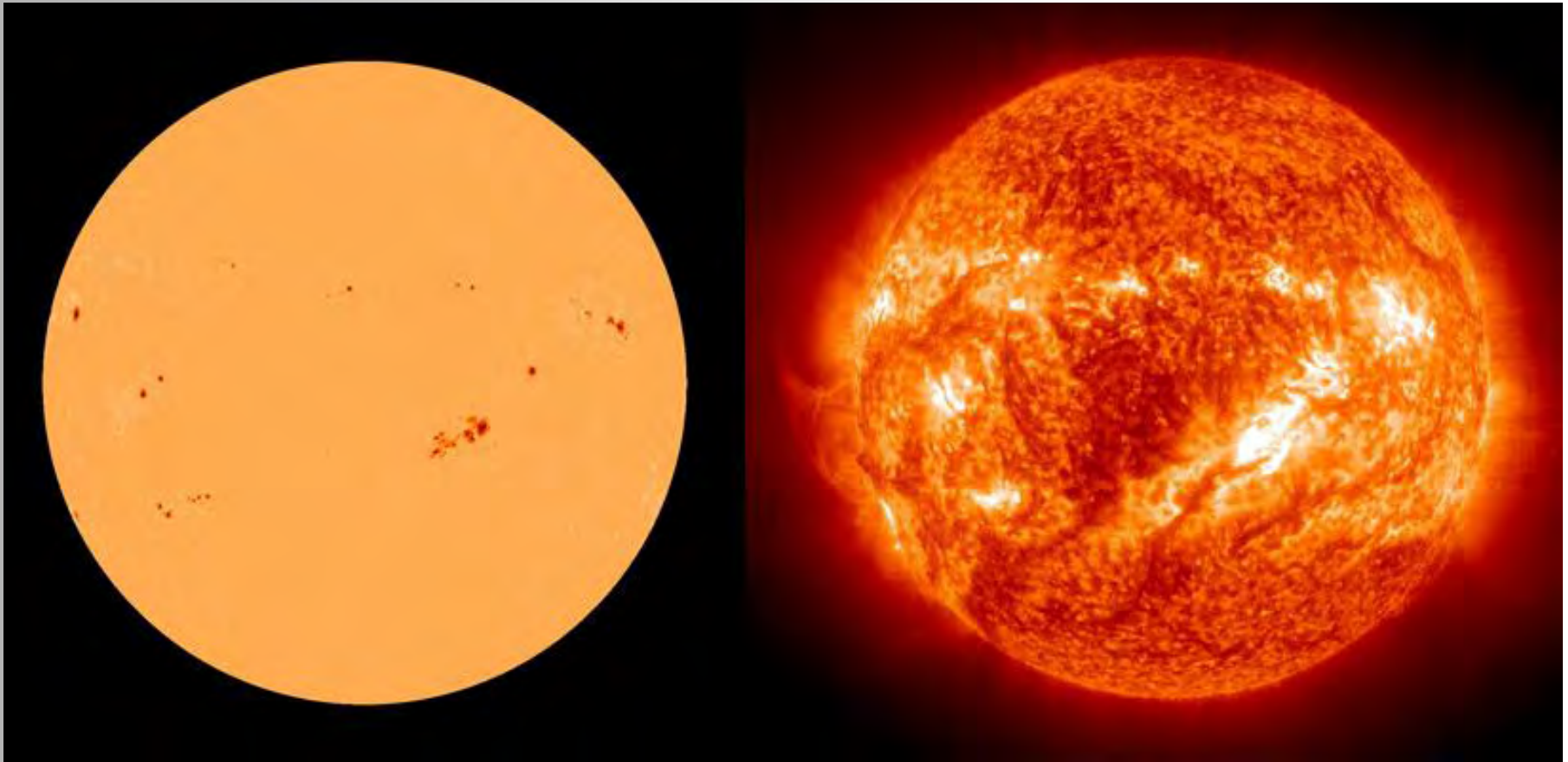
= **a tenth** of greenhouse gas radiative forcing $\approx 1.8 \text{ Wm}^{-2}$

And remember total radiative forcing needed to explain
GMAST rise (with feedbacks) = $\Delta G \approx 4 \text{ Wm}^{-2} \approx 24 \Delta P_{SW}$



The sun seen is Visible and UV light

3rd February 2002

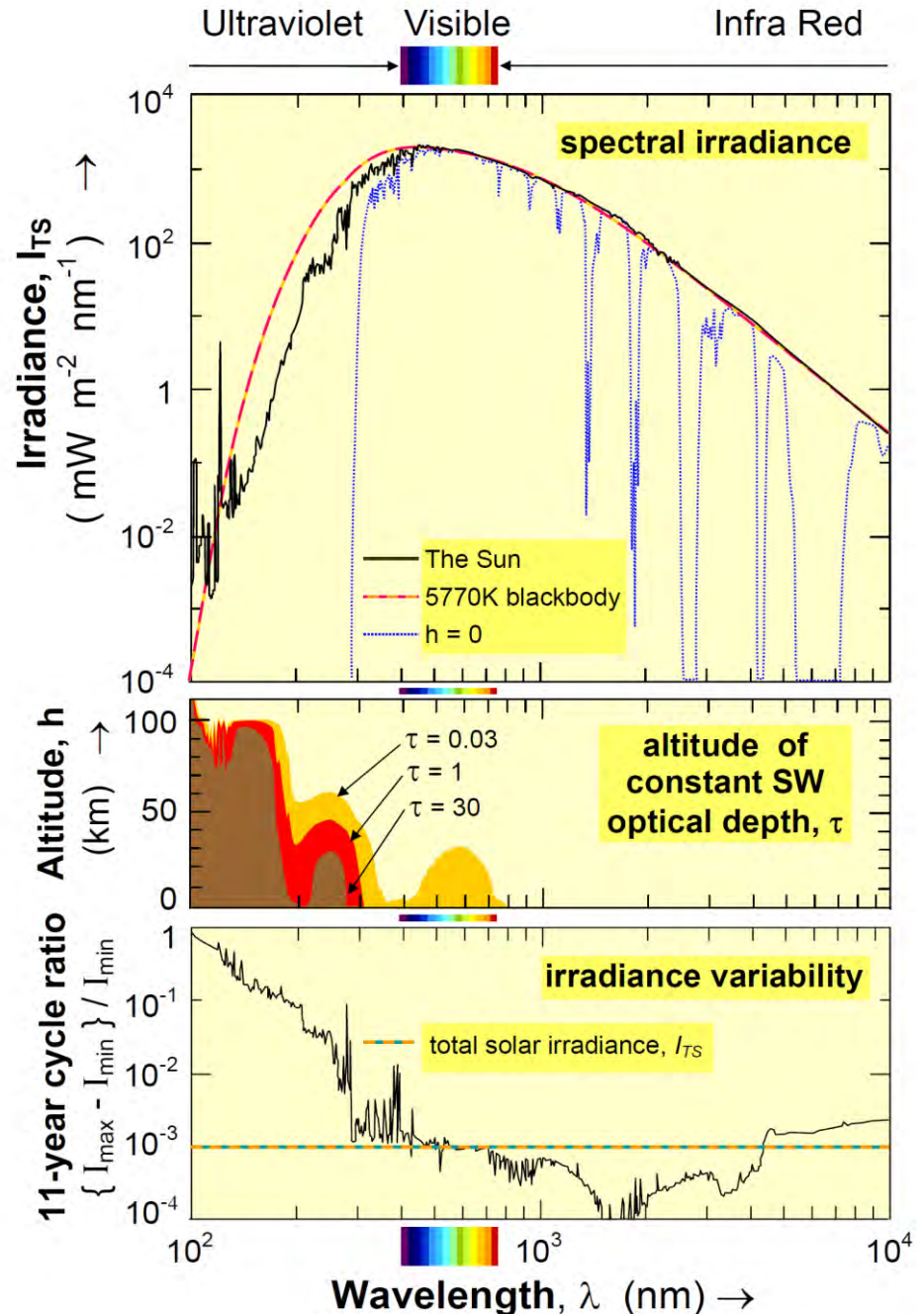


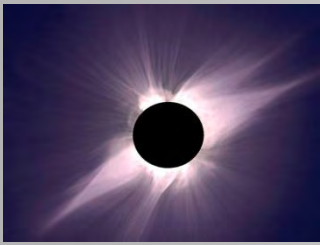
▶ white light, $\lambda = 400-700$ nm

▶ Ultraviolet, $\lambda = 30.4$ nm



- The Sun's e-m radiation spectrum
- Variability is low in parts of spectrum power is greatest
- Variability is highest in UV which is absorbed in the stratosphere



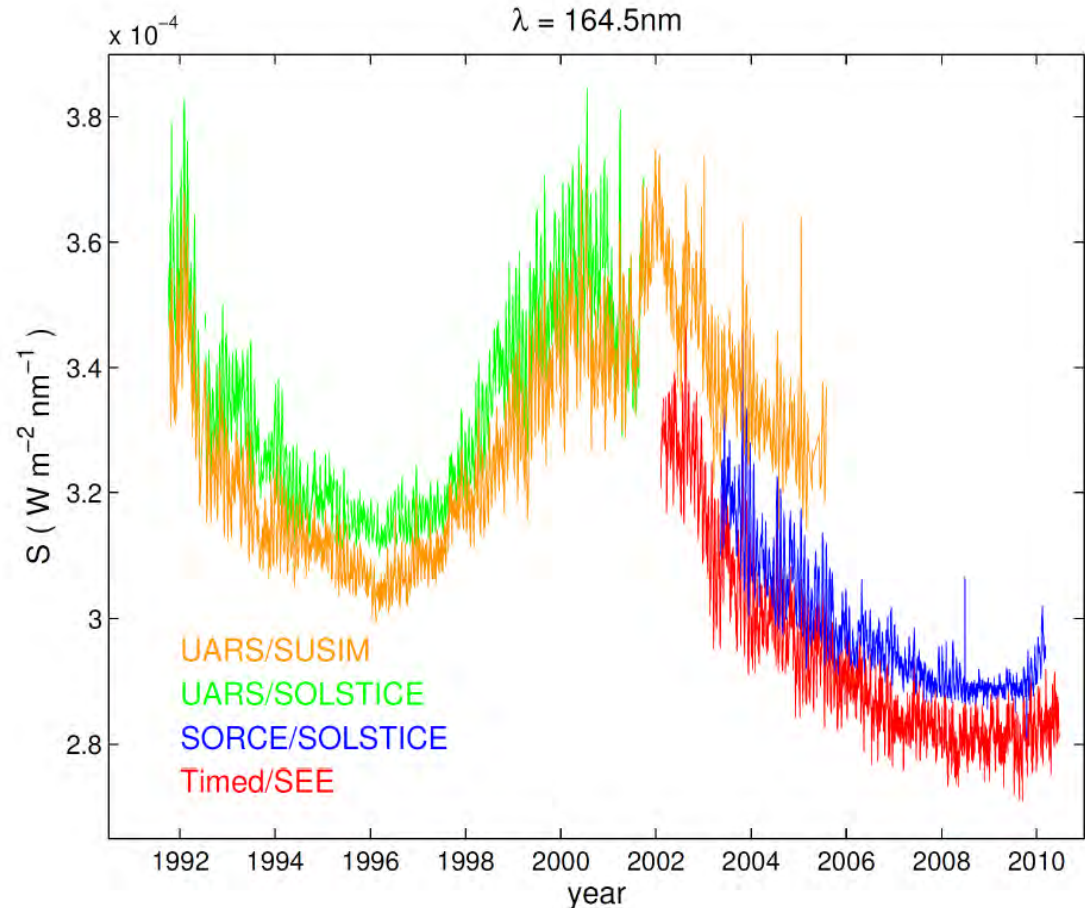


Solar UV data intercalibration

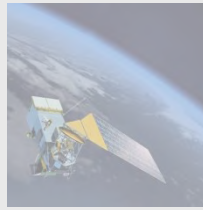
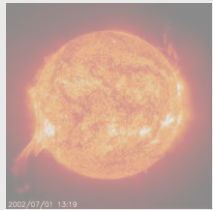
(Lockwood, JGR, 2011)



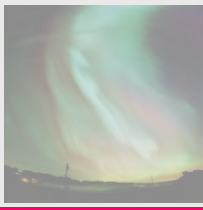
- ▶ e.g. $\lambda = 164.5\text{nm}$
- ▶ data from different satellite and instruments
- ▶ note the “SOLSTICE gap” between the end of UARS/SOLSTICE data and start of SORCE/SOLSTICE data.
- ▶ UARS/SUSIM data not reliable at $\lambda > 205\text{nm}$



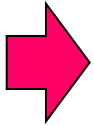
Solar Variability: Effects on Climate?



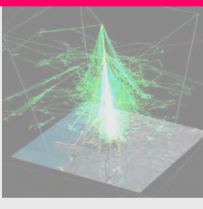
Solar Outputs



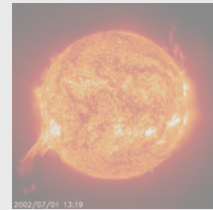
Solar Variability



Global Effects

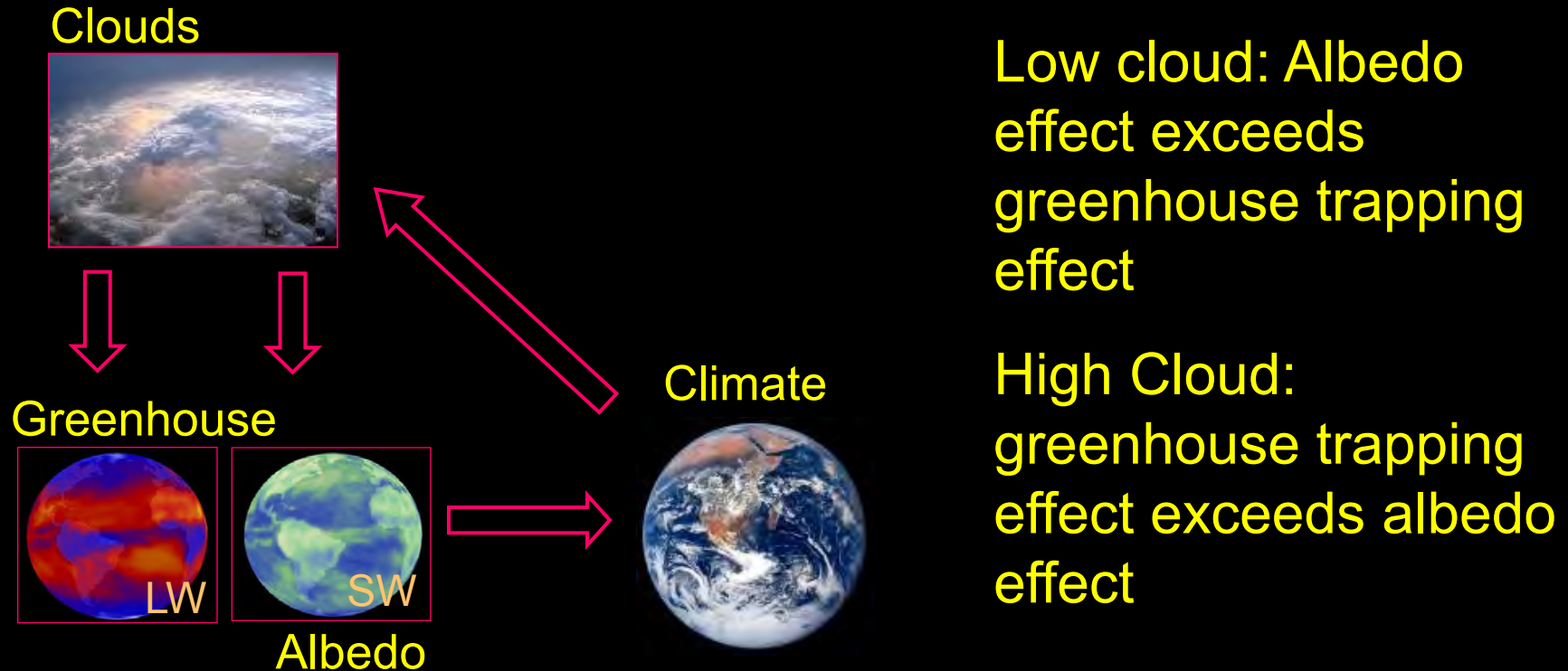


Regional & Seasonal Effects



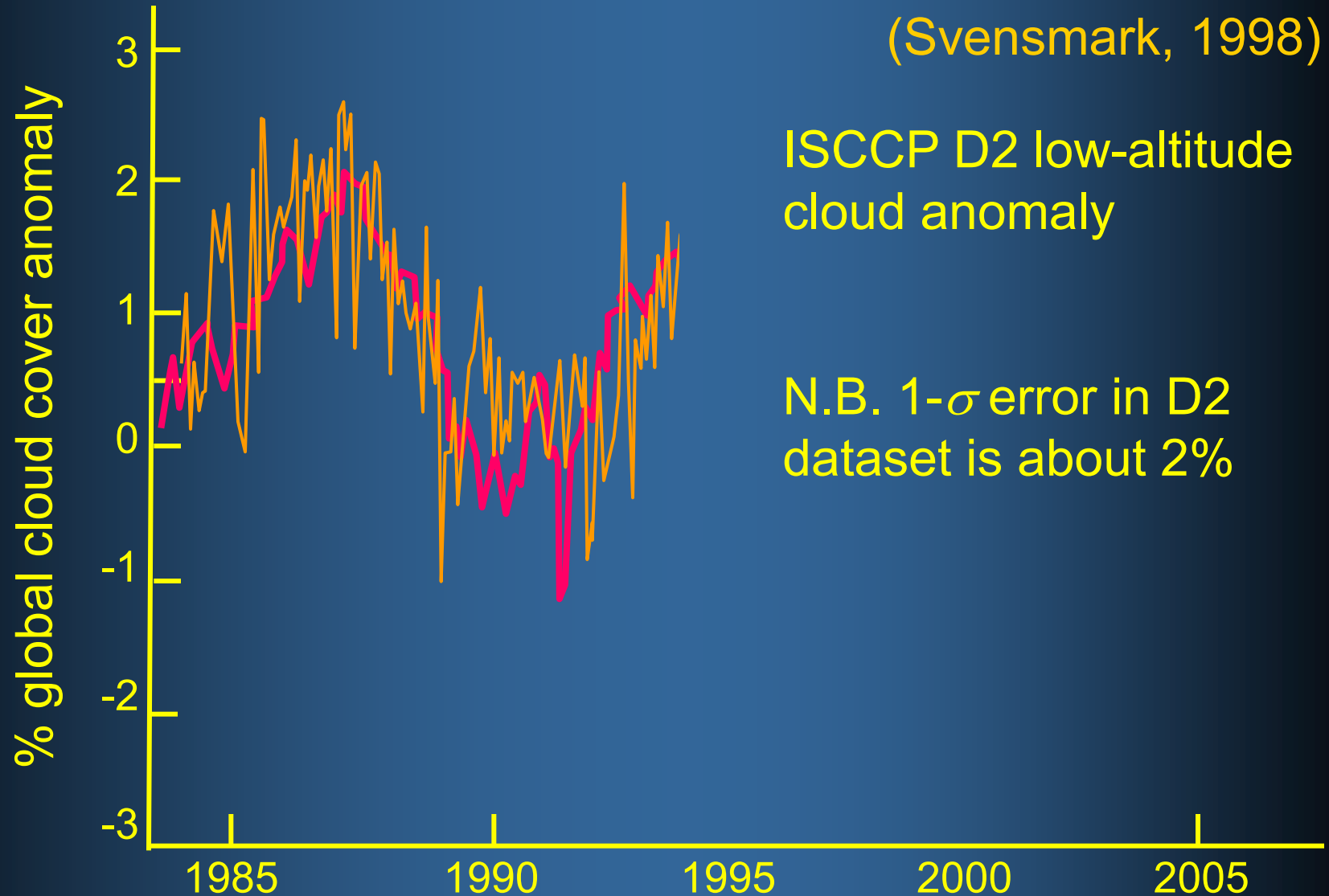
The Future

Cosmic Rays & Clouds

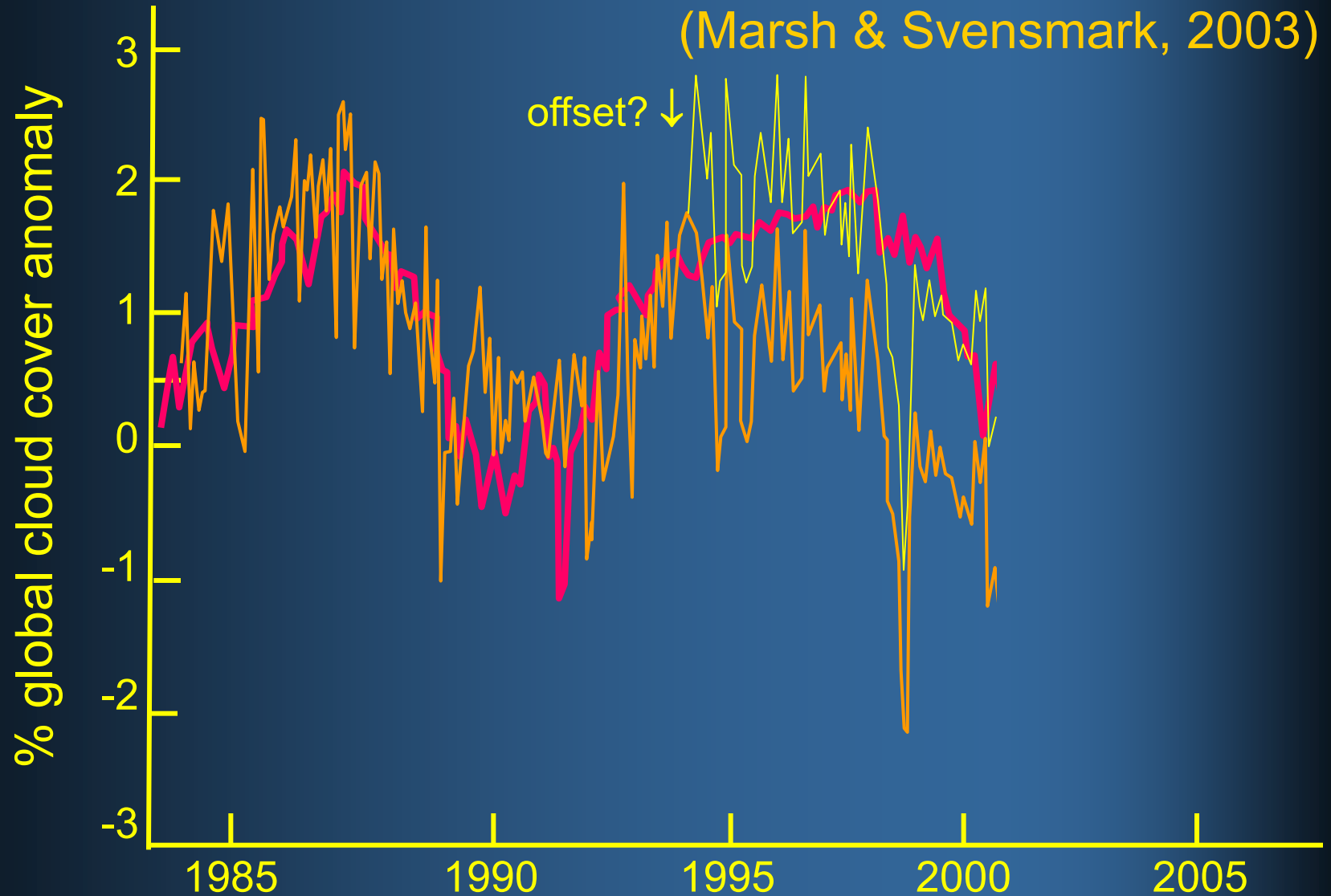


GCR fluxes fell over 1900-1985 so to contribute to warming Earth, they would have to generate low altitude cloud (so reduced albedo exceeds reduced greenhouse trapping)

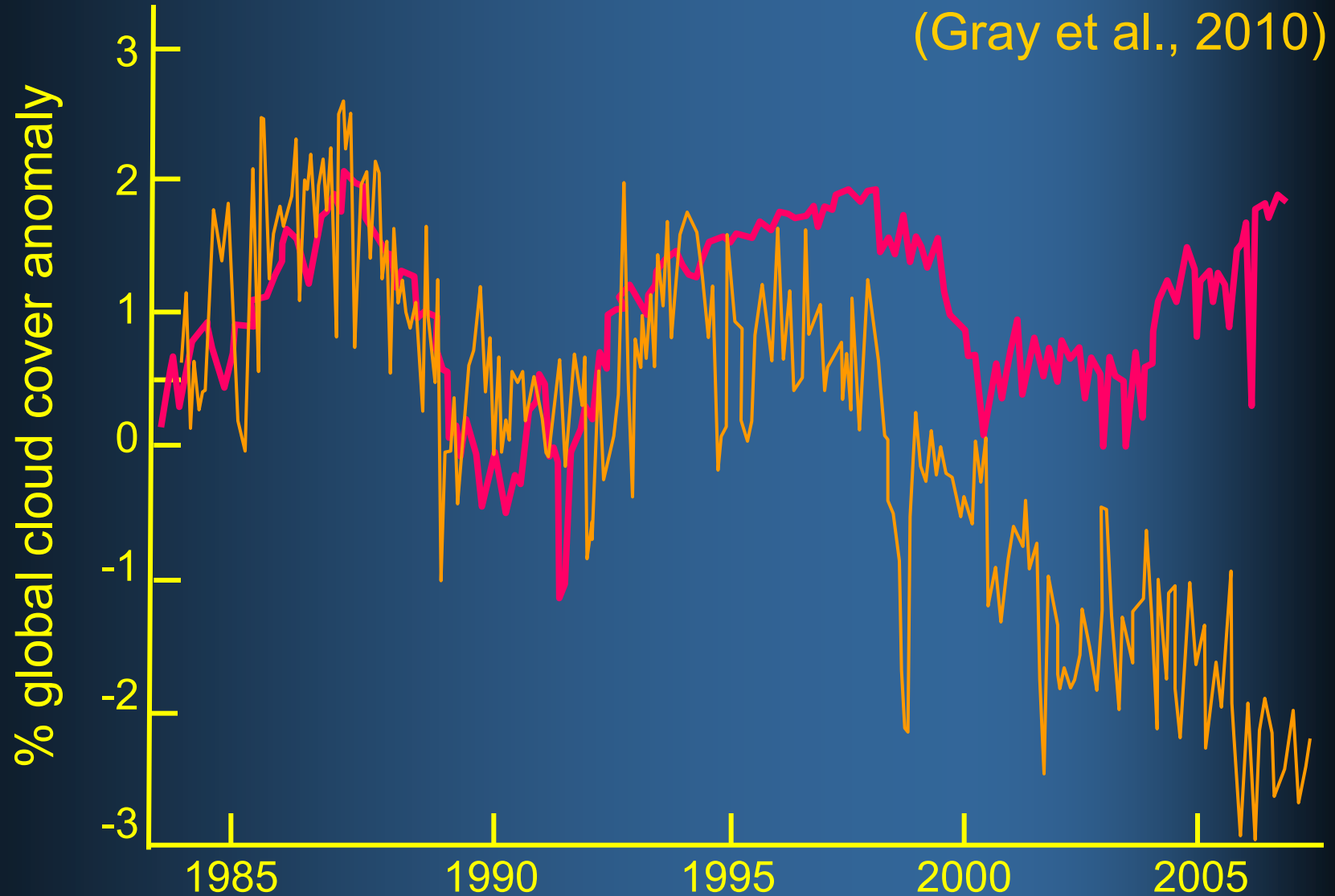
Global Cloud Cover Variation



Global Cloud Cover Variation



Global Cloud Cover Variation



New Evidence: Diffuse Fraction (DF)

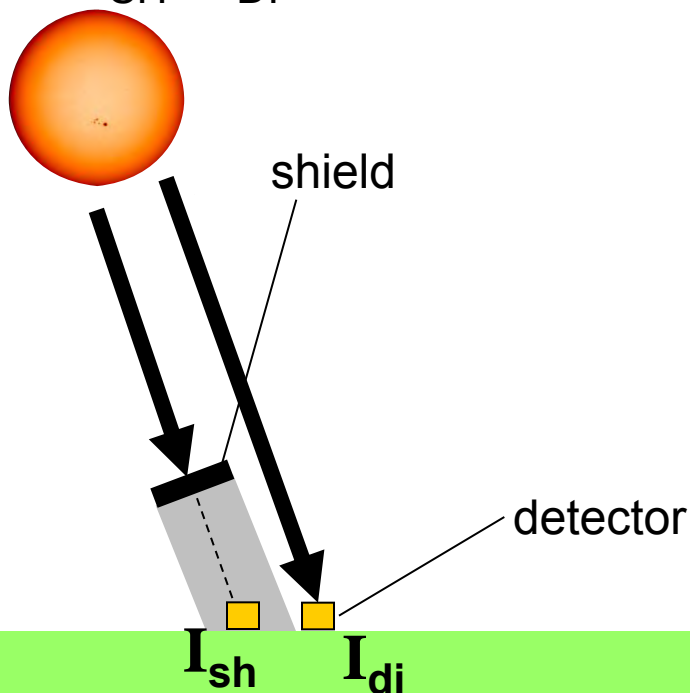
- ▶ Intensity in direct sunlight = I_{DI}
- ▶ Intensity in shade = I_{SH}
- ▶ Diffuse fraction, $DF = I_{SH} / I_{DI}$

- ▶ Measured at a number of sites since the 1950s

CLEAR SKY

$$I_{SH} \approx 0$$

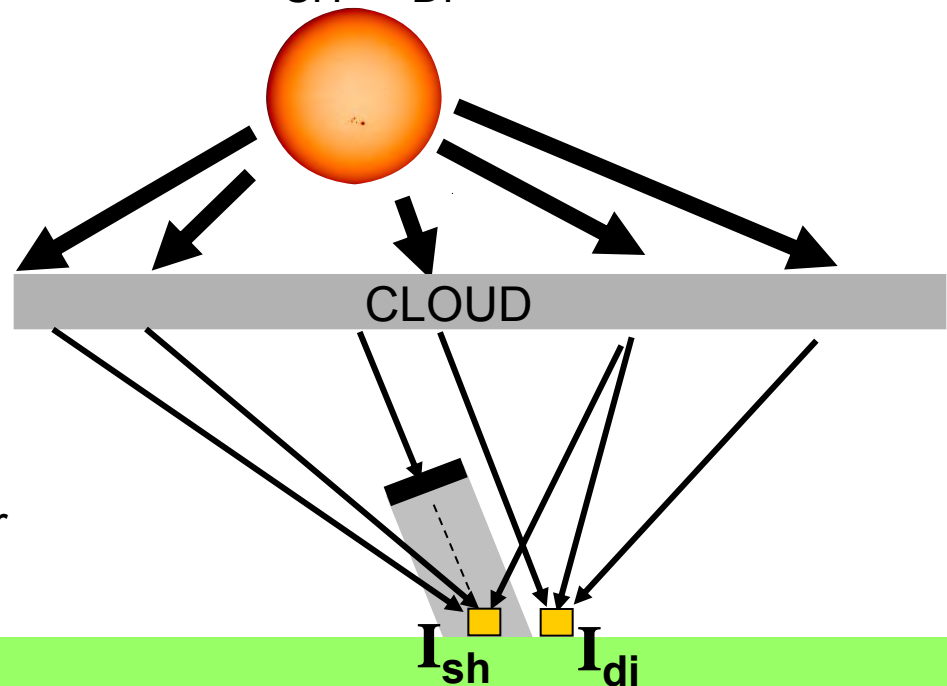
$$DF = I_{SH} / I_{DI} \approx 0$$



CLOUDY SKY (and/or aerosols)

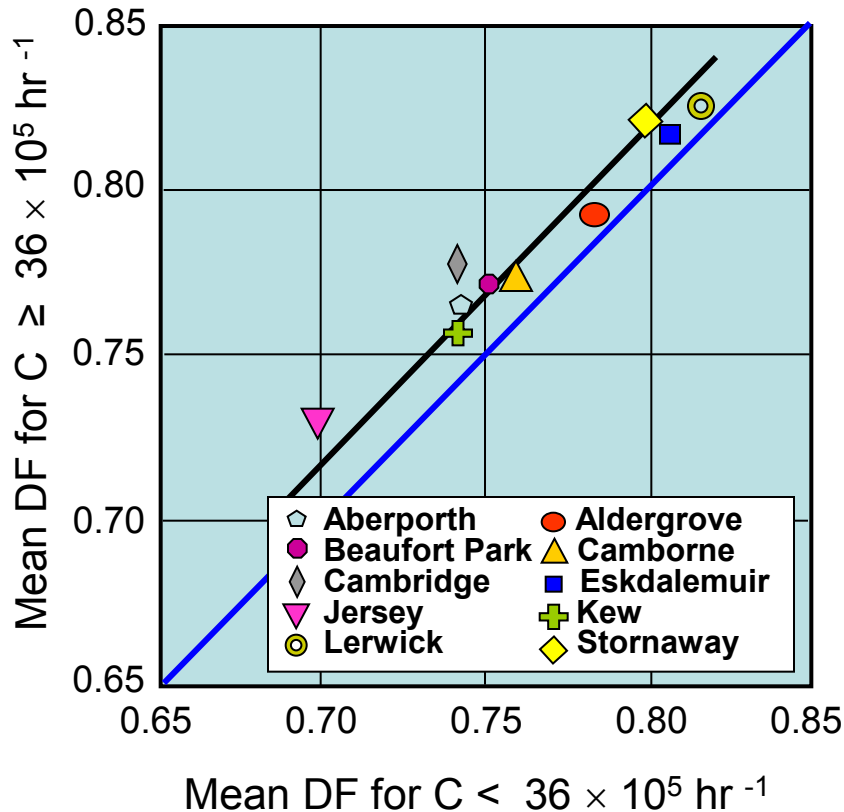
$$I_{SH} \approx I_{DI}$$

$$DF = I_{SH} / I_{DI} \approx 1$$

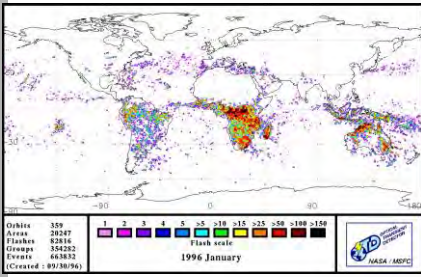


Global Cloud Cover Variation

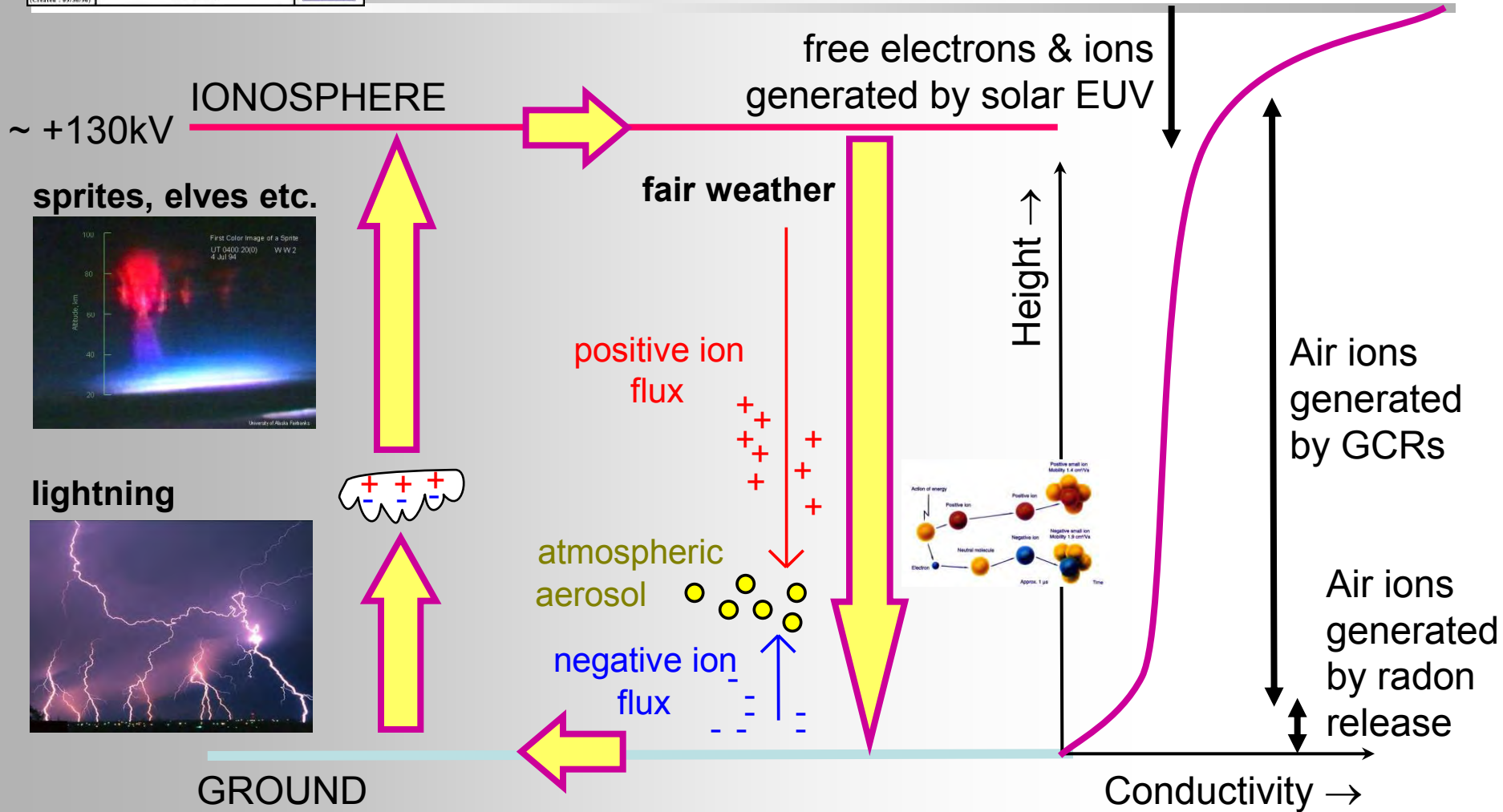
Harrison and Stephenson, Proc Roy. Soc (2006)



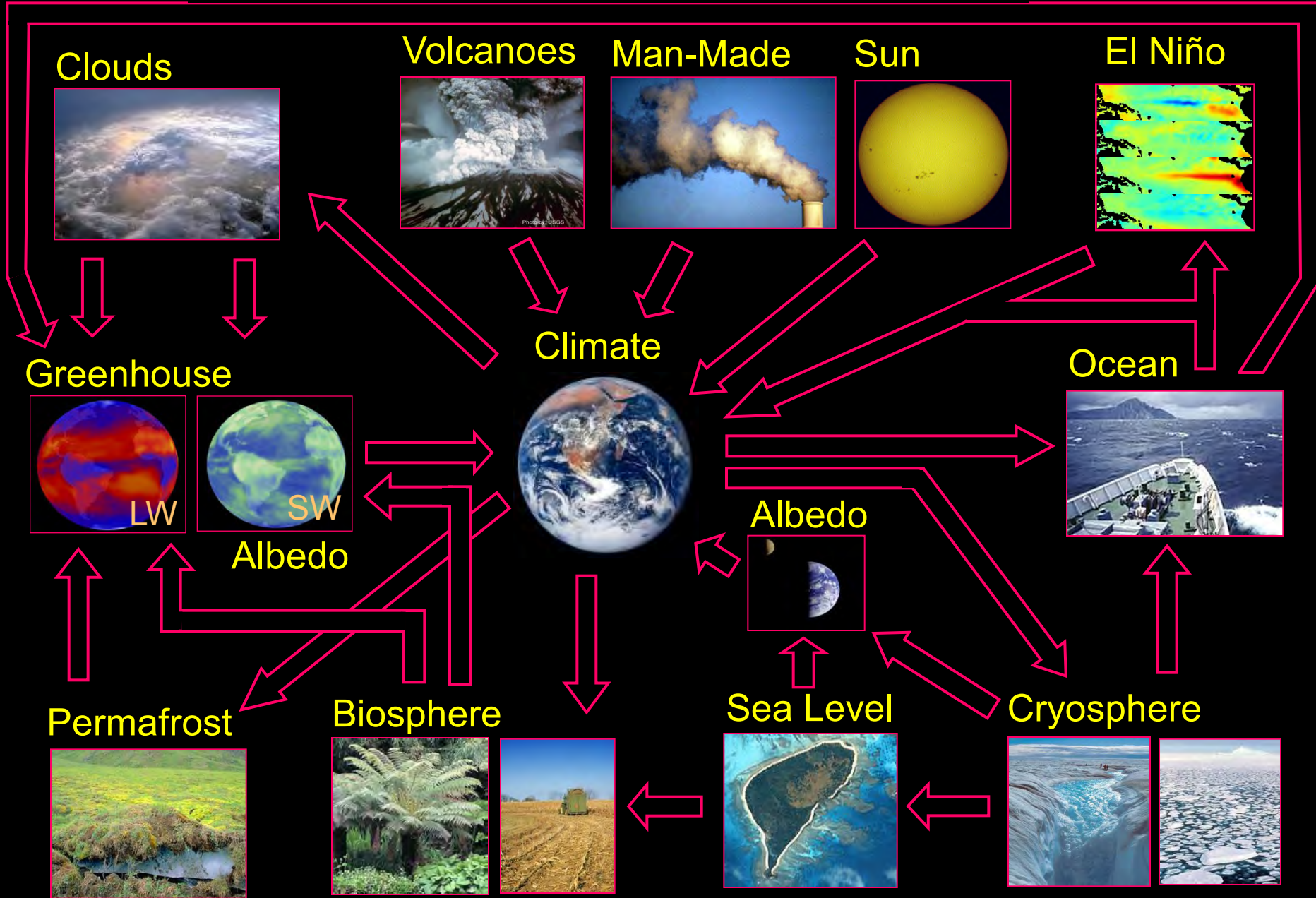
- Average DF for various stations in UK since 1950's
- Sorted according to the galactic cosmic ray flux ($>3\text{GeV}$) at Climax, C
- Mean DF for $C >$ threshold consistently exceeds mean DF for $C <$ threshold



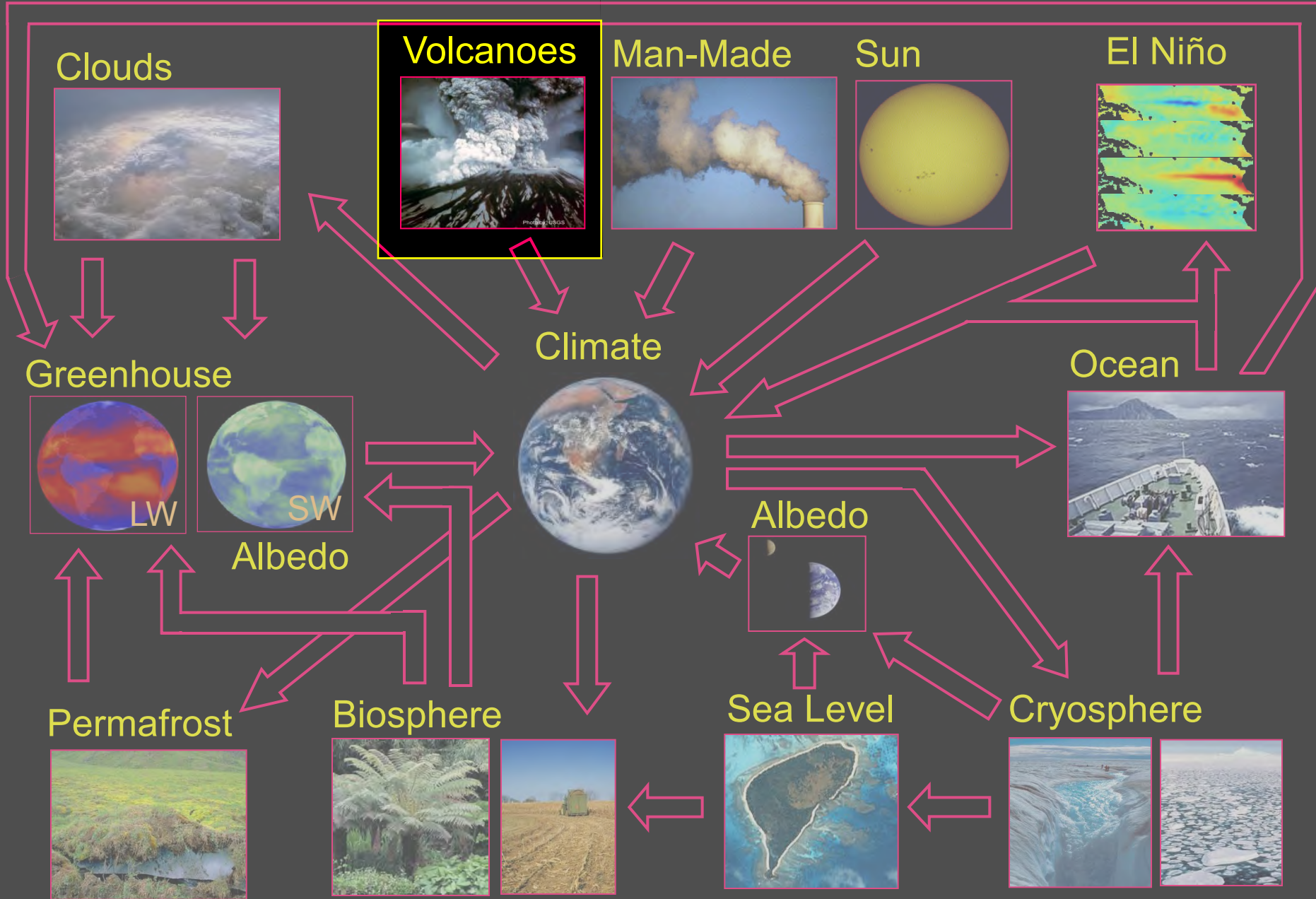
Global Electric Circuit



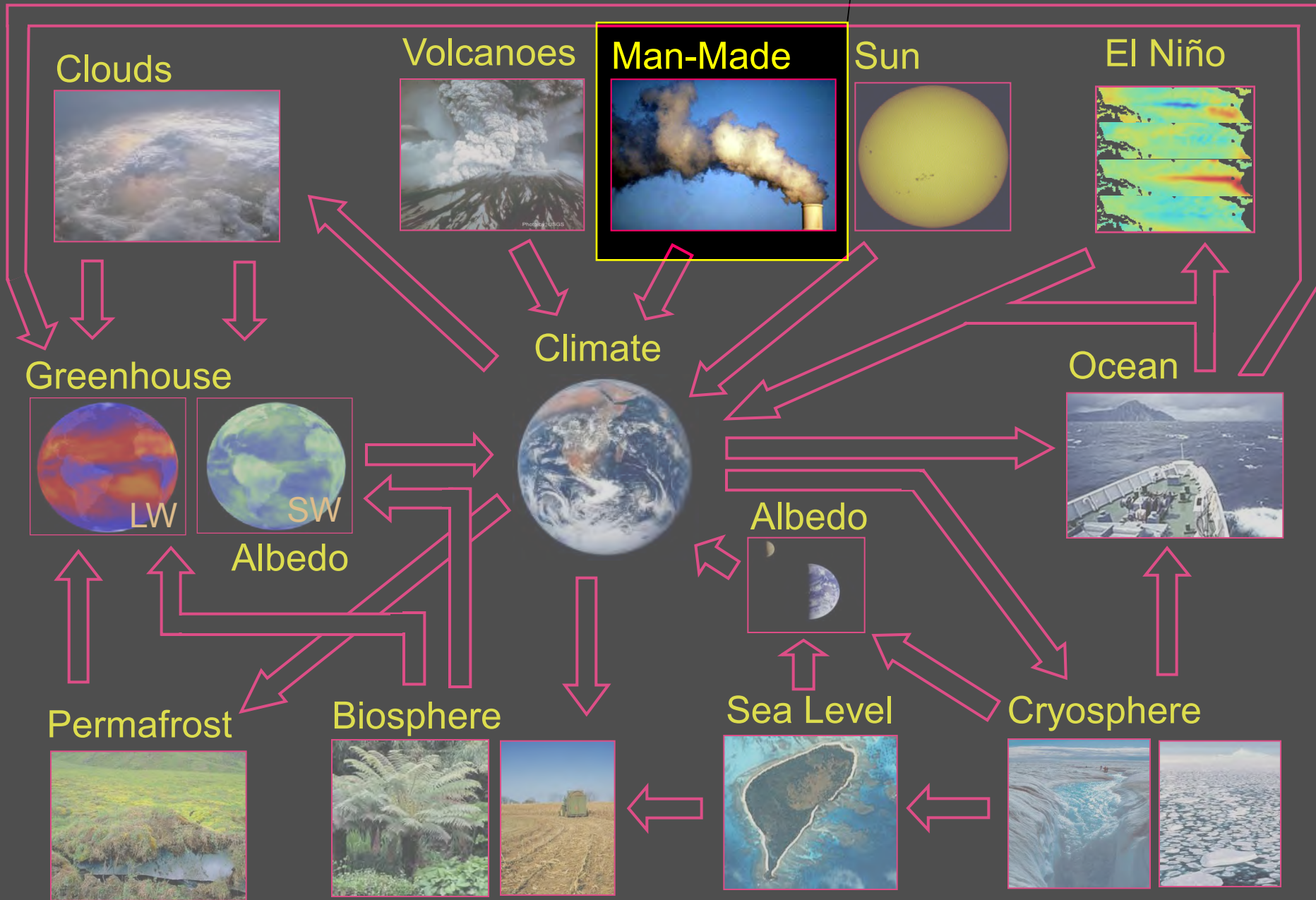
Climate System



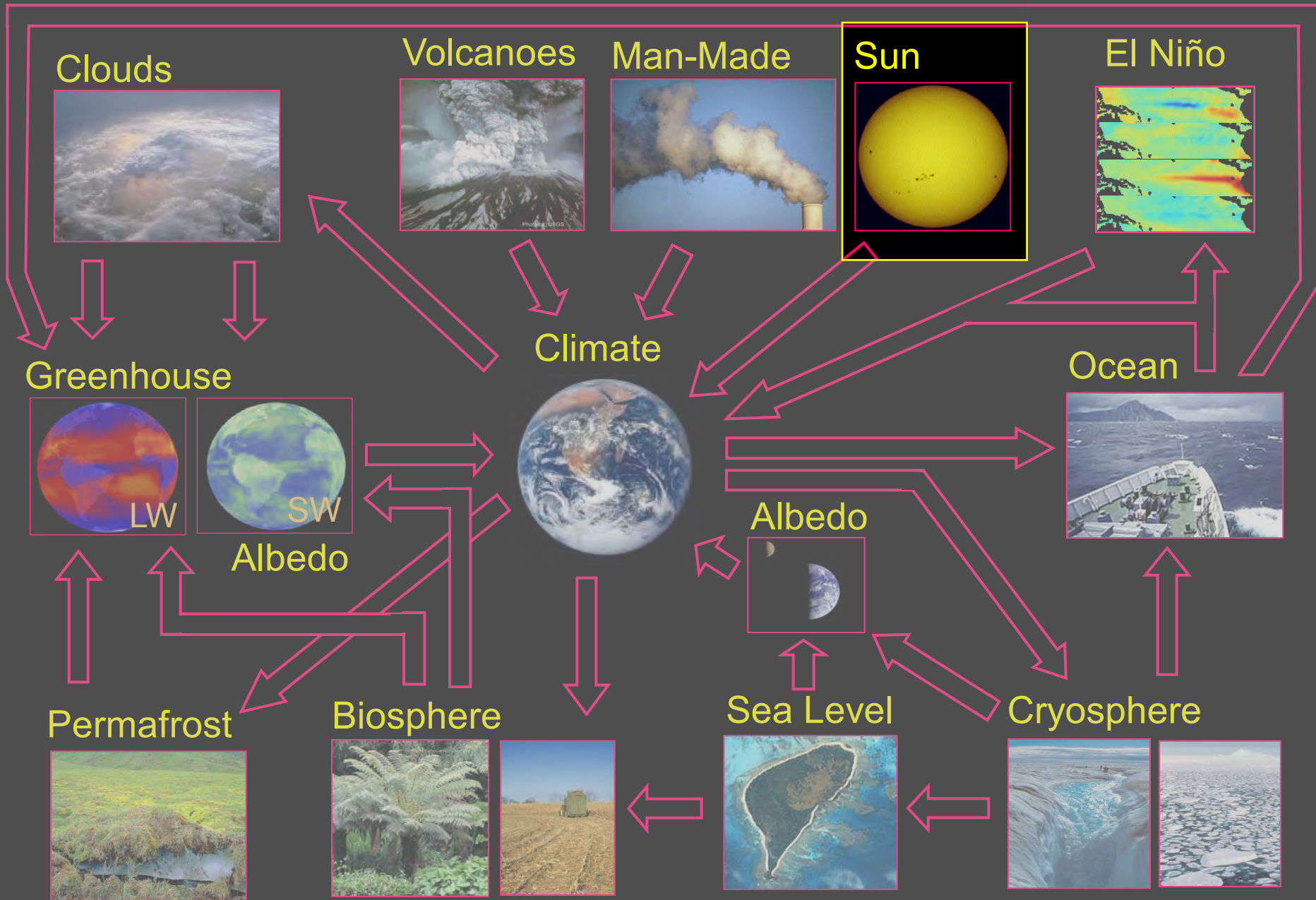
Geological Effects



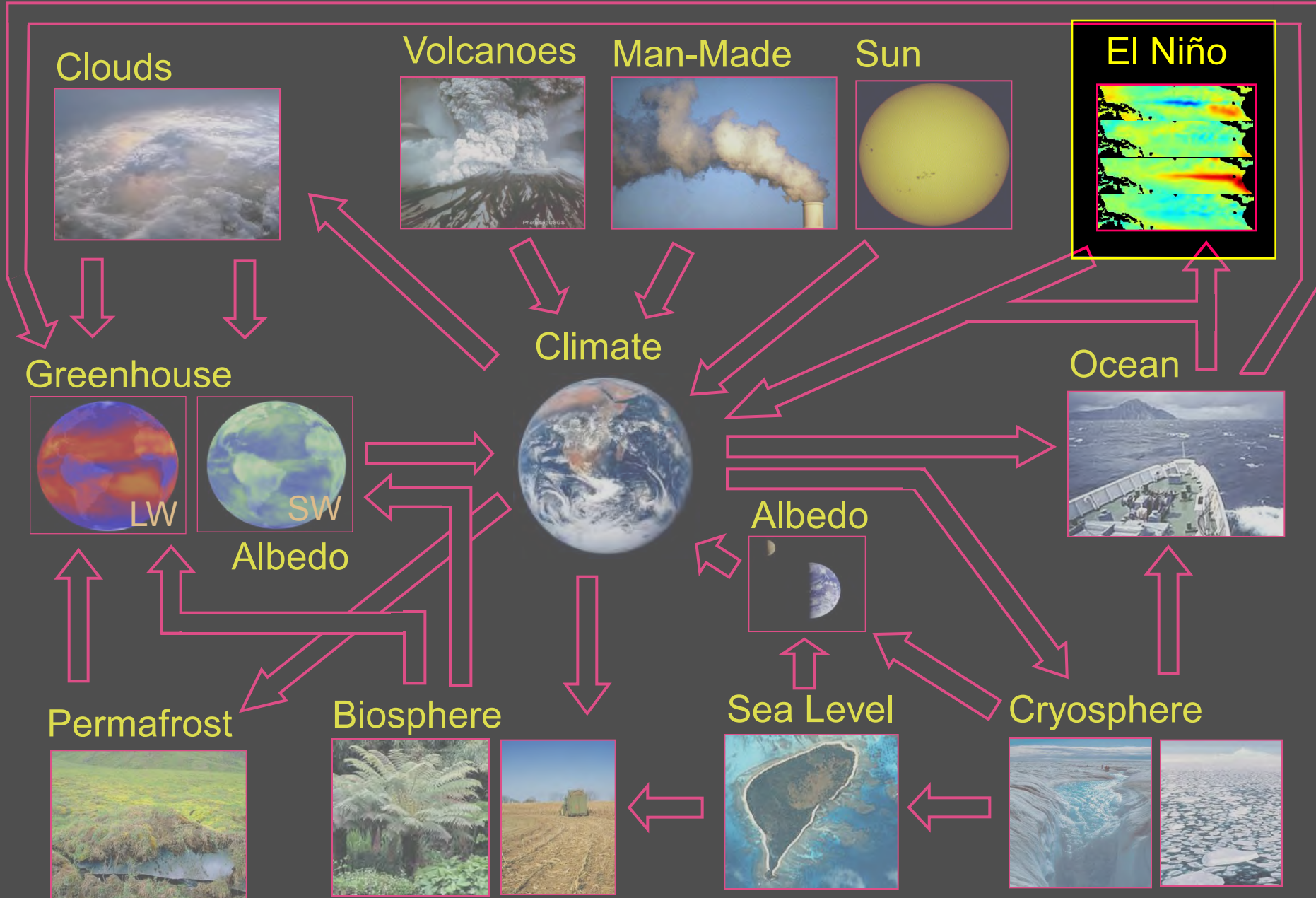
Anthropogenic Effects



Solar Influence



Short-Timescale Ocean Energy Exchange



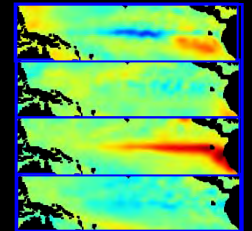
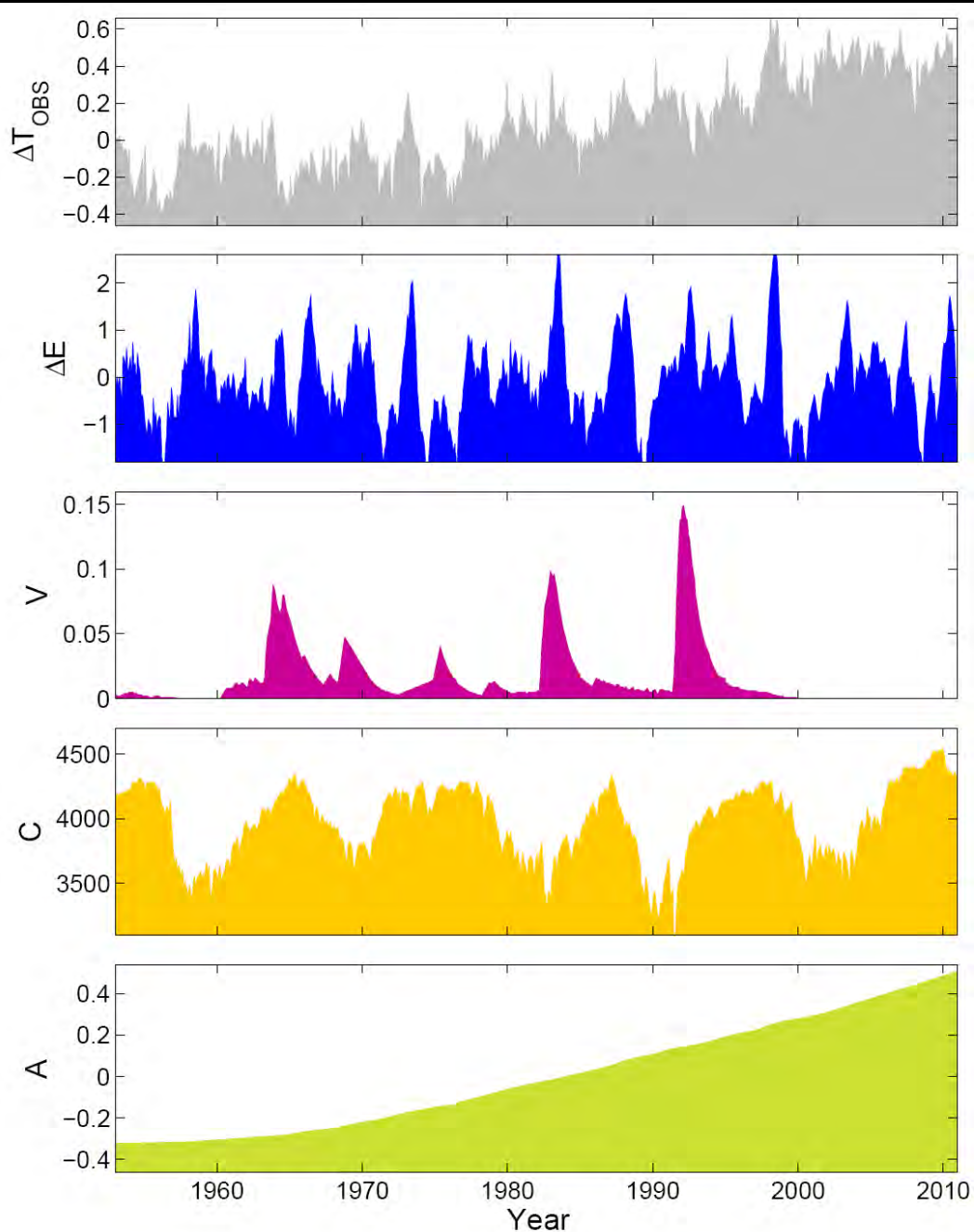
Observed Global
Surface Air
Temperature
Anomaly, ΔT_{OBS}

ENSO N3.4 index
Anomaly, ΔE

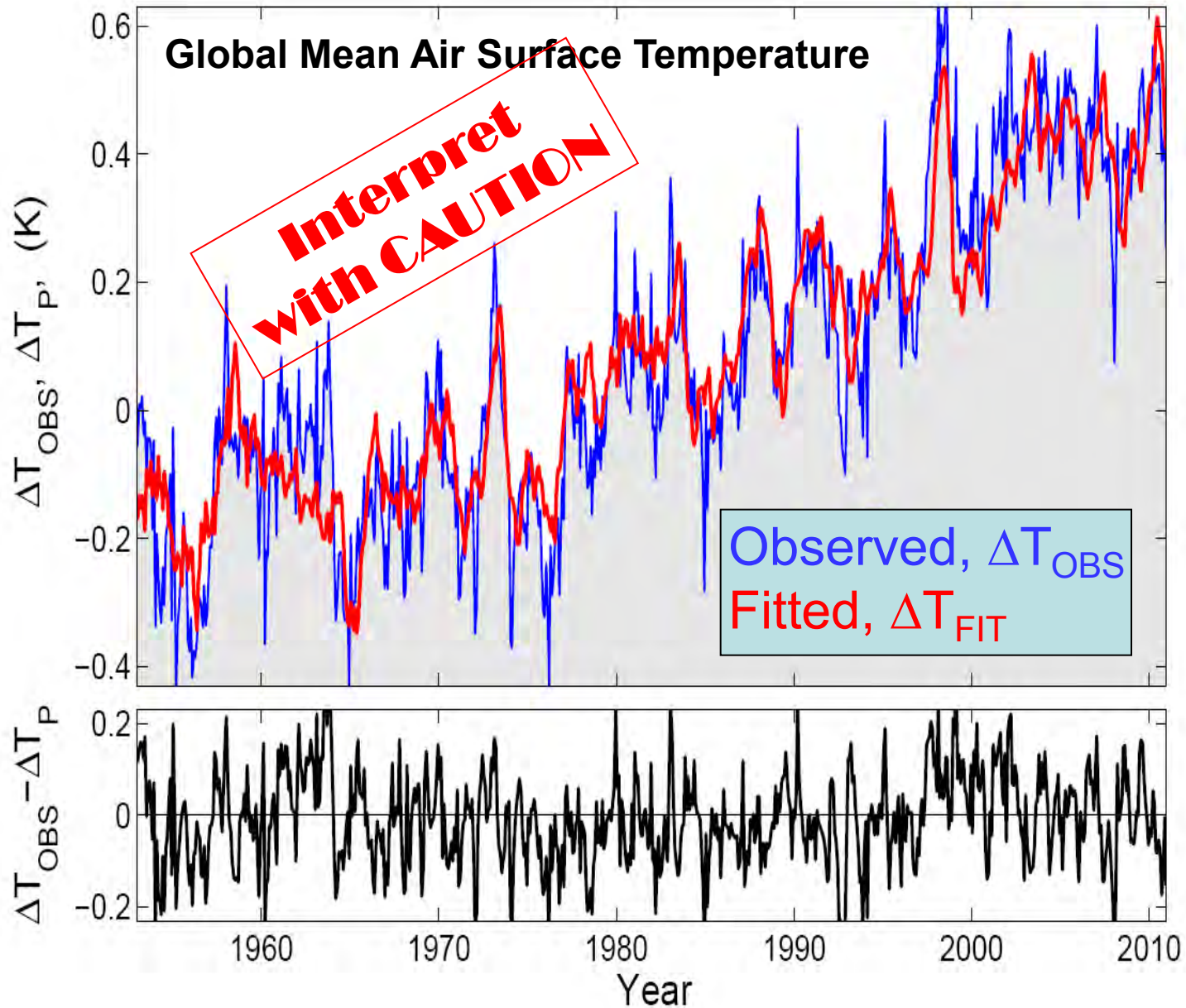
Mean Optical
Depth (AOD) at
550 nm, V

Cosmic Ray
Counts at
Climax, C

Anthropogenic
forcing, A ,
(greenhouse
gases, aerosols, &
land use change)



- fit to observed GMAST anomaly obtained using the Nelder-Mead simplex (direct search) method



(Lockwood, 2008)

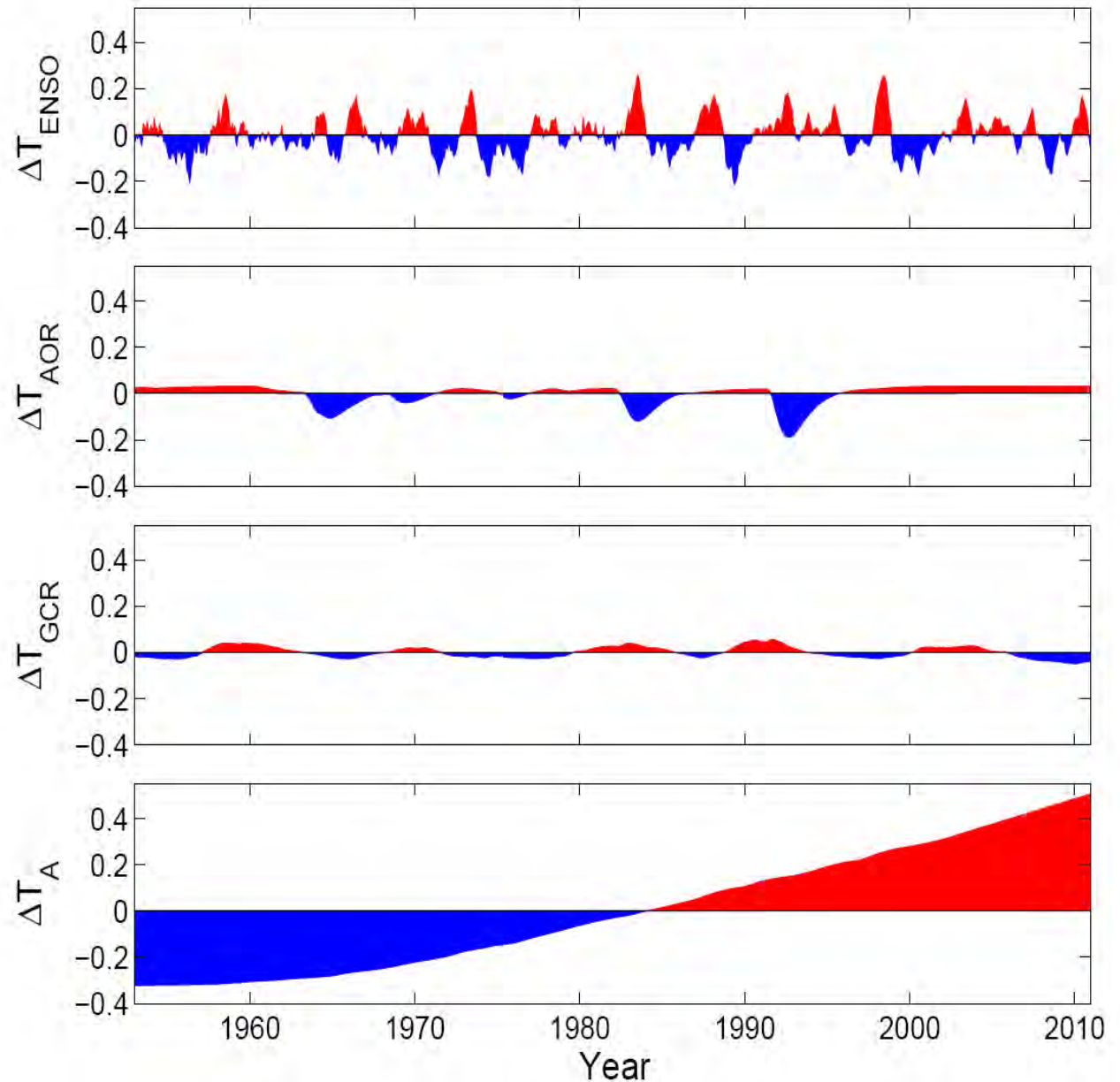


- when a fit has too many degrees of freedom
- can start to fit to the noise in the training subset, which is not robust throughout the data (fit has no predictive power)
- recognised pitfall when quasi-chaotic behaviours give large internal noise such as in climate science¹ and population growth²
- often not recognised in space physics where systems tend to be somewhat more deterministic with lower internal variability.

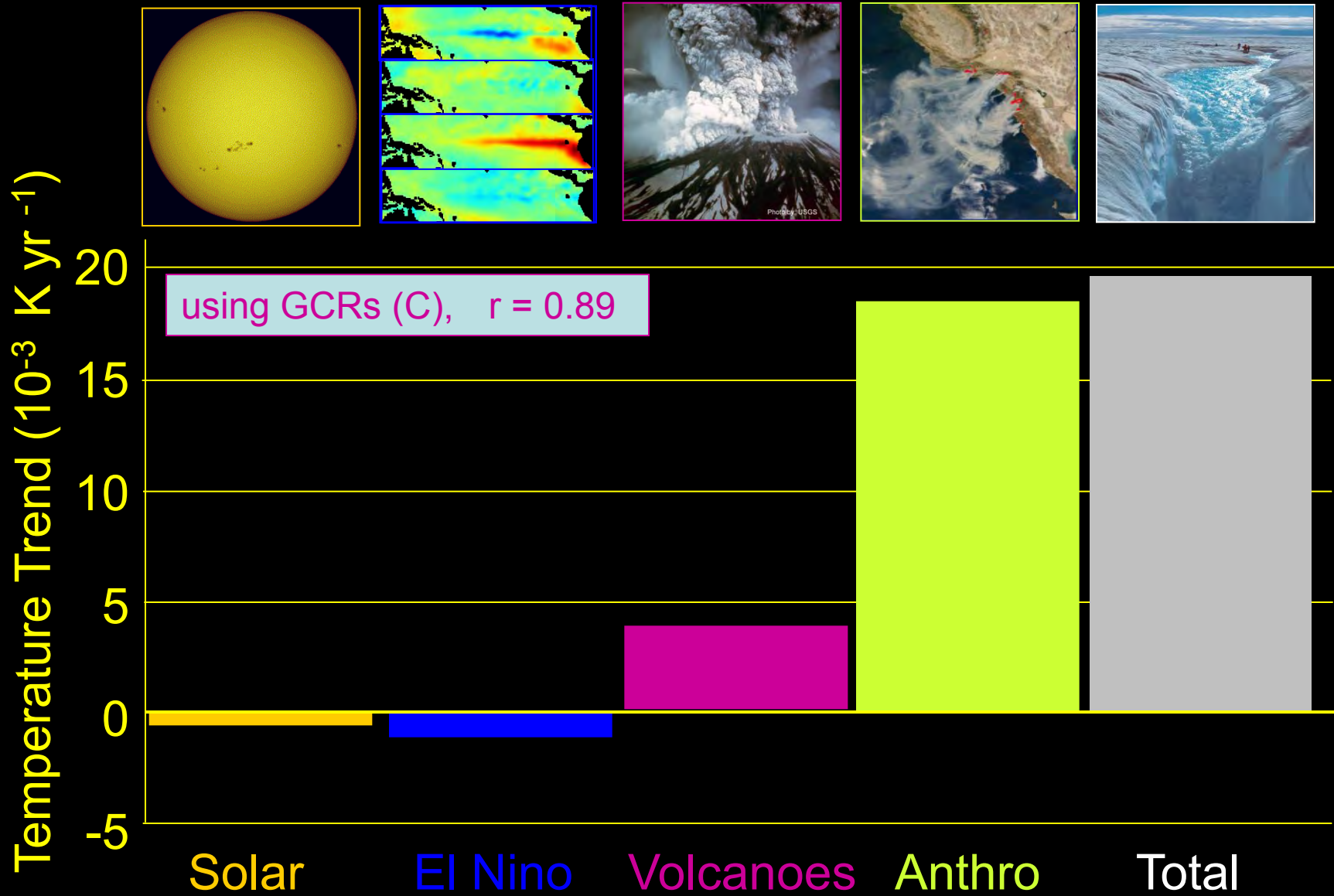
¹ e.g. Knutti et al. (2006) *J. Climate*, DOI: 10.1175/JCLI3865.1

² e.g. Knappe and de Valpine (2011) *Proc. Roy. Soc. London B*, DOI: 10.1098/rspb.2010.1333

- Weighted contributions to best fit variation, T_p (uses Climax GCR counts to quantify solar effect)

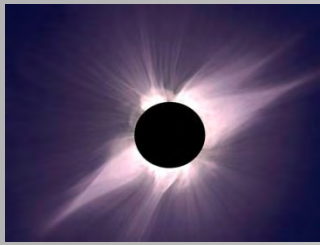


(updated from Lockwood, 2008)



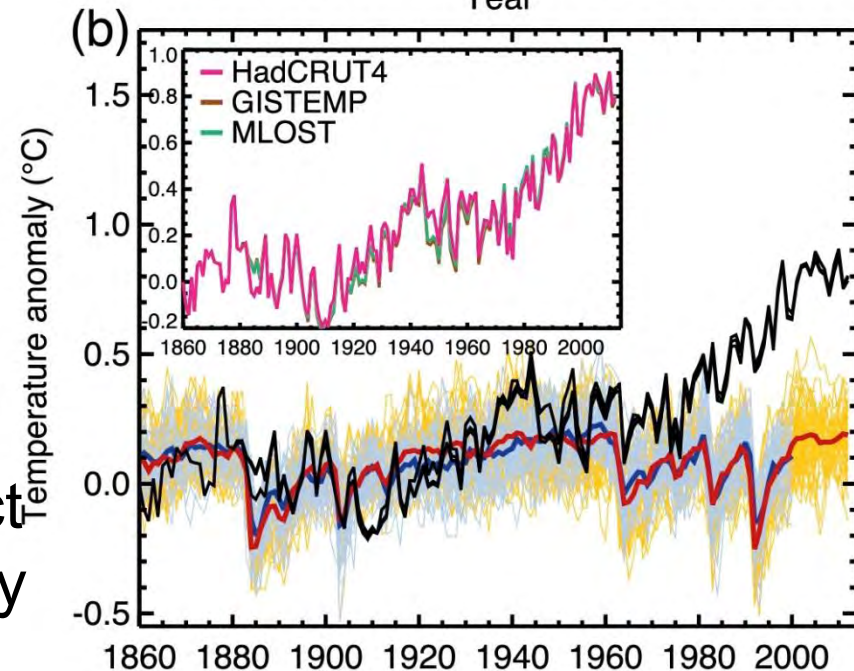
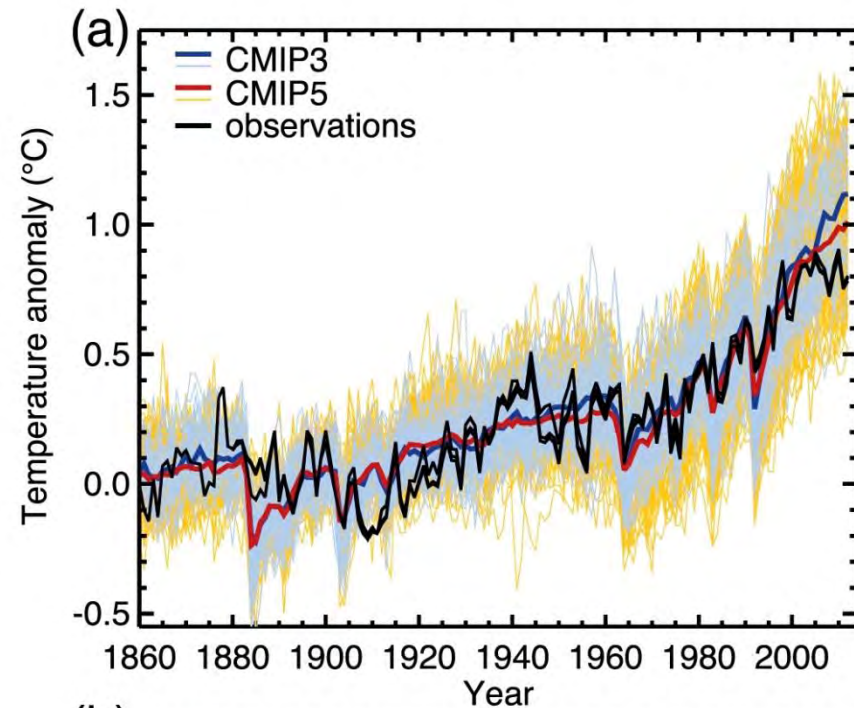
1987-present

(Lockwood, 2008)

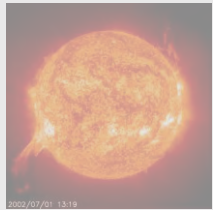


Detection-Attribution

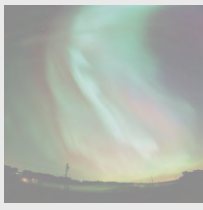
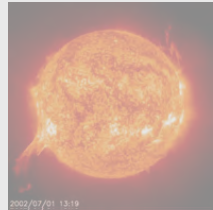
- Use models to avoid over-fitting problem
- The idea is that models, started from slightly different initial conditions, can reproduce the internal variability of the climate system
- Produce an ensemble of many model runs for set inputs and then compare mean or median with observations
- Runs with no anthropogenic effect differ from observed GMAST rise by more than the internal noise level



Solar Variability: Effects on Climate?



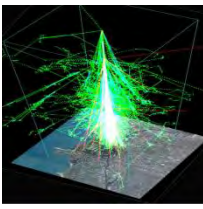
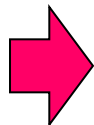
Solar Outputs



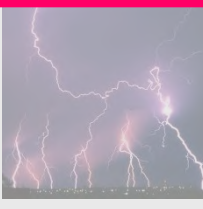
Solar Variability



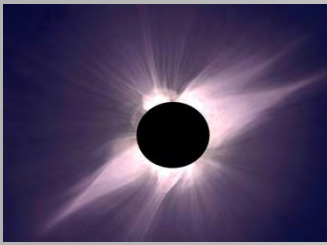
Global Effects



Regional & Seasonal Effects

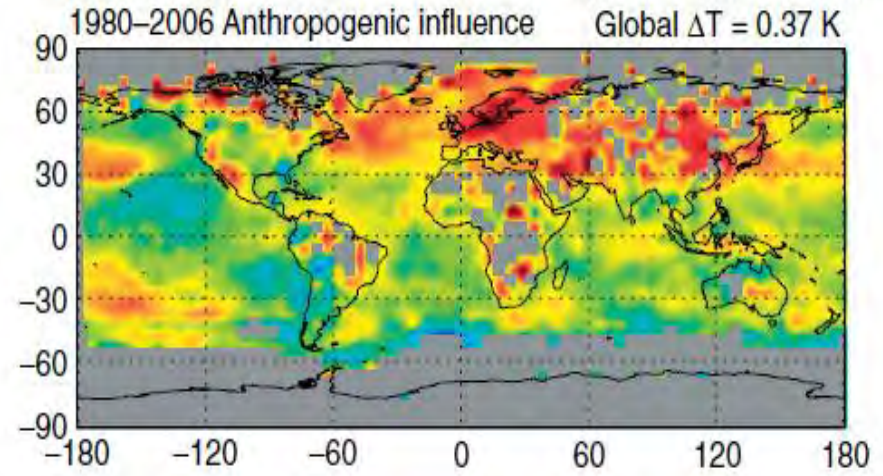
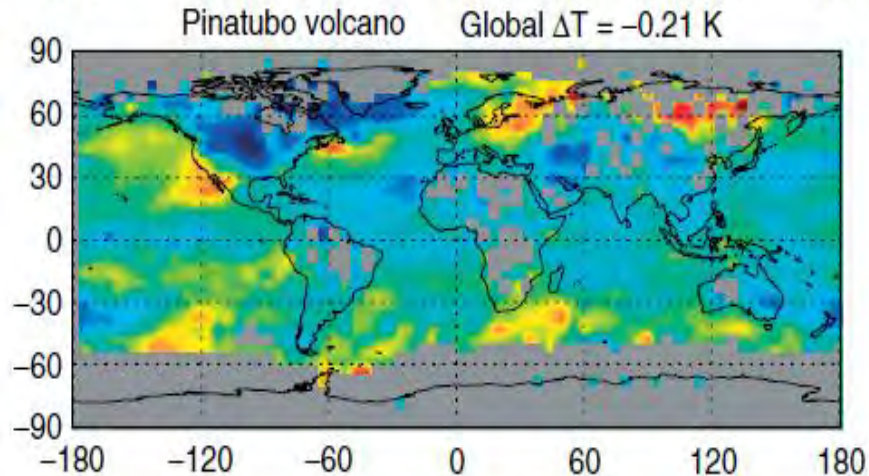
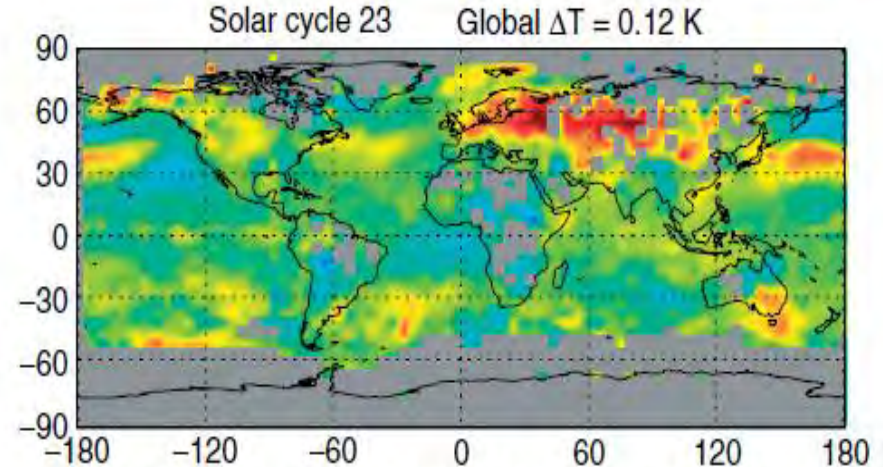
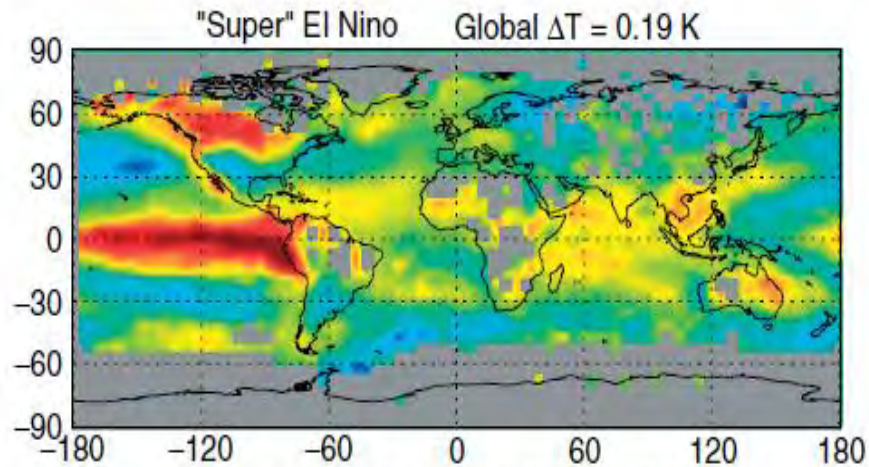


The Future

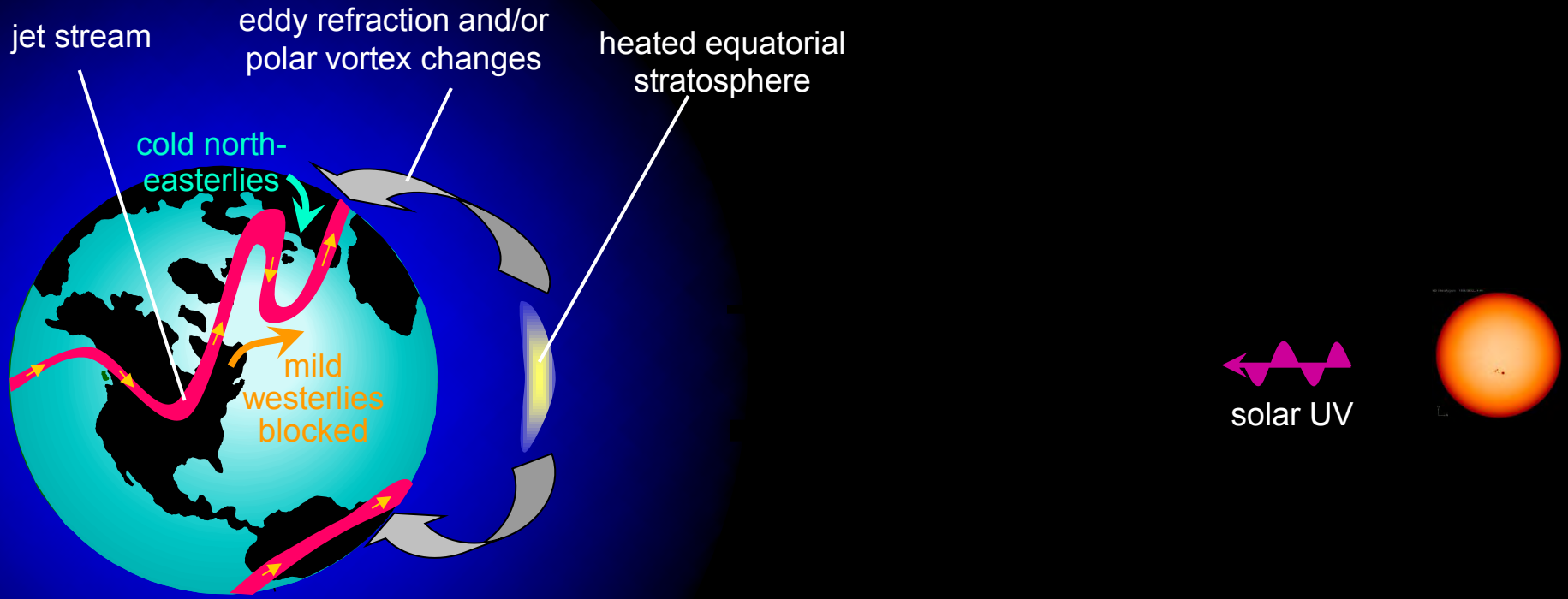


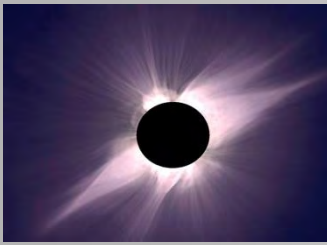
Regional Analysis

(Lean and Rind, 2008)



“Top-down” Solar Modulation

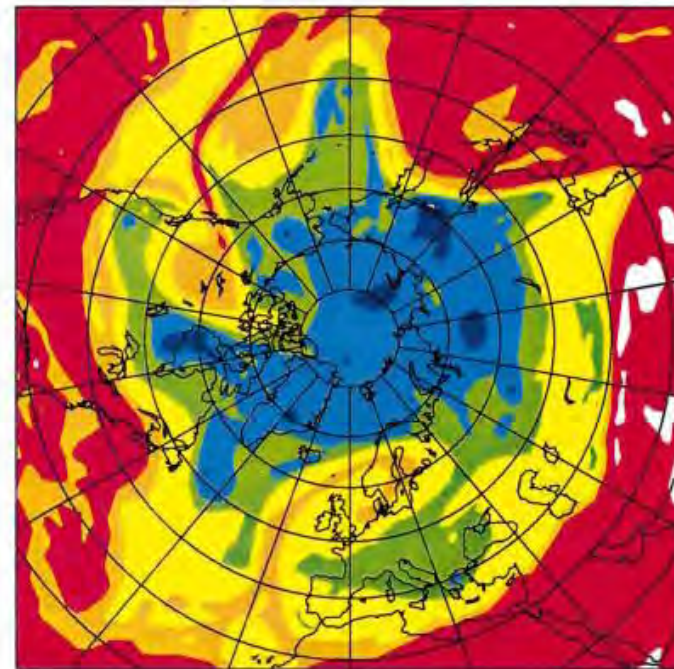
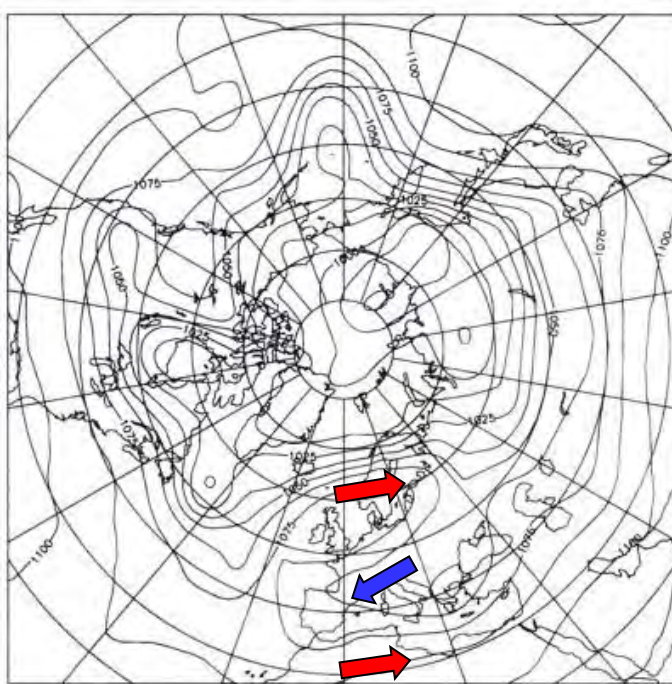




Atlantic blocking events

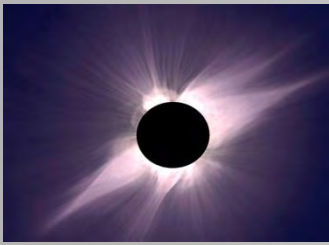
(Plelly and Hoskins, 2003)

- ▶ blocking events are large long-lived anticyclones which disrupt easterly flow of storms, bifurcating the jet stream and, in winter, causing cold winds from the east over Europe



■ 285 - 300 ■ 300 - 315 ■ 315 - 330 ■ 330 - 345 ■ 345 - 360 ■ 360 - 375

Example at 12UT, 21 Sept, 1998: on the potential vorticity $PV=2$ surface
(a) 250-hPa geopotential height (b) potential temperature θ (K)

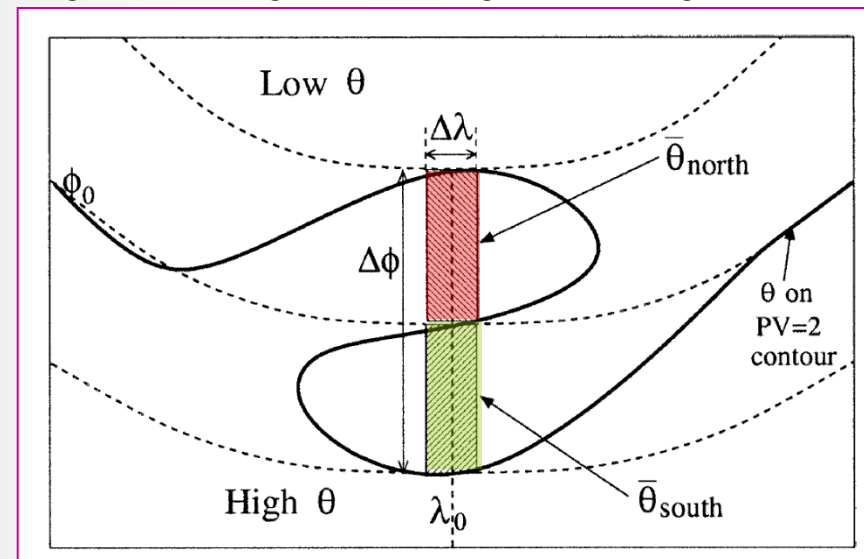


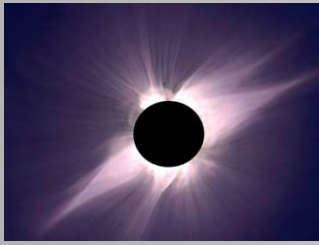
Blocking Intensity Indices

- ▶ *Lejenäs and Økland (1983)* required a region of easterly winds and used $Z(\lambda, \phi_o + \Delta\phi/2) - Z(\lambda, \phi_o - \Delta\phi/2)$ where Z is a constant height geopotential, λ is the longitude and ϕ the latitude
- ▶ *Barriopedro et al. (2006, 2008)* used $BI = 100 \times \{ [Z(\lambda_o, \phi_o) / RC] - 1 \}$ where $RC = \{ Z(\lambda_o + \Delta\lambda, \phi_o) - Z(\lambda_o - \Delta\lambda, \phi_o) \} / 2$

- ▶ *Pelly and Hoskins (2006, 2008)* used mean potential temperature θ in the red and green areas of the plot $B =$

$$(2/\Delta\phi) \int_{\phi_o}^{\phi_o + \Delta\phi/2} \theta \, d\phi - (2/\Delta\phi) \int_{\phi_o - \Delta\phi/2}^{\phi_o} \theta \, d\phi$$

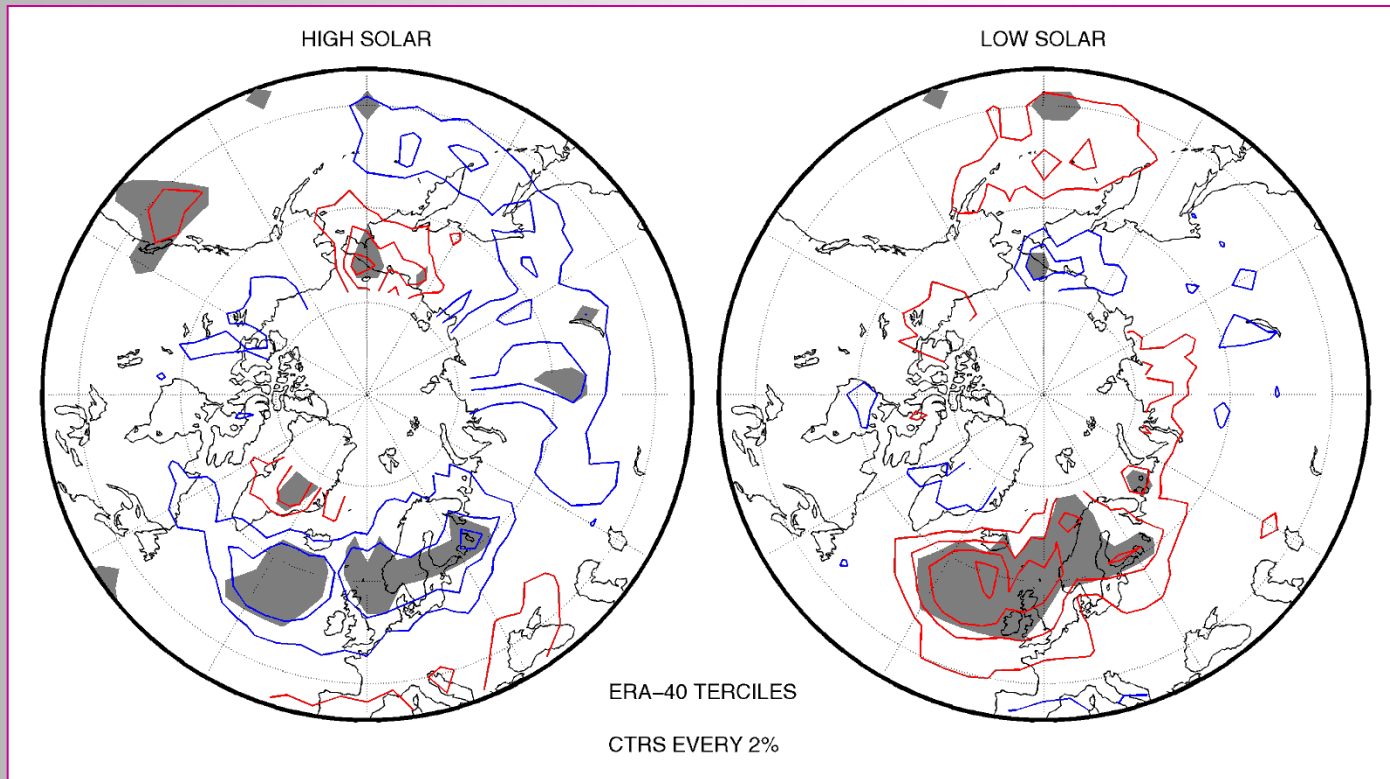




ERA-40 Analysis of Blocking Index

(change of terciles relative to whole set)

(*Woollings et al, GRL.,2010*)



► sorted using open solar flux F_S

High/Low solar activity gives reduced/enhanced (up to 8%) blocking over east Atlantic and Europe (symmetric effect)

Consistent and localised effect

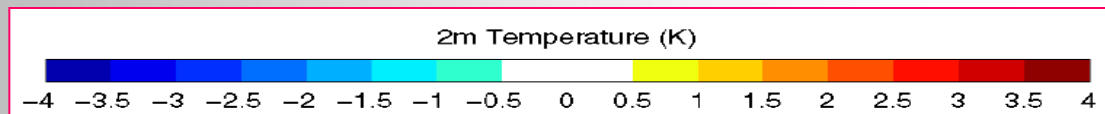
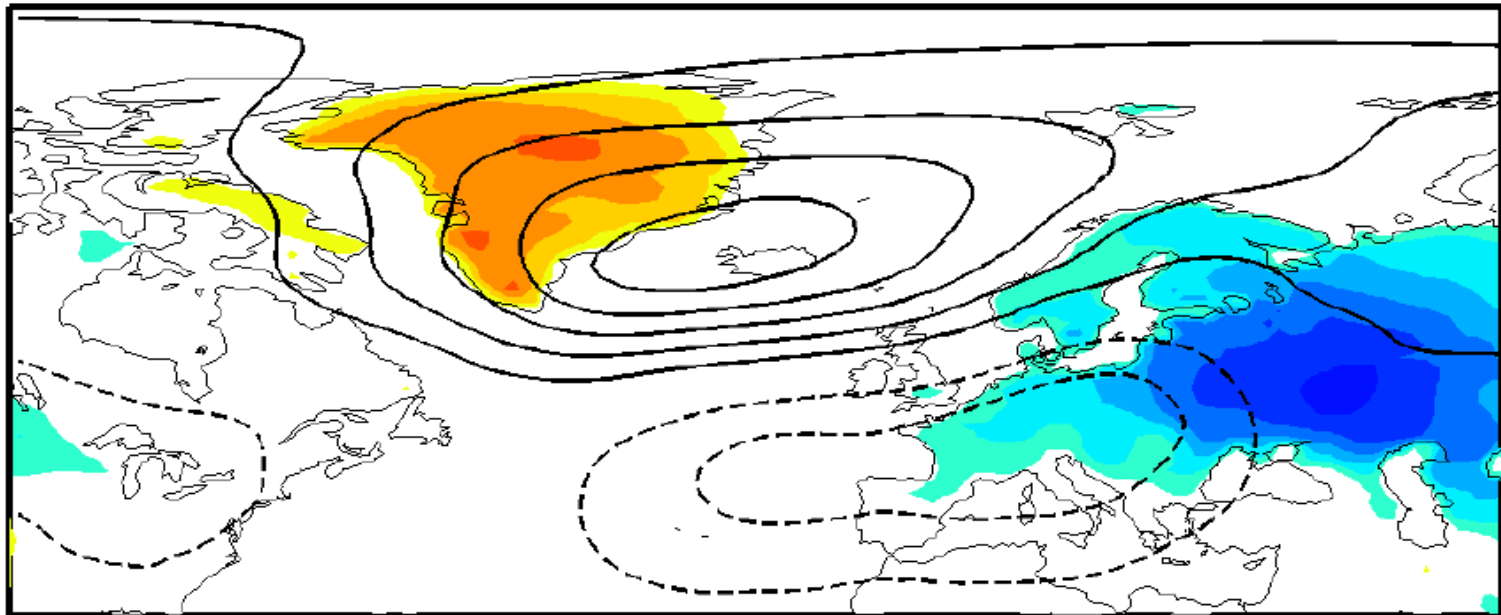
Grey area shows significance from Monte-Carlo technique > 95%



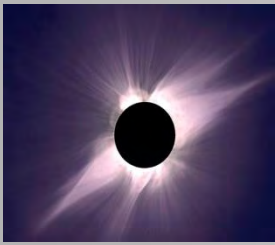
ERA-40 Analysis of DJF temperatures & circulation (difference of high and low tercile subsets)

(Woollings et al, GRL.,2010; see also Barriopedro et al., JGR, 2008)

SOLAR: LOW – HIGH



- ▶ sorted using open solar flux F_S
- Low solar activity gives lower surface temperatures in central England
- Effect much stronger in central Europe
- Analysis shows a distinct system to NAO

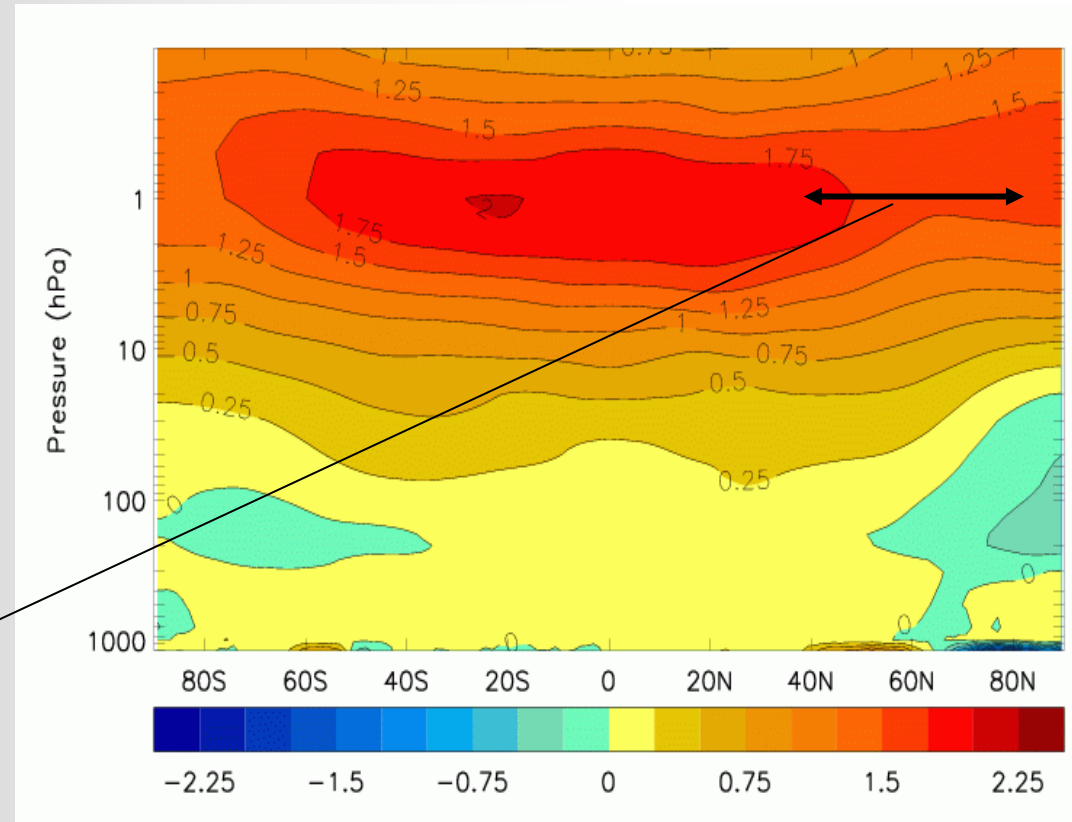


Modelled solar maximum-solar minimum temperatures



(Ineson et al, Nature Geosci., 2011)

- ▶ Heating effect only (no [O₃] change)
- ▶ HADGEM3rev1.1 GCM, 85 atmos and 42 ocean levels.
- ▶ Uses the SORCE max-min UV spectrum $S_S(\lambda)$
- ▶ Increased meridional temperature gradient → increase in westerly flow



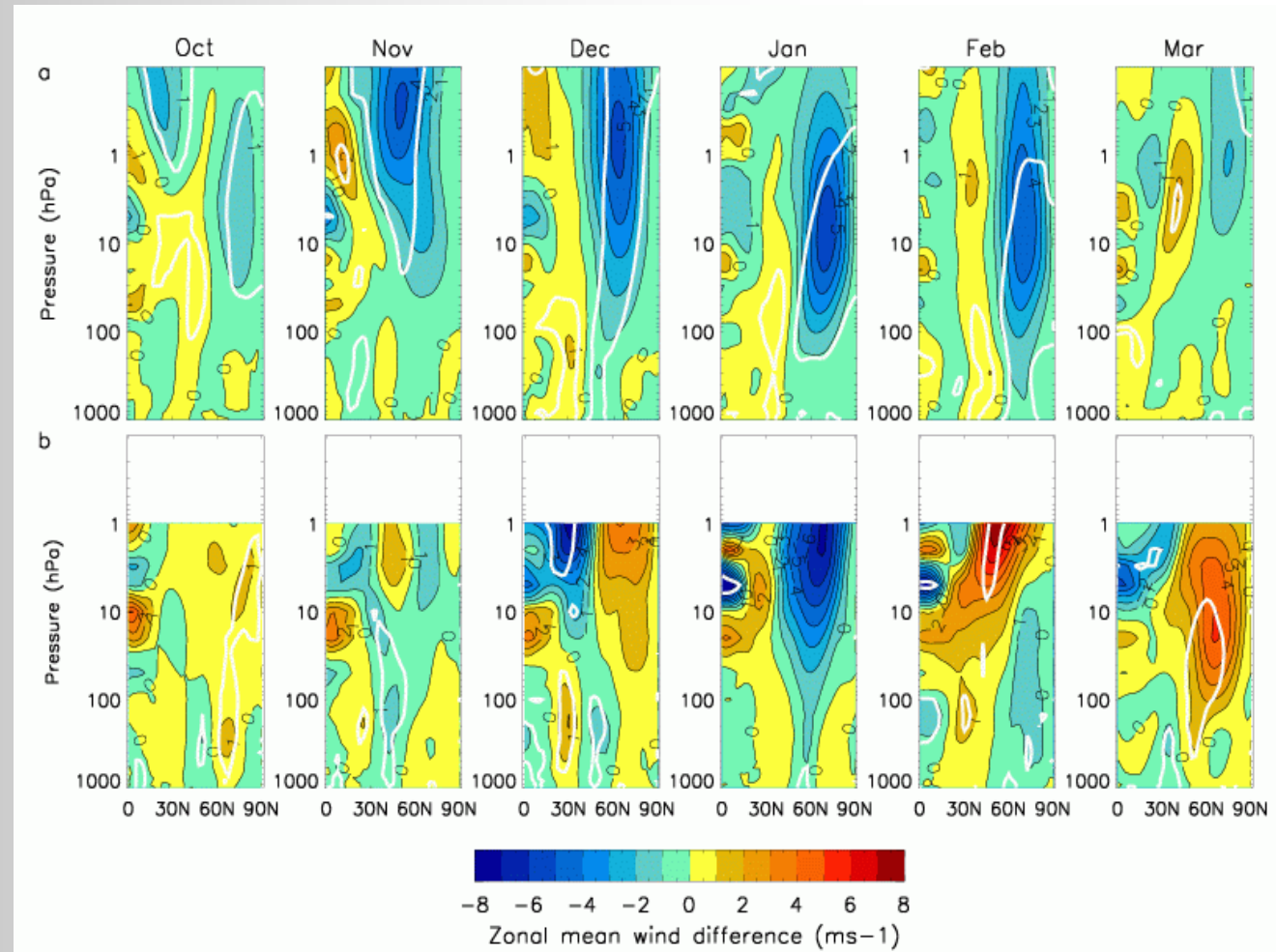


Modelled solar maximum-solar minimum zonal wind speed

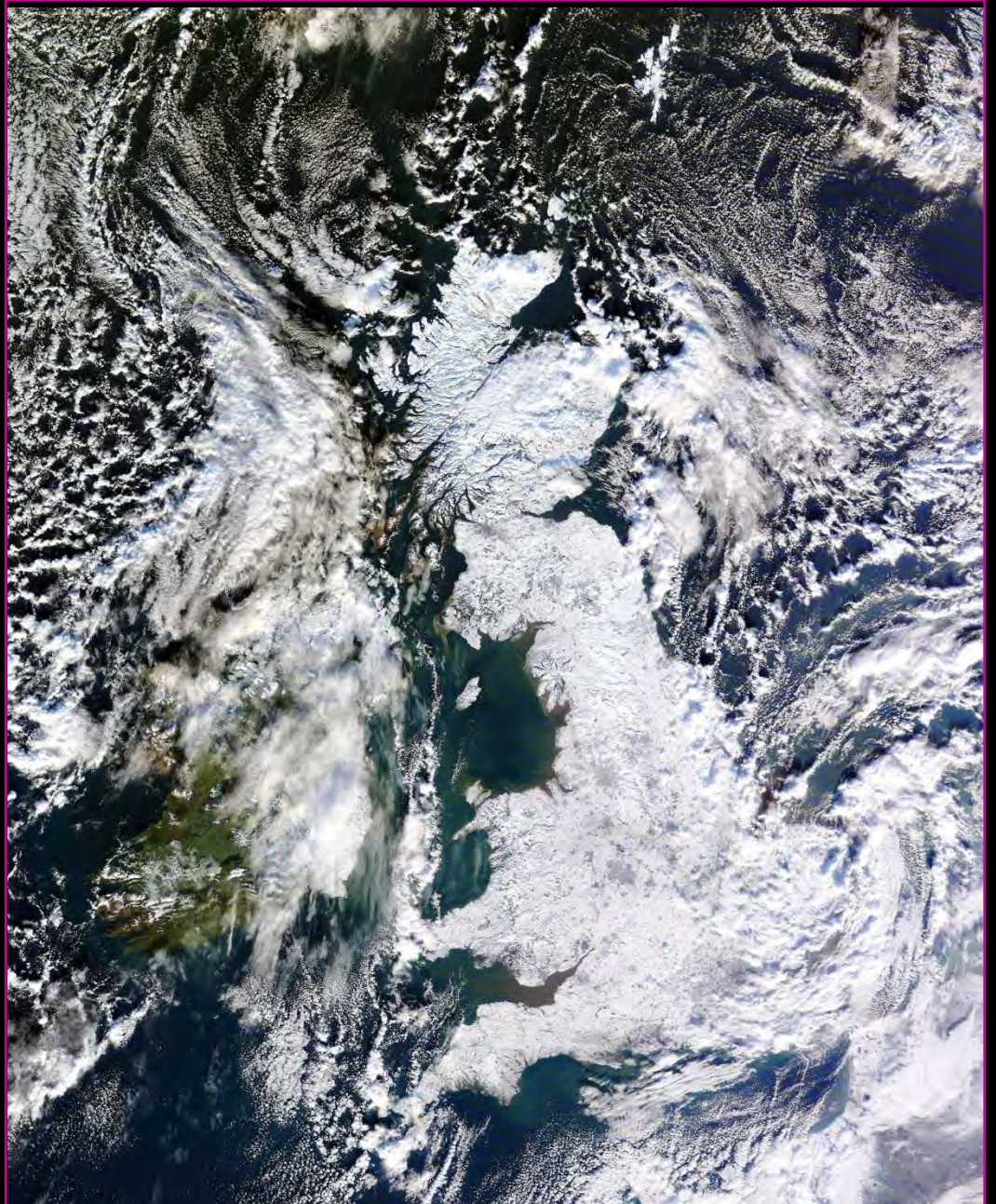


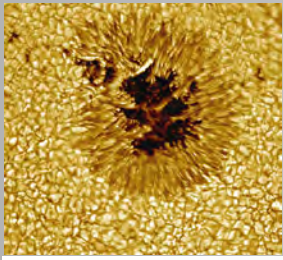
- ▶ Modelled downward and northward propagation of easterly wind anomaly (by Eliassen-Palm flux divergence)
- ▶ seen in ERA40+ data
- ▶ c.f. Kodera and Kuroda, 2002; Matthes et al., 2006

(Ineson et al, Nature Geosci., 2011)

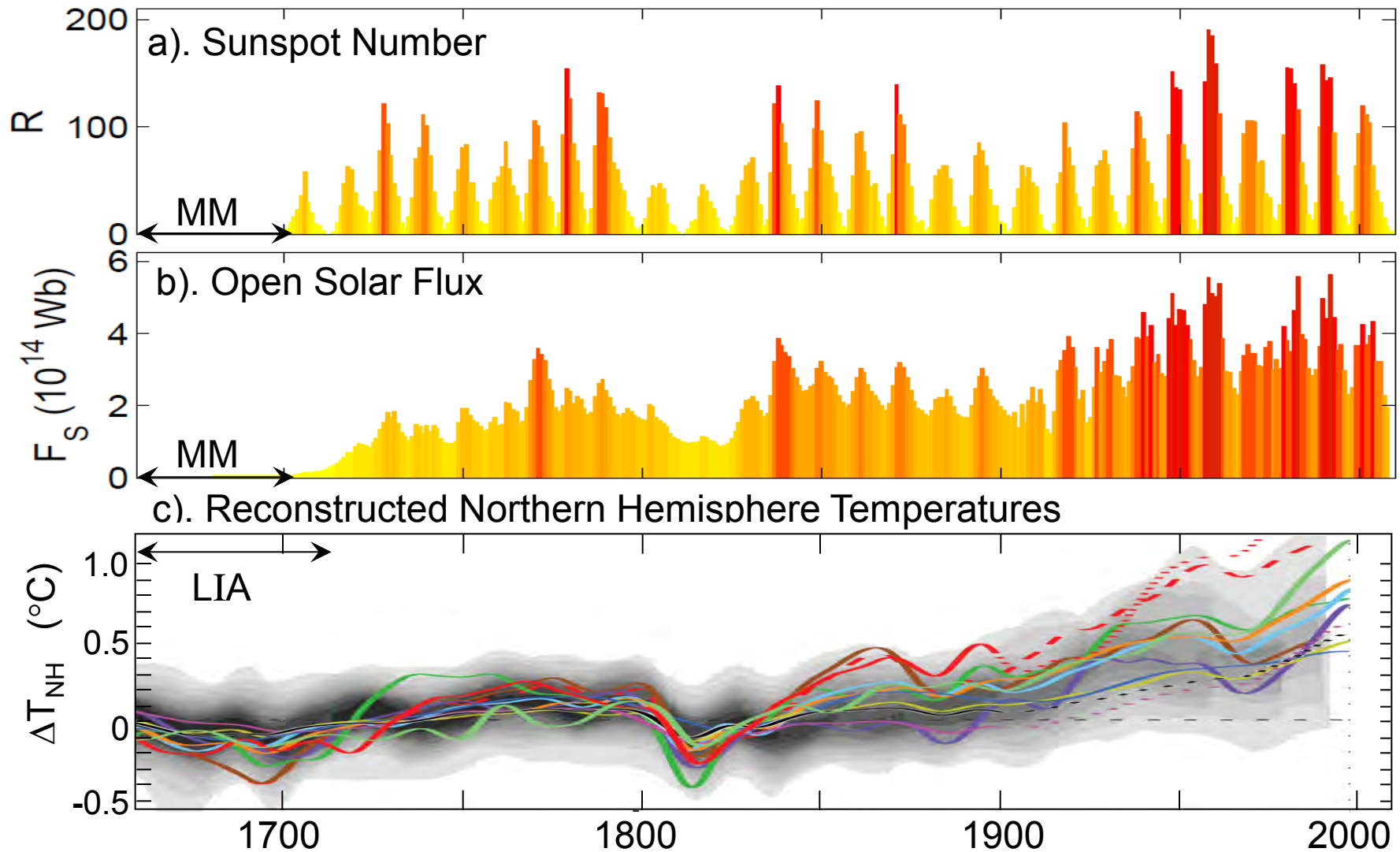


January 2010





Maunder Minimum & the “Little Ice Age”



January 1683

A Frost Fair on the Thames in London. The river froze in central London relatively frequently during the Maunder Minimum of sunspot activity



“An exact and lively mapp ... with an alphabetical explanation of the most remarkable figures”



“An exact and lively
mapp ... with an
alphabetical explanation
of the most remarkable
figures”

H. The Musick Booth



“An exact and lively mapp ... with an alphabetical explanation of the most remarkable figures”

I. The Printing Booth



“An exact and lively mapp ... with an alphabetical explanation of the most remarkable figures”

E. The Roast Beefe Booth



“An exact and lively
mapp ... with an
alphabetical explanation
of the most remarkable
figures”

N. The Boat drawne with
a Hors



“An exact and lively
mapp ... with an
alphabetical explanation
of the most remarkable
figures”

Q. The Bull Baiting



“An exact and lively mapp ... with an alphabetical explanation of the most remarkable figures”

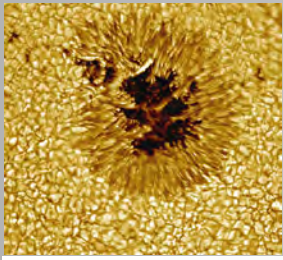
C. The Tory Booth



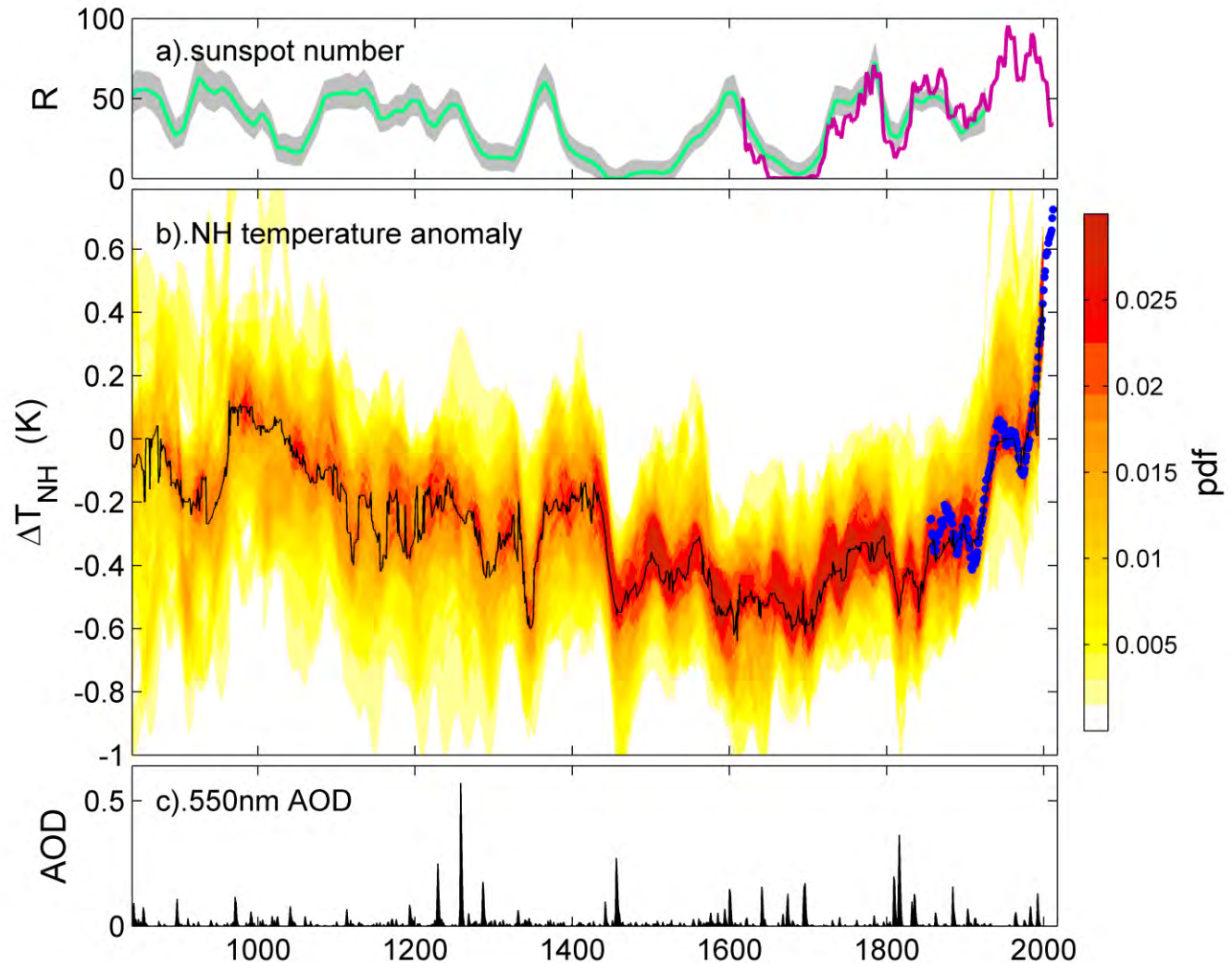
“An exact and lively mapp ... with an alphabetical explanation of the most remarkable figures”

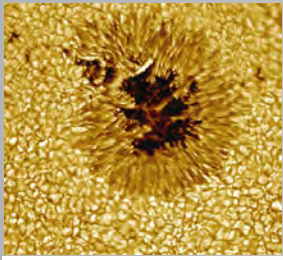
Z. London Bridge



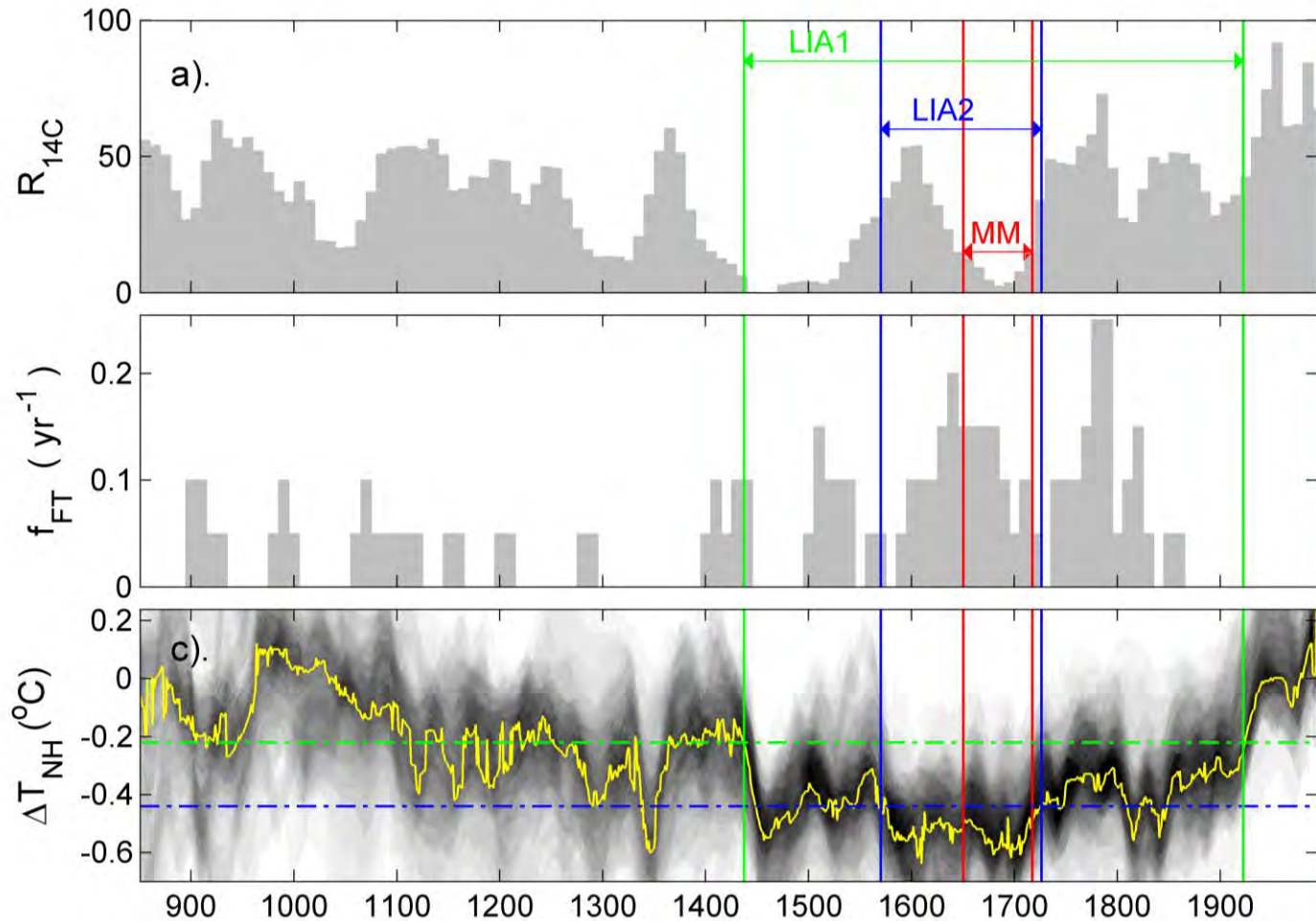


Maunder Minimum & the “Little Ice Age”

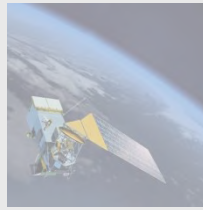
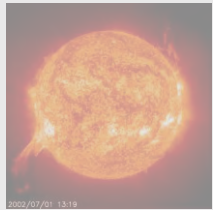




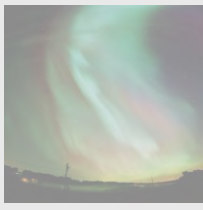
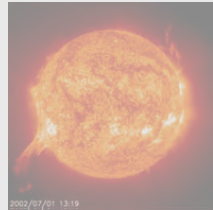
Maunder Minimum & the “Little Ice Age”



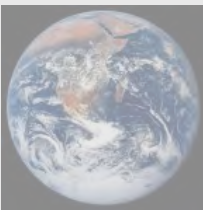
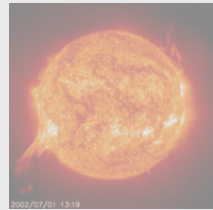
Solar Variability: Effects on Climate?



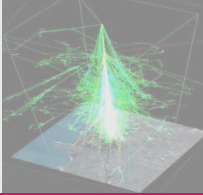
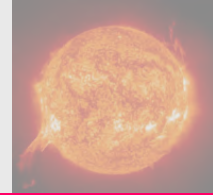
Solar Outputs



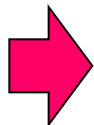
Solar Variability



Global Effects



Regional & Seasonal Effects



The Future



Predictions for the future



“It is not important to predict the future, but it is important to be prepared for it”



Pericles,

Athenian orator, statesman and general
c. 495 – 429 BC

“It is not important to know the future, but to shape it”



Antoine de Saint Exupéry,

French writer and aviator
1900 - 1944

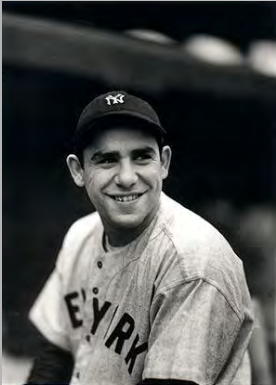


“Prediction is very hard — especially when it’s about the future”

Niels Bohr

Danish Physicist

1885 – 1962



“Never make predictions — especially about the future”

Lawrence Peter ("Yogi") Berra

American Baseball Player, coach and author

1925 –

Who also said

“I never said half the things I really said.”

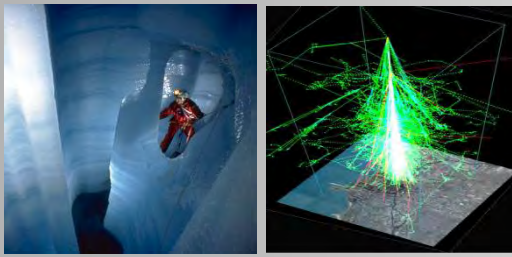
"It ain't over 'till it's over"

"When you come to a fork in the road, take it."

"It's like déjà vu all over again"

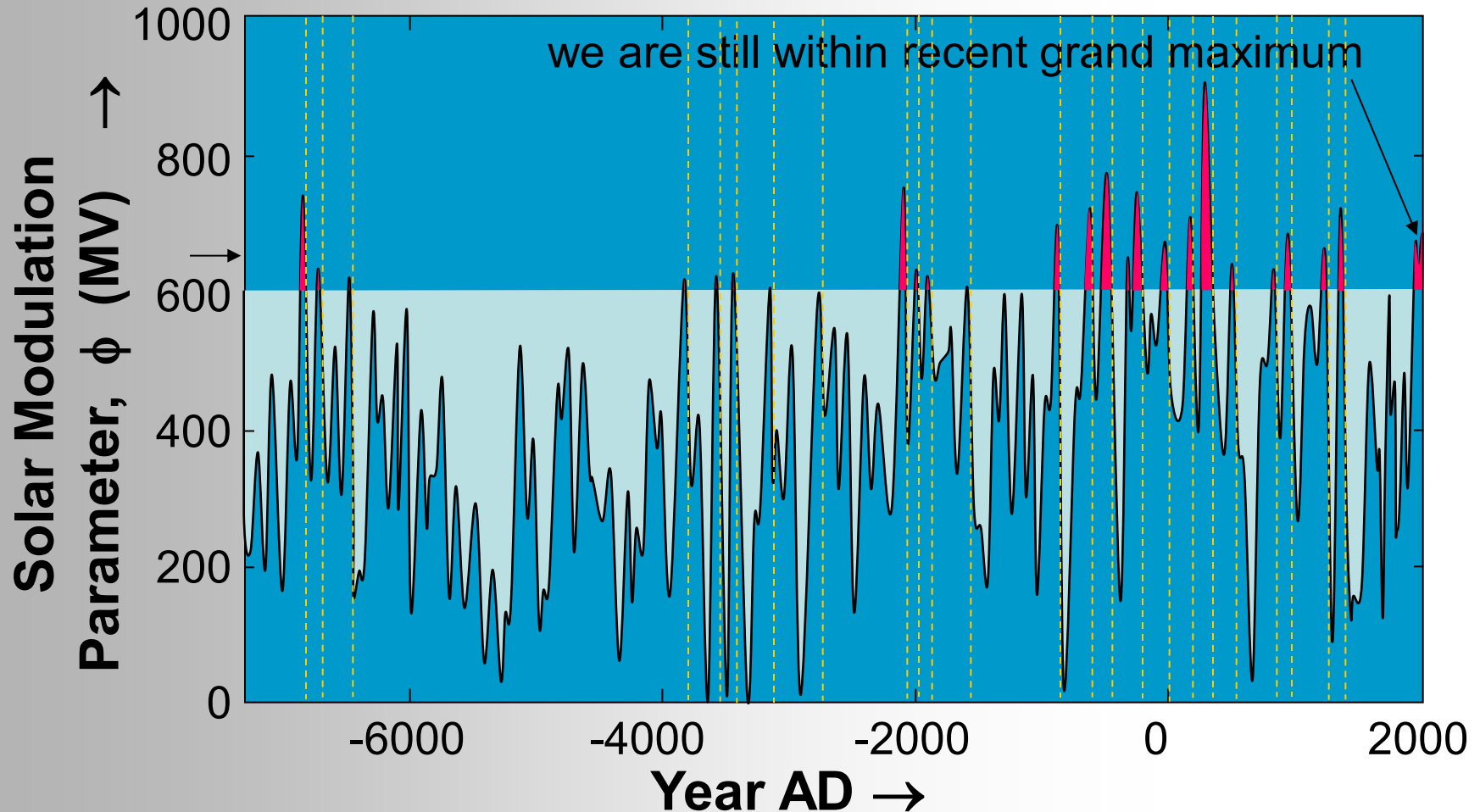
"Always go to other peoples' funerals, otherwise they won't come to yours."

“I don't have nightmares about my team – you've got to be able to sleep before you can have nightmares”

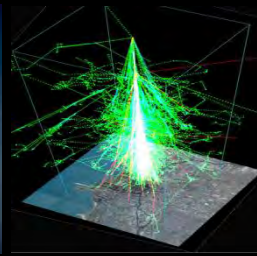


Millennial Variation

ϕ composite (25-year means) from cosmogenic isotopes by Steinhilber et al. (2008)



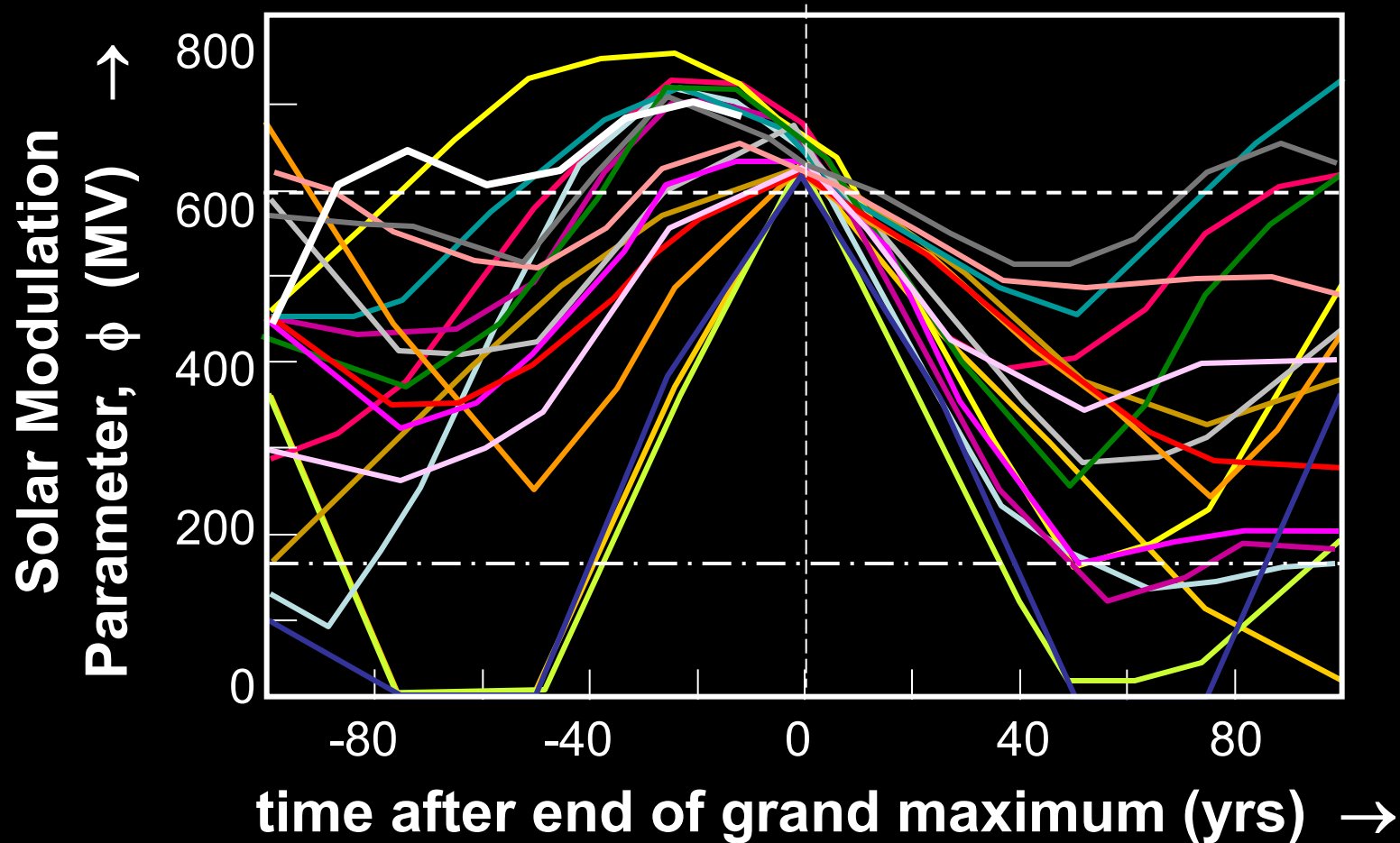
composite from Solanki et al., 2004; Vonmoos et al., 2006 & Muscheler et al., 2007

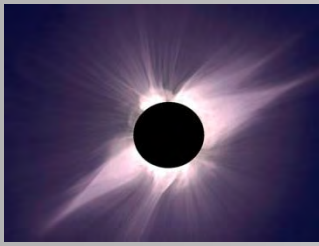


Superposed epoch study of the end of grand maxima

(24 events in 9000 yrs)

↓ end of grand solar maximum



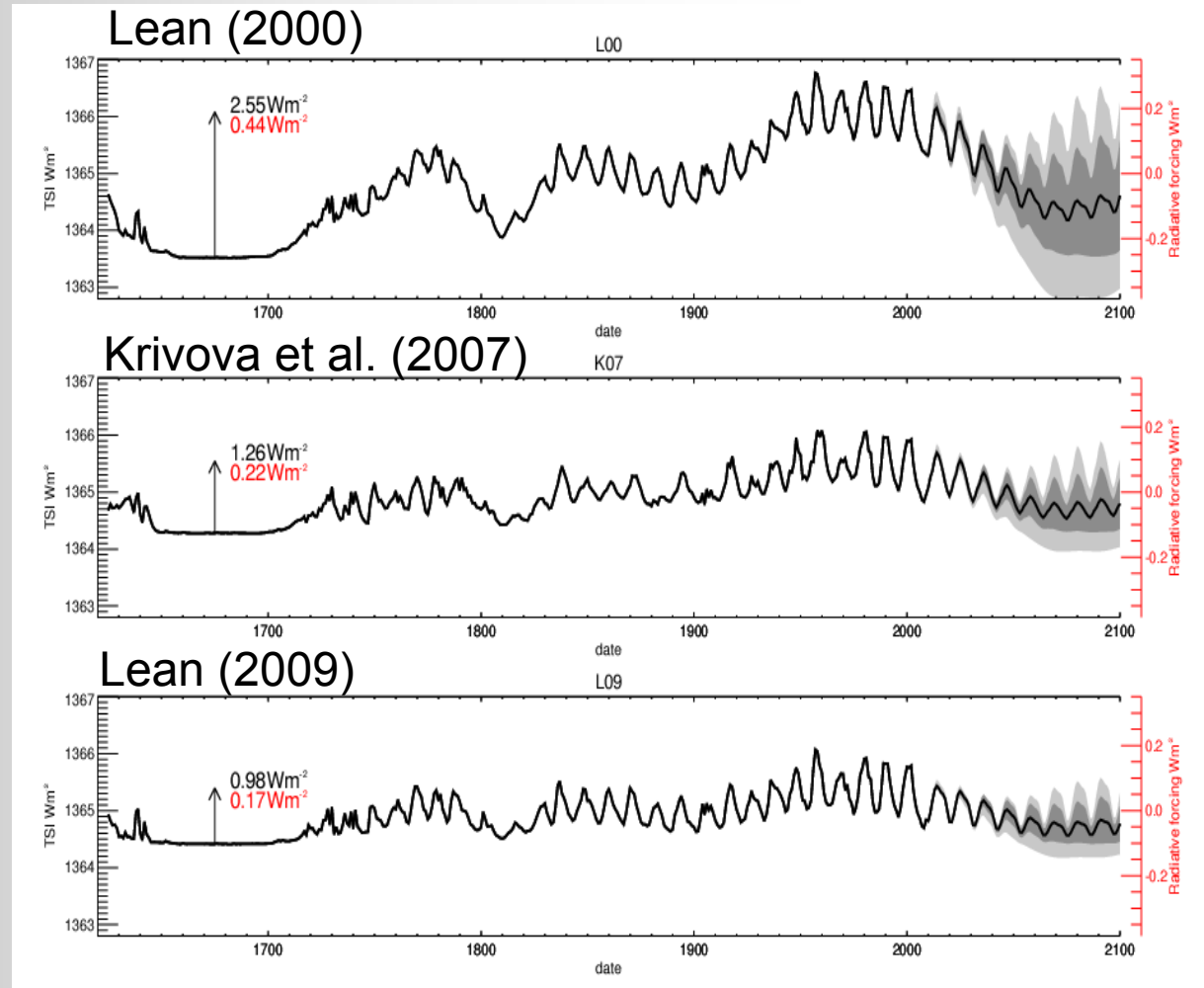


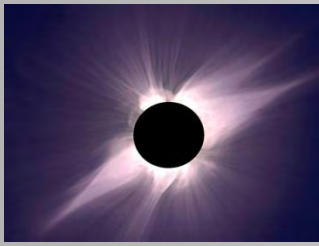
Future TSI Variation?

(Jones, Lockwood and Stott, JGR 2011)



- using the relationship of TSI and GCRs
- & relationship between solar cycle amplitude and the mean



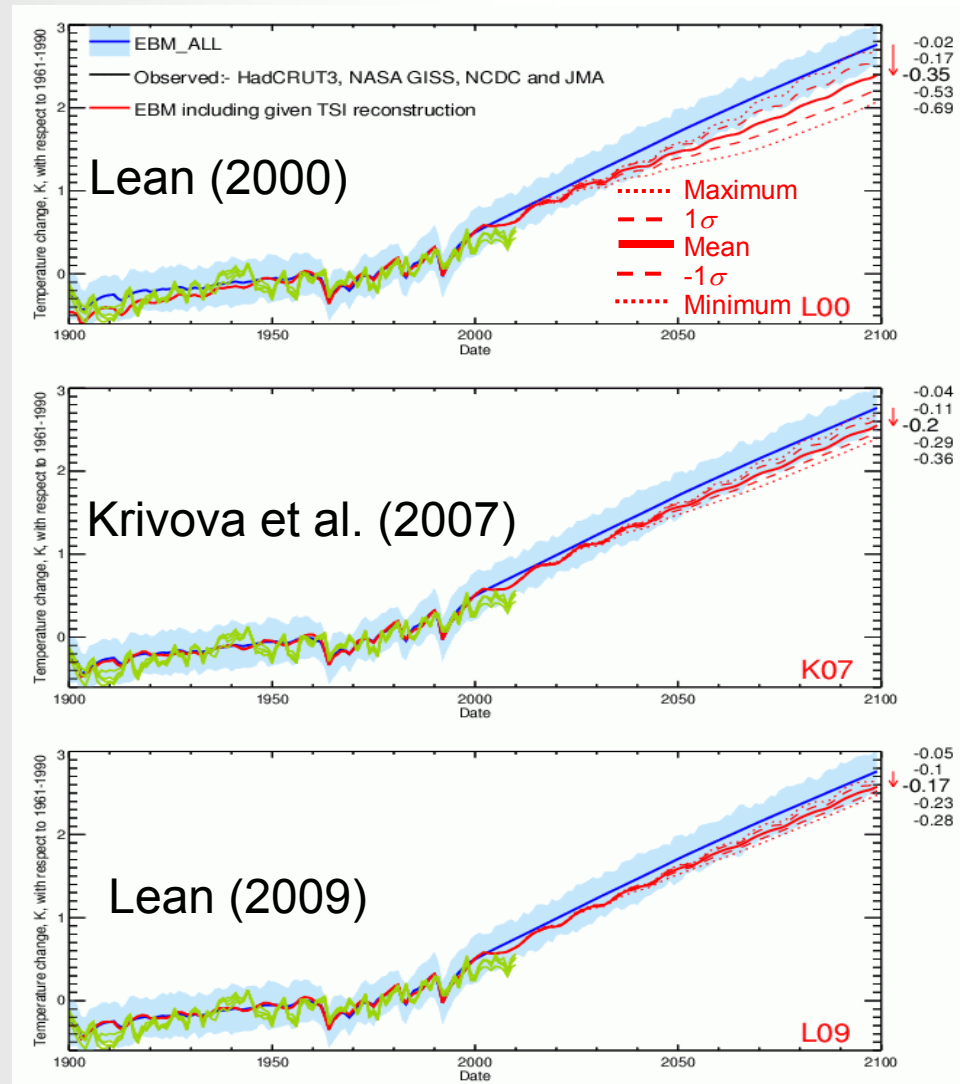


GMAST Predictions – EBM tuned to HadCM3

(Jones, Lockwood and Stott, *JGR in press*, 2011)



- use B2 SRES emissions scenario
- no future volcanic forcing
- solar responses have been scaled to match a maximum possible solar cycle amplitude of 0.1K.



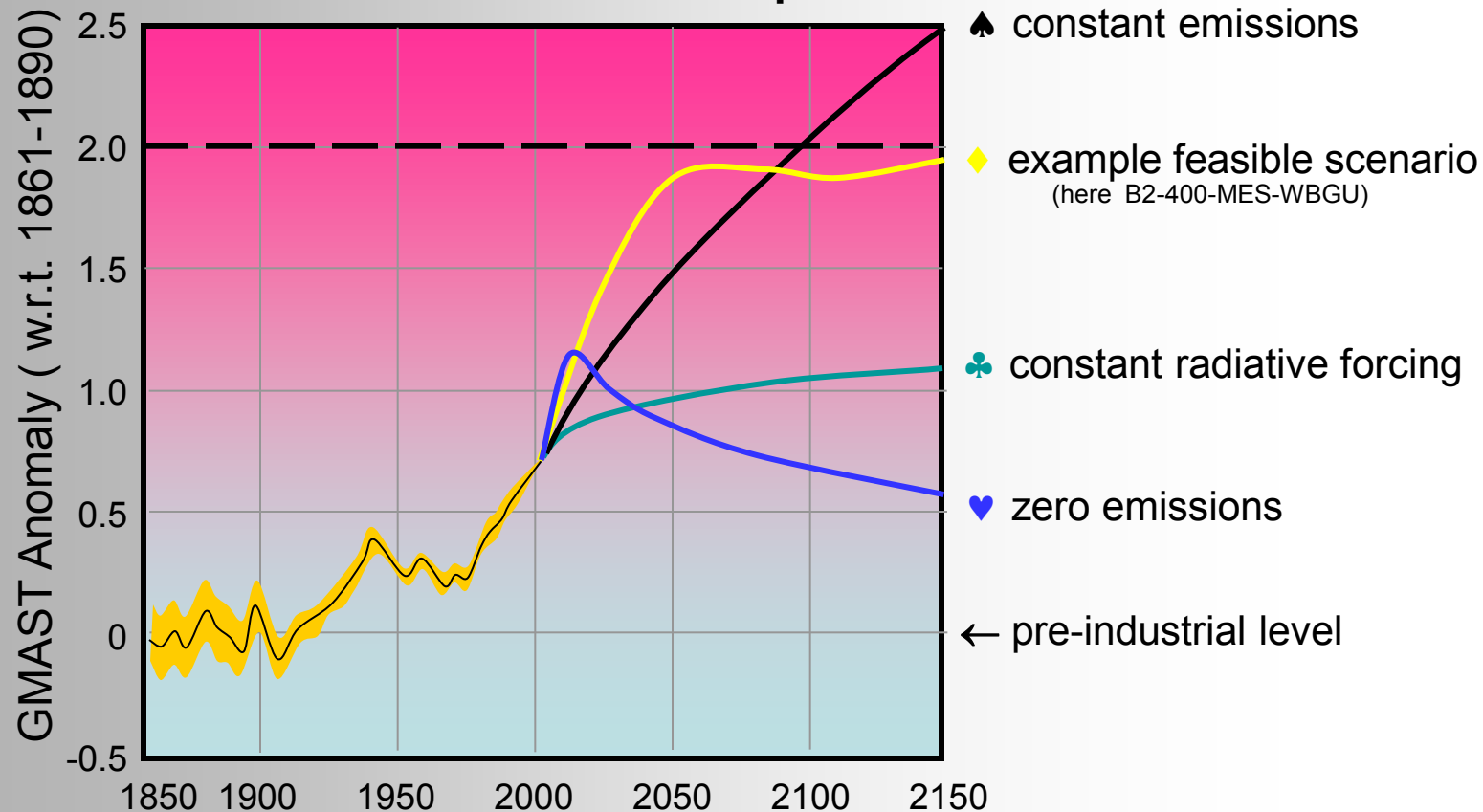


Temperature Commitment

Climate Sensitivity 2.8°C



Global Mean Air Surface Temperature



(Hare & Meinshausen, 2006)



Handling Uncertainty

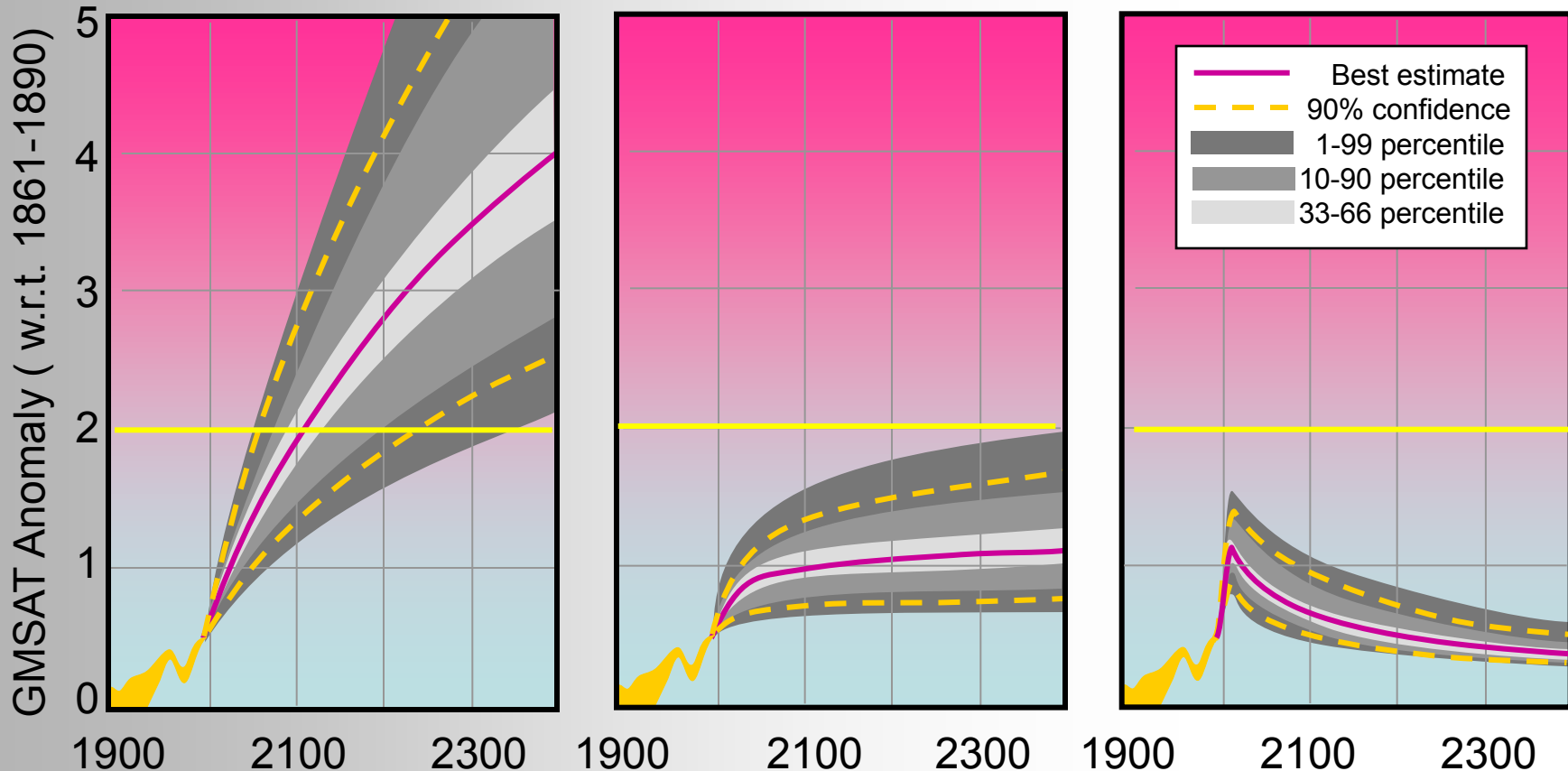
- For IPCC lognormal pdf of climate sensitivity

Global Mean Air Surface Temperatures for

Constant emissions

Present forcing

Zero Emissions



(Hare & Meinshausen, 2006)



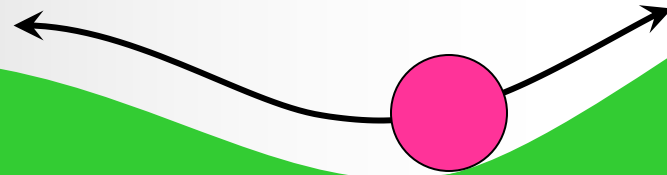
Tipping Points

(Lenton et al., 2007)



System State 2

System State 1



Not necessarily irreversible



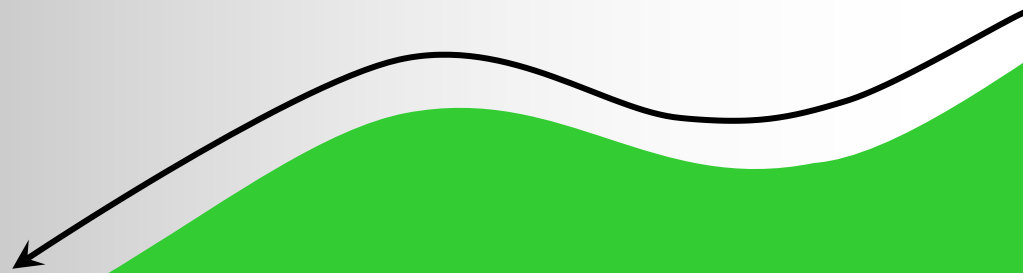
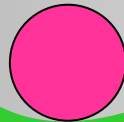
Tipping Points

(Lenton et al., 2007)



System State 2

System State 1





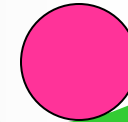
Tipping Points

(Lenton et al., 2007)



System State 2

System State 1



Potential tipping points between climate states are:

- ▶ Atlantic thermohaline circulation disruption
- ▶ Indian monsoon chaotic multistability
- ▶ West African monsoon latitude shift
- ▶ Change in ENSO frequency and/or amplitude
- ▶ West Antarctic ice sheet instability
- ▶ Changes in Antarctic bottom water formation
- ▶ Arctic sea ice loss
- ▶ Greenland ice sheet melting
- ▶ Boreal forest dieback
- ▶ Loss of permafrost and tundra
- ▶ Sahara greening
- ▶ Amazon rainforest dieback

time to...

STOP!!!

but questions most welcome,
now,
over dinner,

or down the pub after

